

Self-consistent gaps and equation of states for neutron star applications

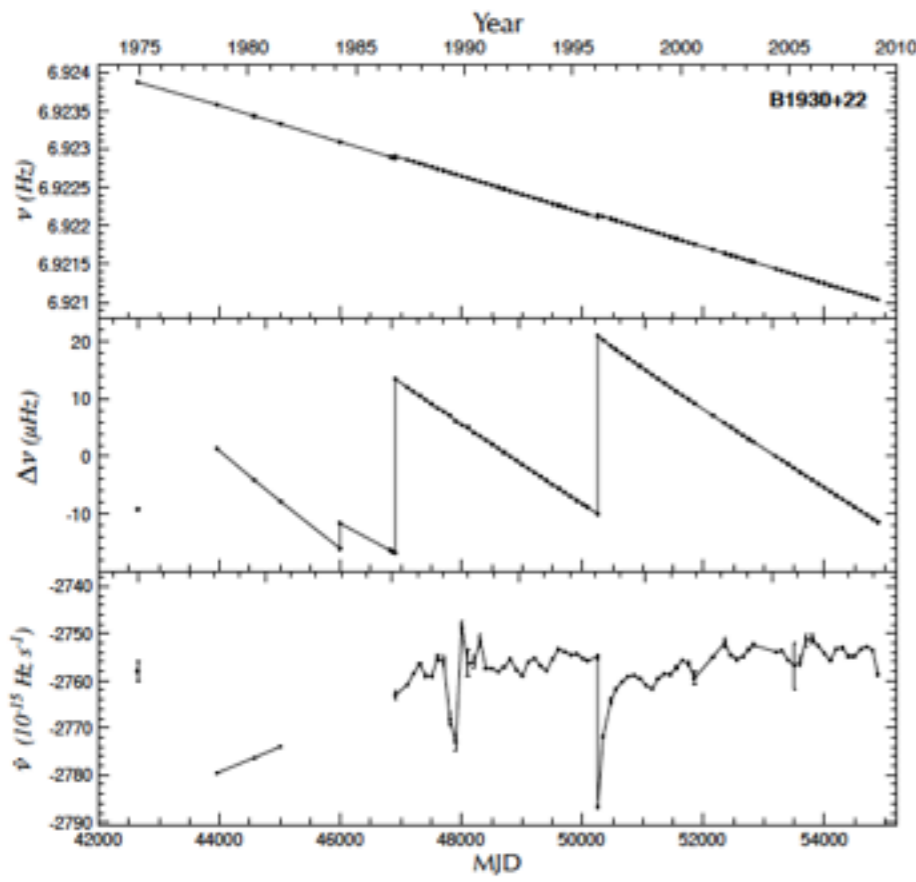
Ding, Rios, *et al.*, *Phys Rev C* **94** 025802 (2016)
Sellaheewa & Rios, *in preparation*

Arnau Rios Huguet
Lecturer & STFC Advanced Fellow
Department of Physics
University of Surrey

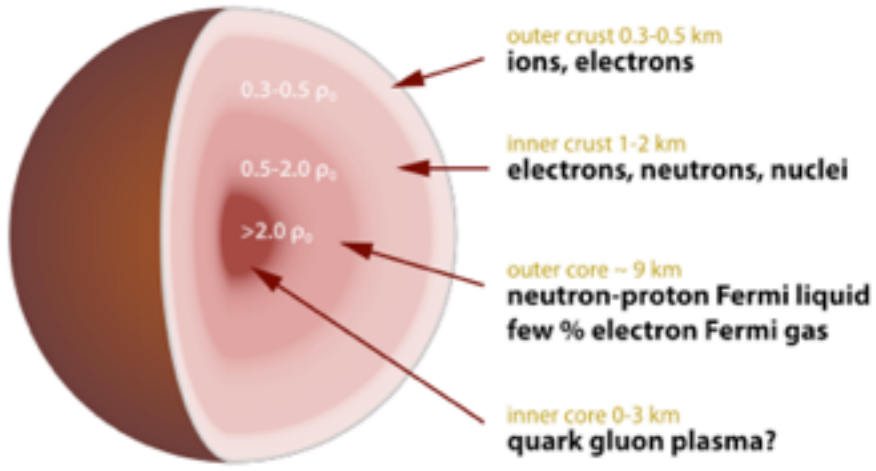
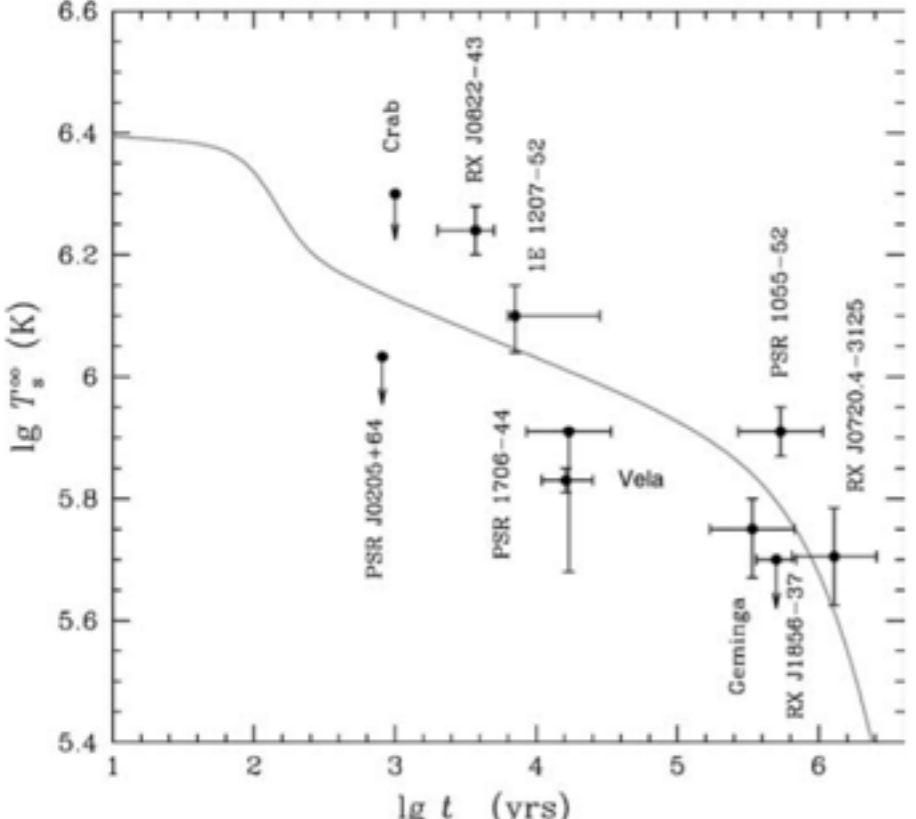
1. NS model with effective forces
2. NS model with realistic forces
3. Correlations and systematic errors

NS data!

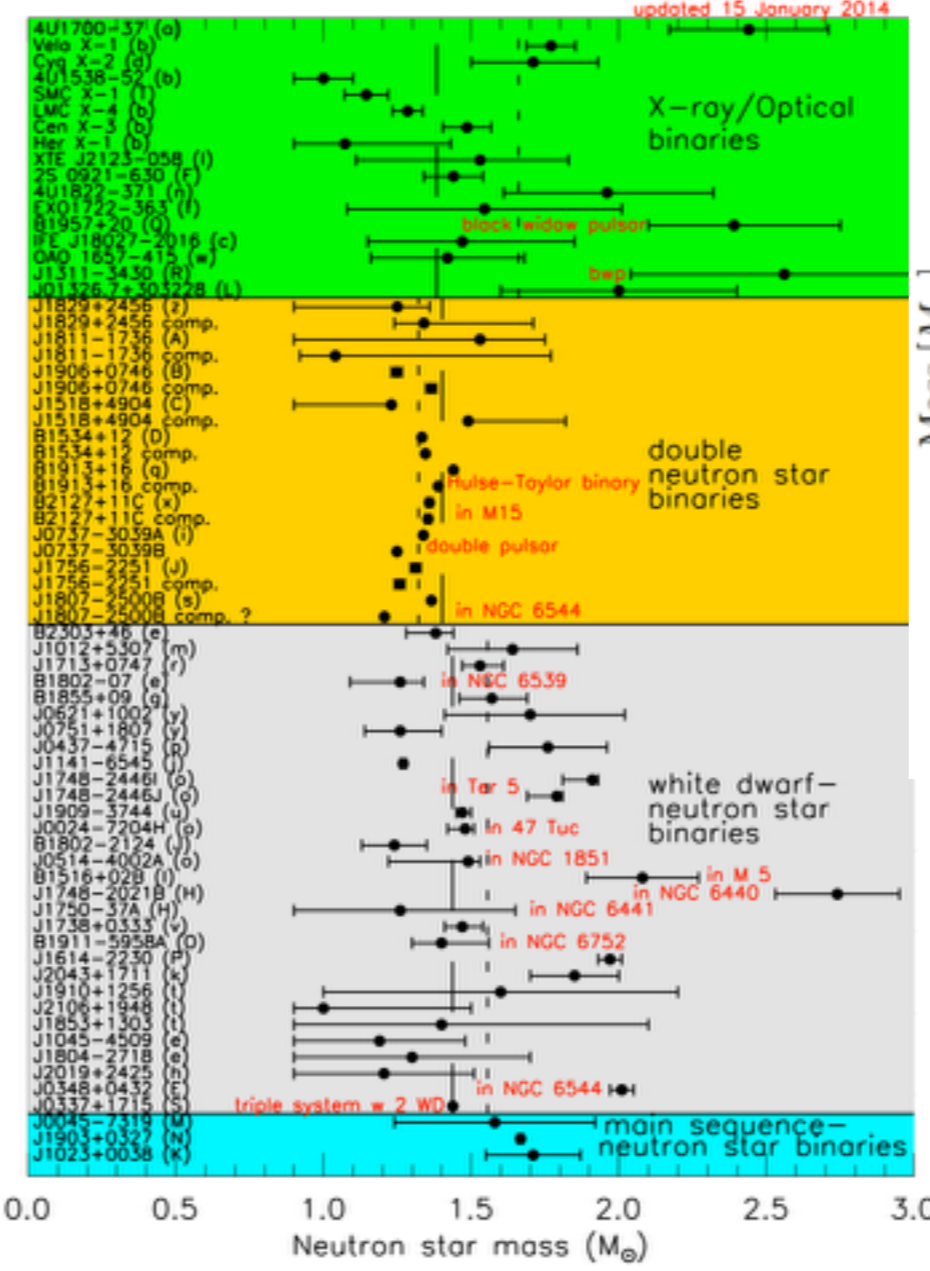
Glitches



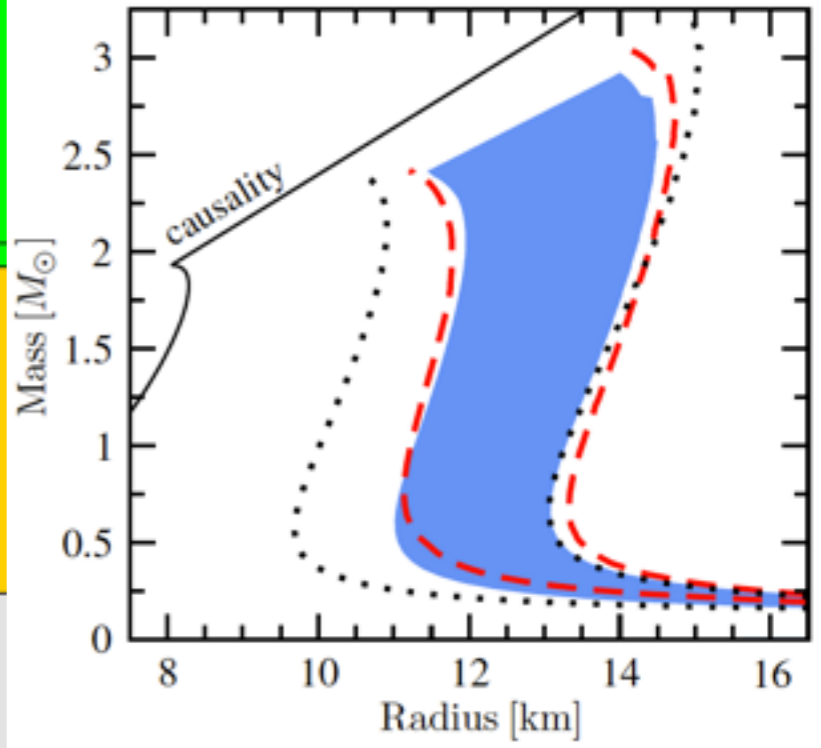
ν cooling



Masses

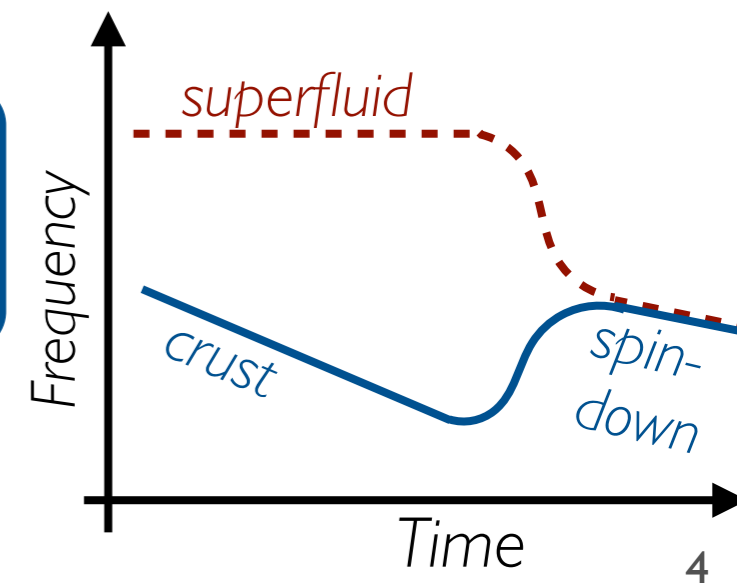
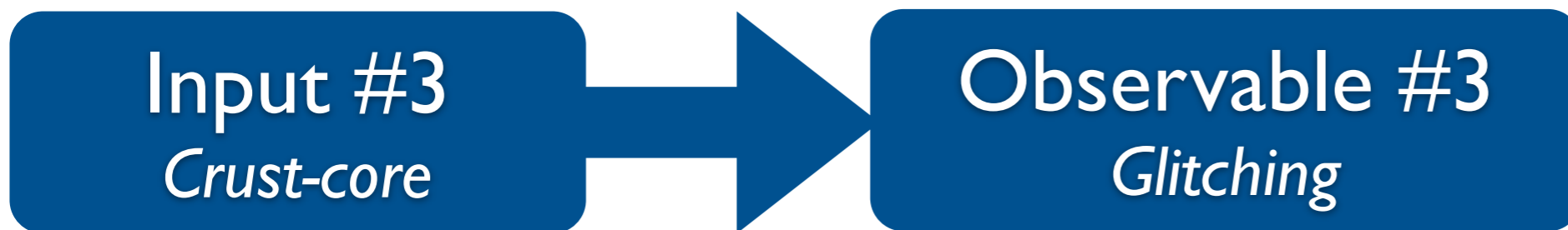
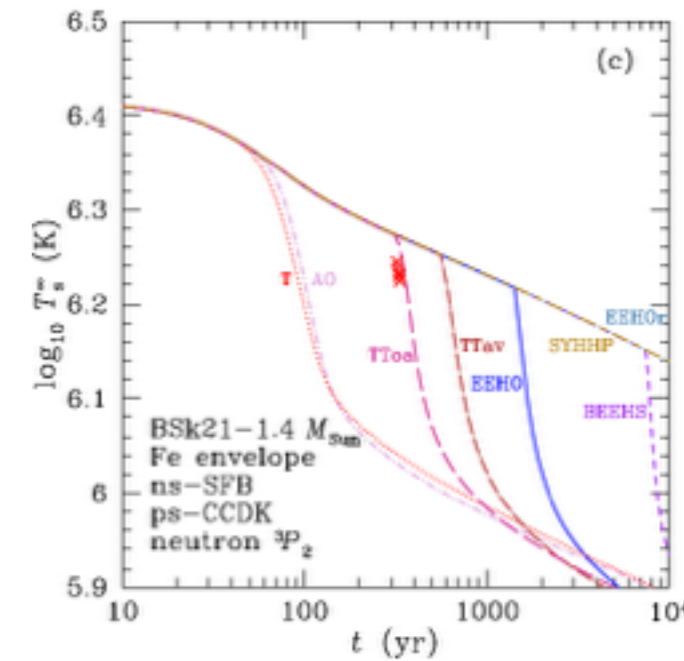
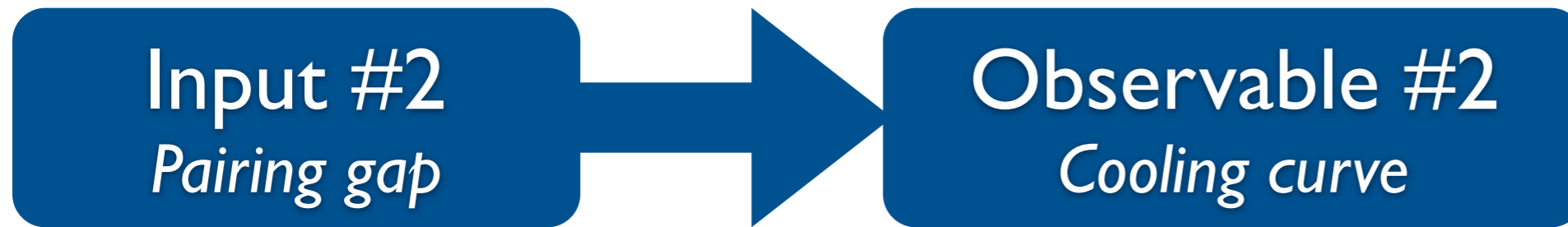
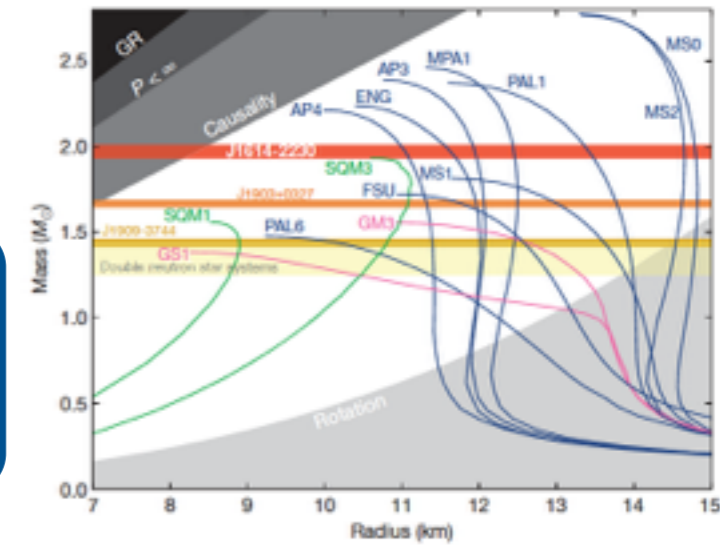
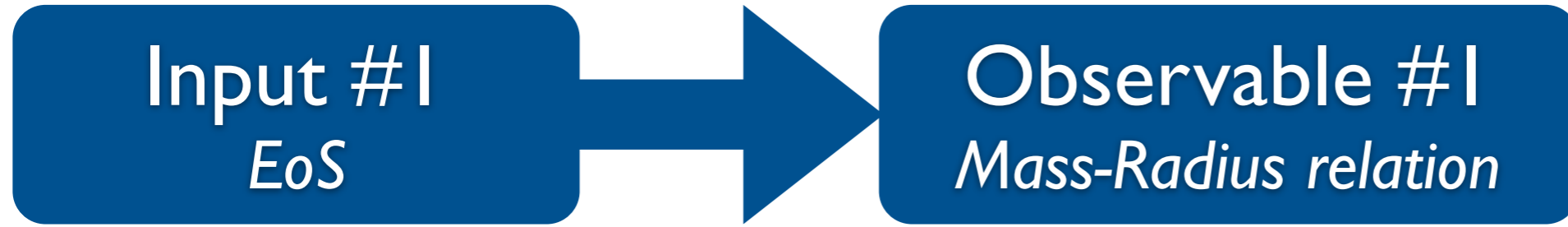


Radii



Hebeler, Lattimer, Pethick, Schwenk
ApJ **773** 11 (2013)

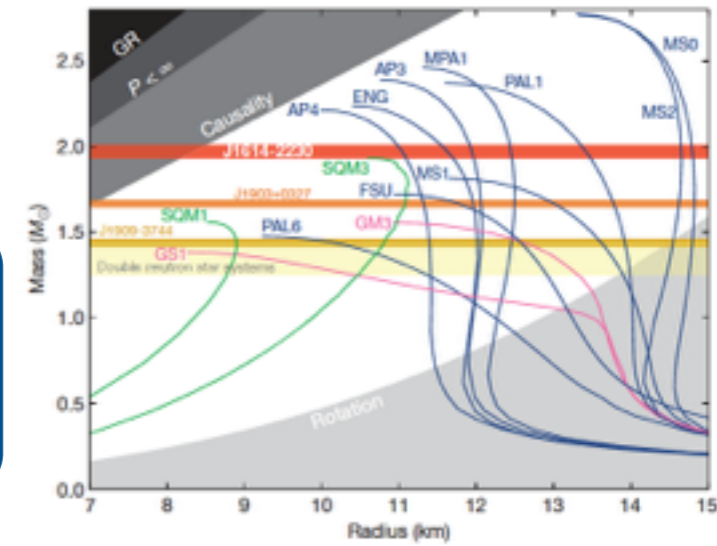
Neutron star modeling



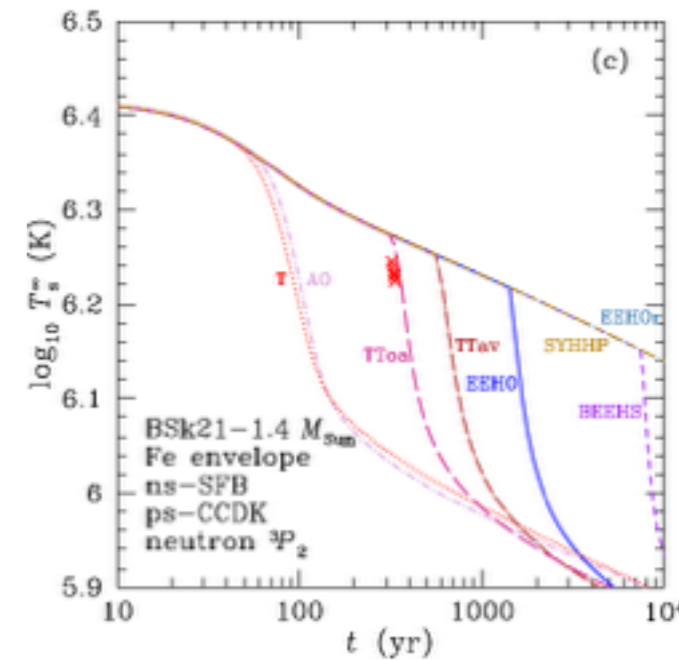
Neutron star modeling

Input
Consistent many-body theory

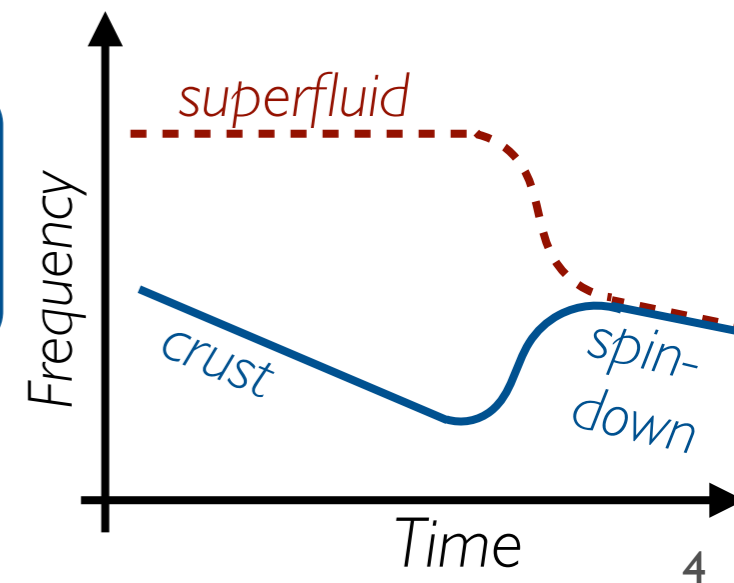
Observable #1
Mass-Radius relation



Observable #2
Cooling curve



Observable #3
Glitching



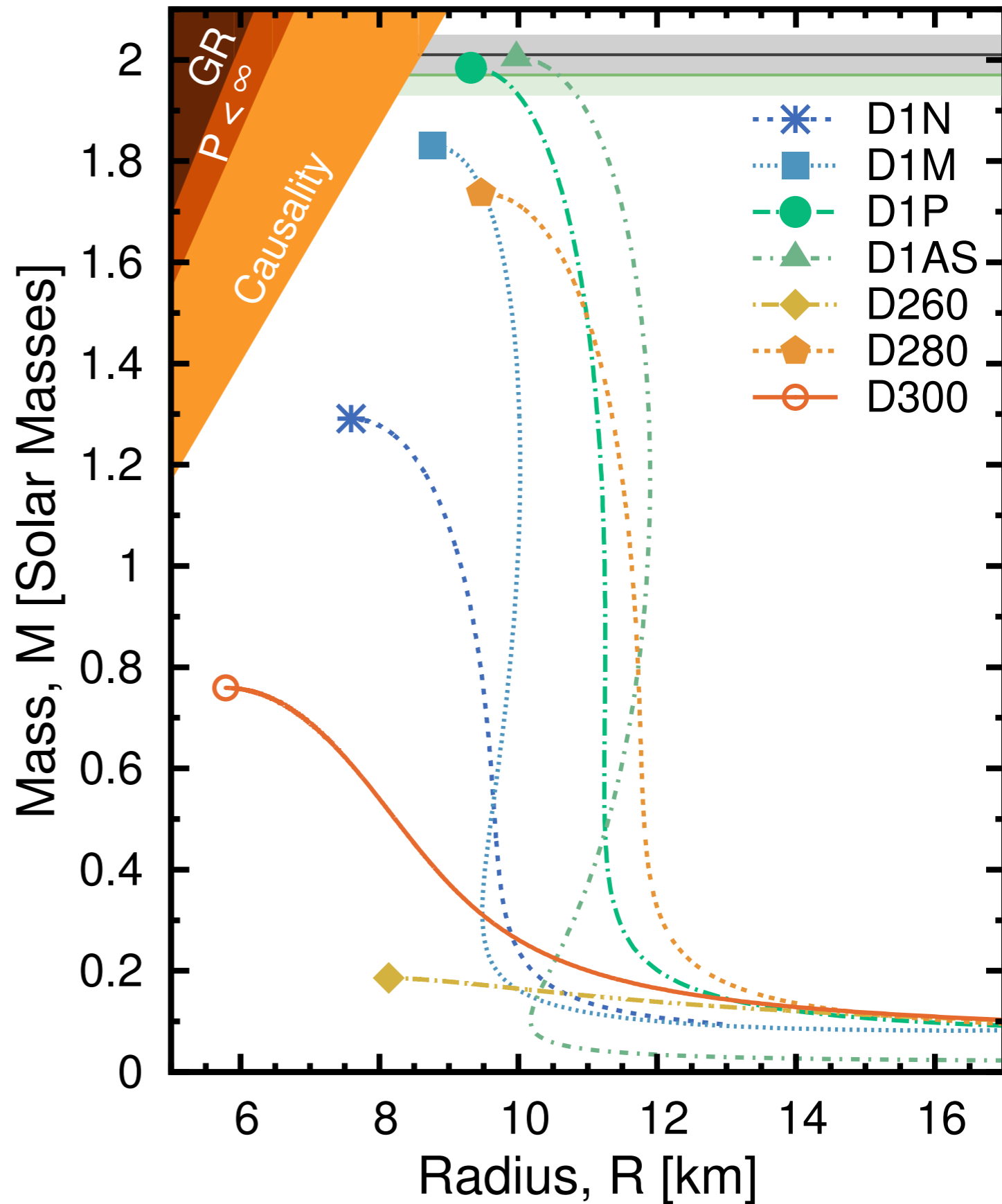
A step forward towards consistency

The Gogny force



$$V(\vec{r}) = \sum_{i=1,2} \left(W_i + B_i P_\sigma - H_i P_\tau - M_i P_\sigma P_\tau \right) e^{-\frac{r^2}{\mu_i}} + \sum_{i=1,2} t_0^i (1 + x_0^i P_\sigma) \rho^{\alpha_i} \delta(\vec{r}) + iW_0(\sigma_1 + \sigma_2) [\vec{k}' \times \delta(\vec{r}) \vec{k}]$$

- 11 parametrizations
- Popular in **fission** studies
- Any good for isospin?
- Only 7 give stable NSs!
- **Pairing?**

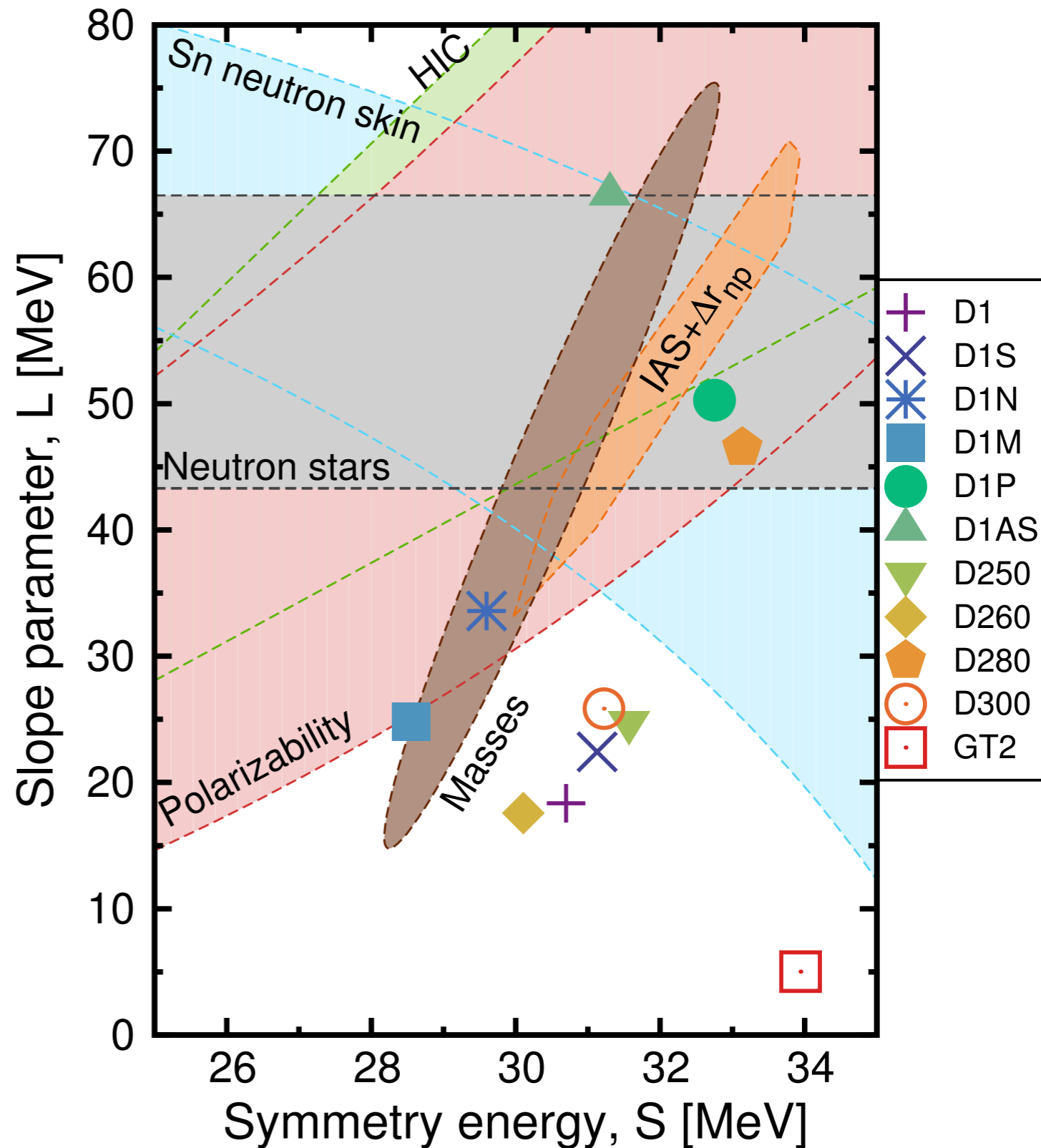


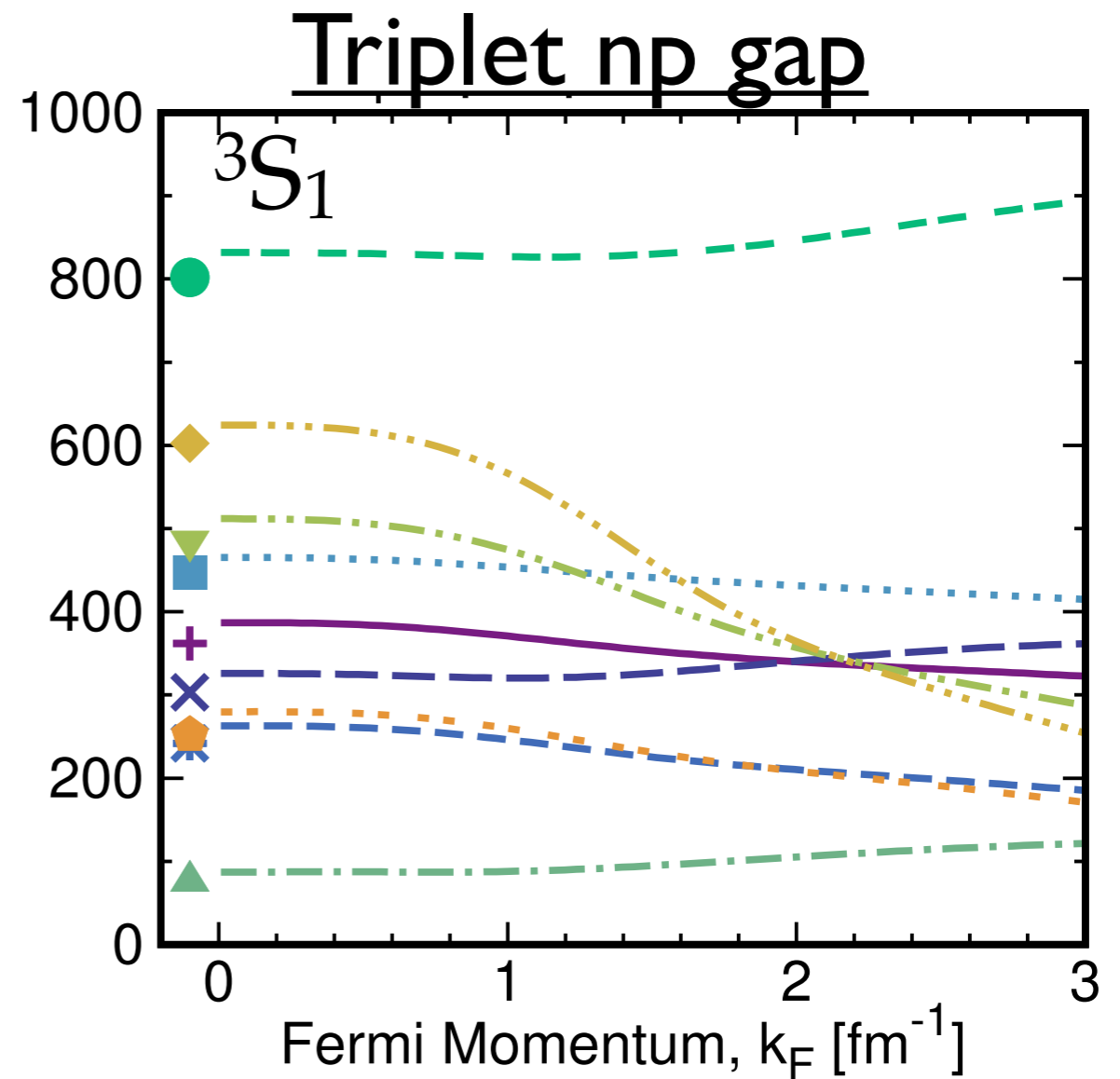
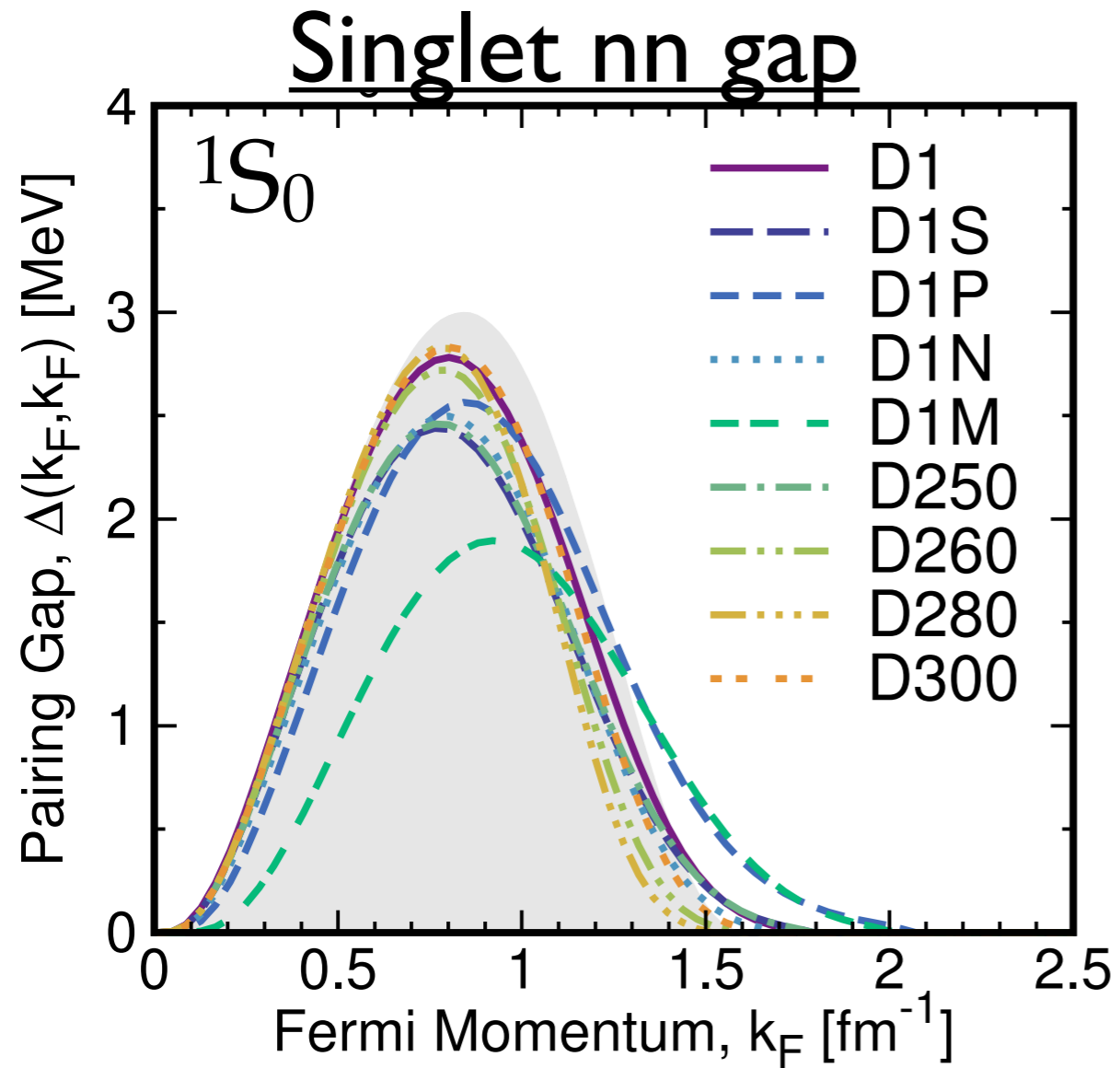
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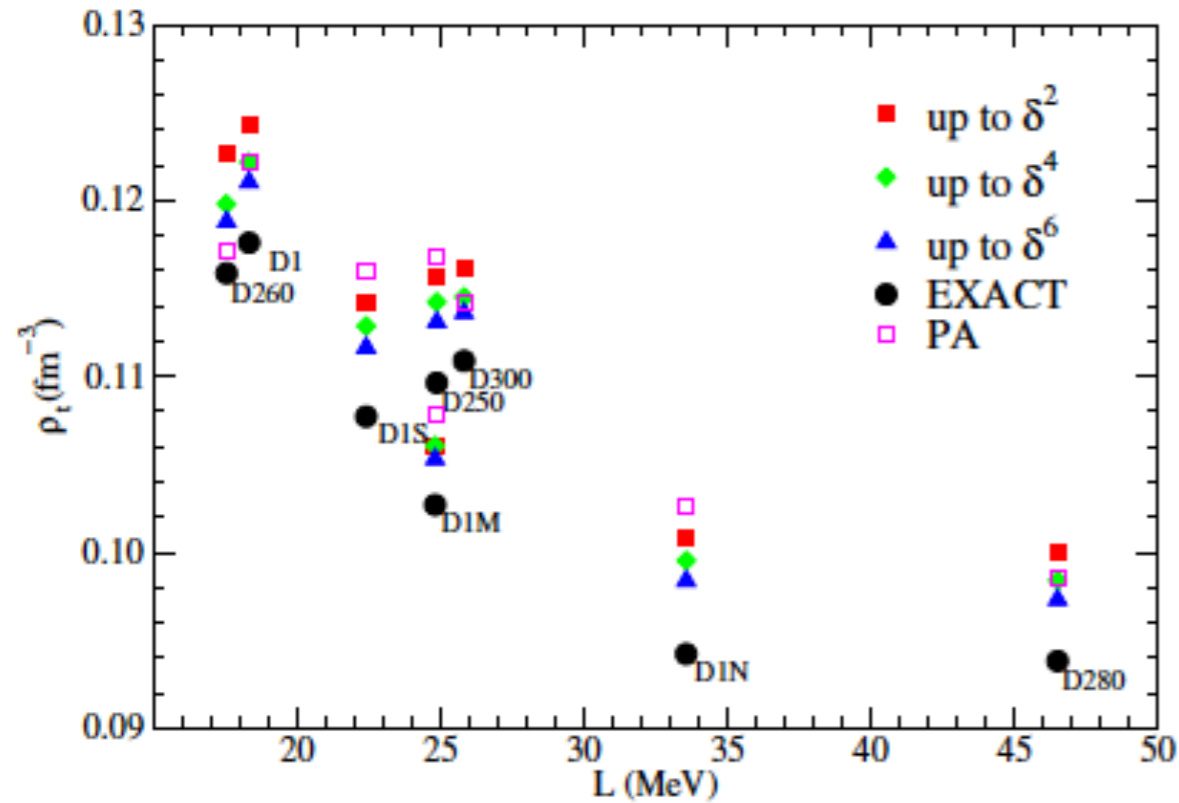
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BCS equation

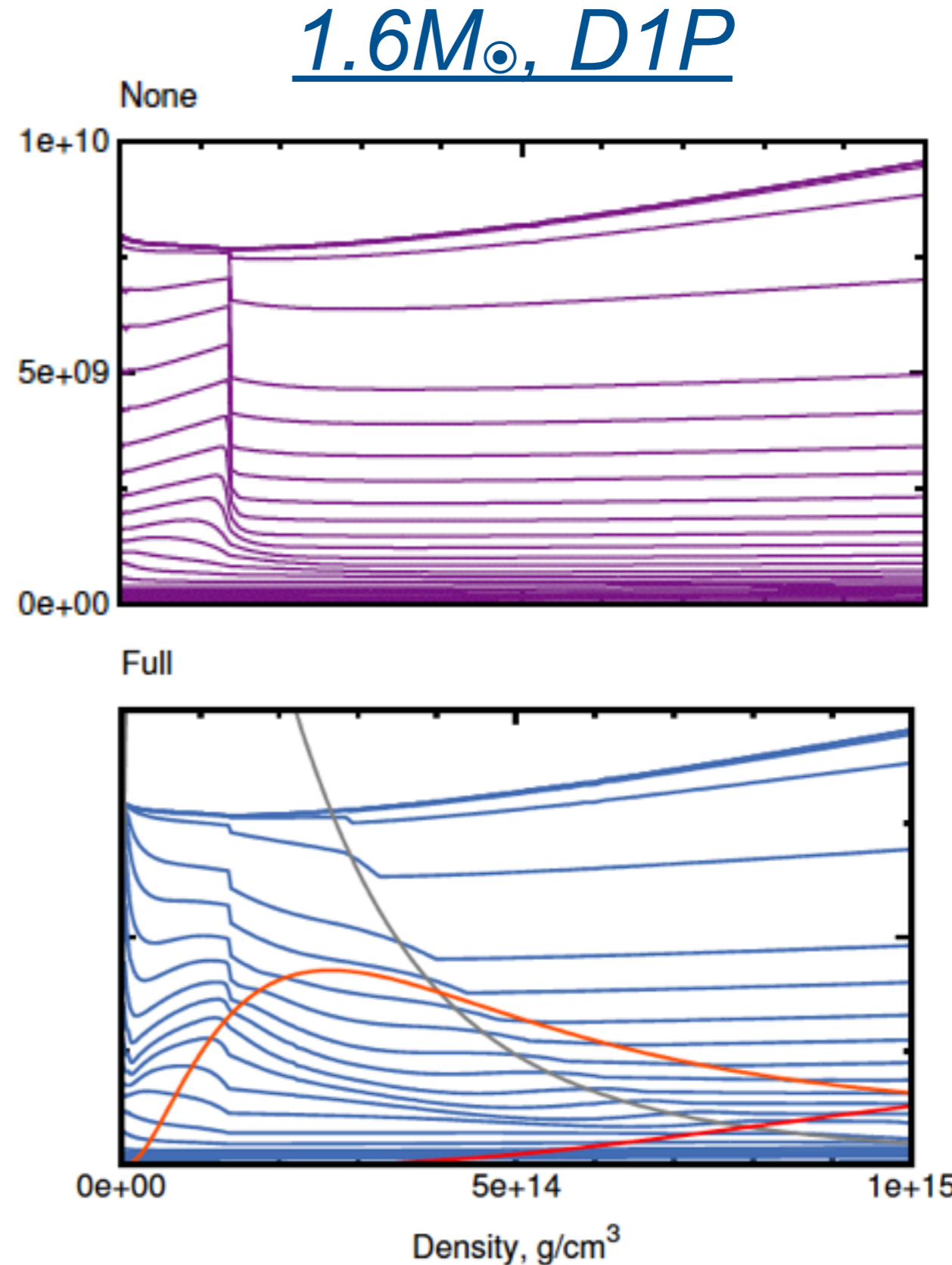
$$\Delta_k^L = - \sum_{L'} \int_{k'} \frac{\langle k | V_{nn}^{LL'} | k' \rangle}{2\sqrt{\chi_{k'}^2 + |\Delta_{k'}|}} \Delta_{k'}^{L'} + \quad \chi_k = \varepsilon_k - \mu$$



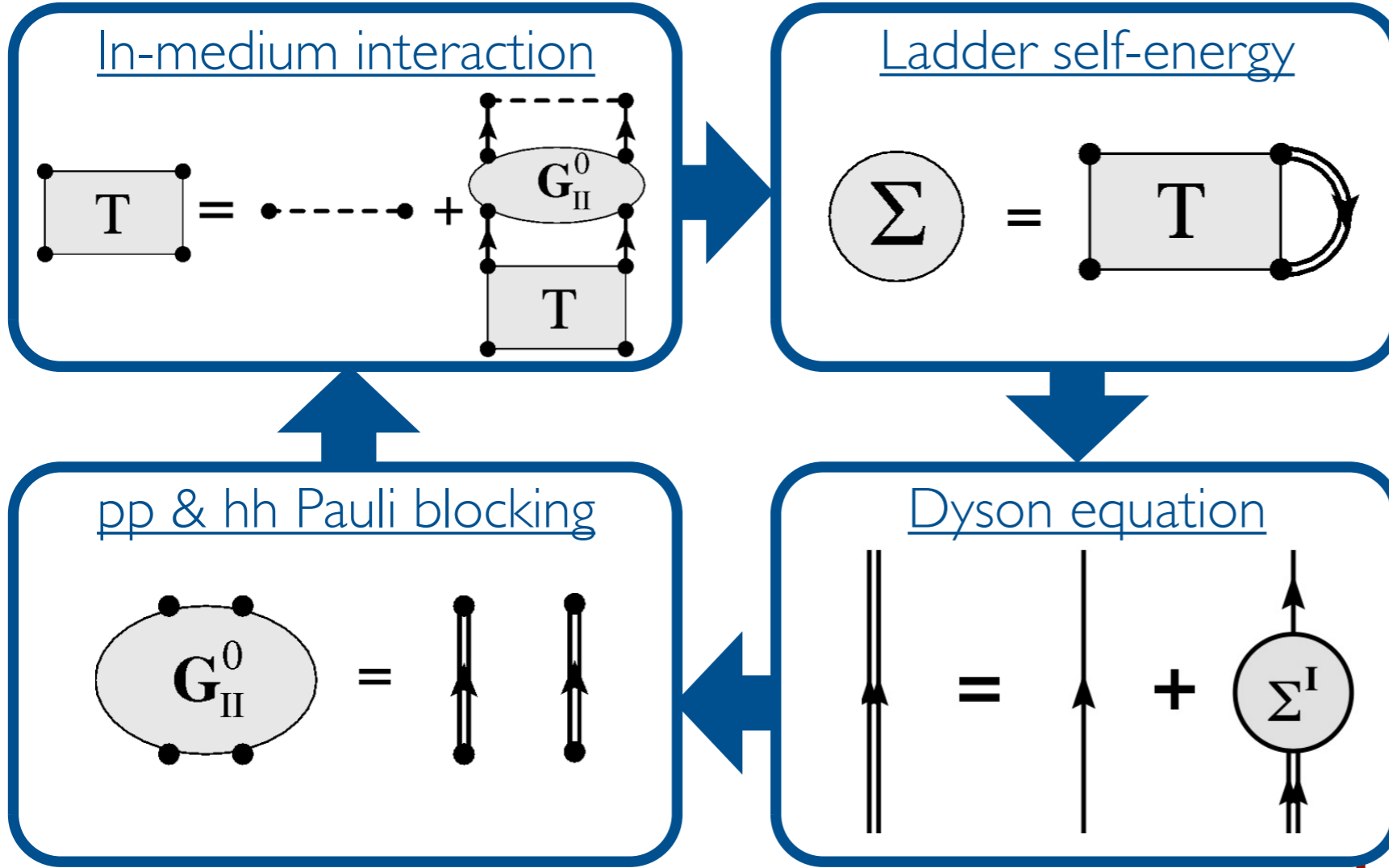
Gonzalez-Boquera, Centelles, Vinas & A. Rios, *in preparation*

Ingredients (D1P)

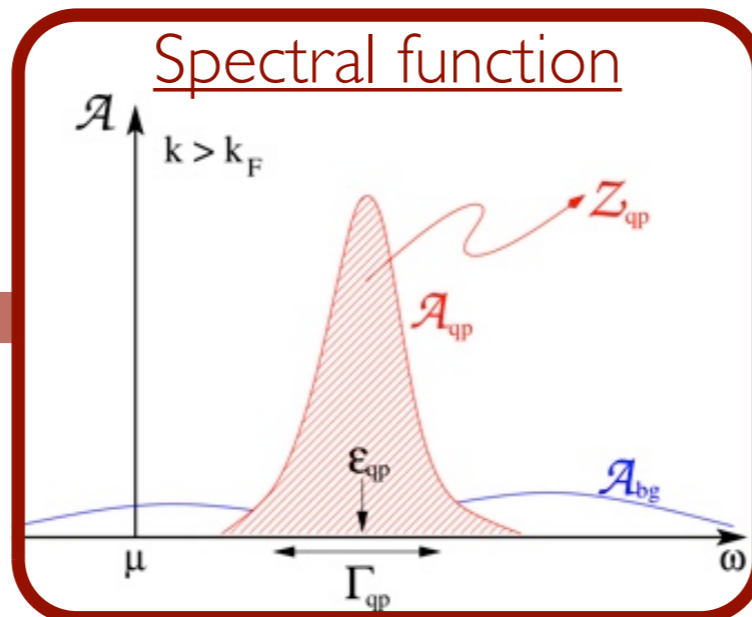
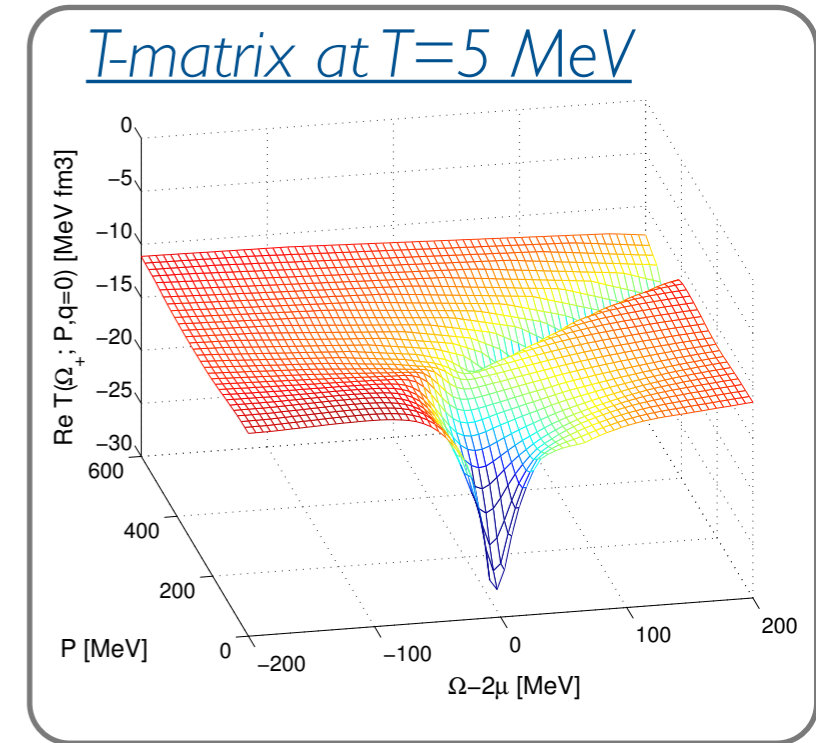
- (a) Mass of pulsar, $1.6M_{\odot}$ ✓
- (b) EoS (except crust) ✓
- (c) Internal composition ✓
- (d) **Pairing gaps** (n 1S_0 & 3P_2) ✓
- (e) Atmosphere composition



SCGF Ladder approximation



- Self-consistent **resummation**
- **Energy** and **momentum** integral
- @**Finite T** (Matsubara)



One-body properties
Momentum distribution
Thermodynamics & EoS
Transport

Ramos, Polls & Dickhoff, NPA **503** 1 (1989)
 Alm *et al.*, PRC **53** 2181 (1996)
 Dewulf *et al.*, PRL **90** 152501 (2003)
 Frick & Muther, PRC **68** 034310 (2003)
 Rios, PhD Thesis, U. Barcelona (2007)
 Soma & Bozek, PRC **78** 054003 (2008)
 Rios & Soma PRL **108** 012501 (2012)

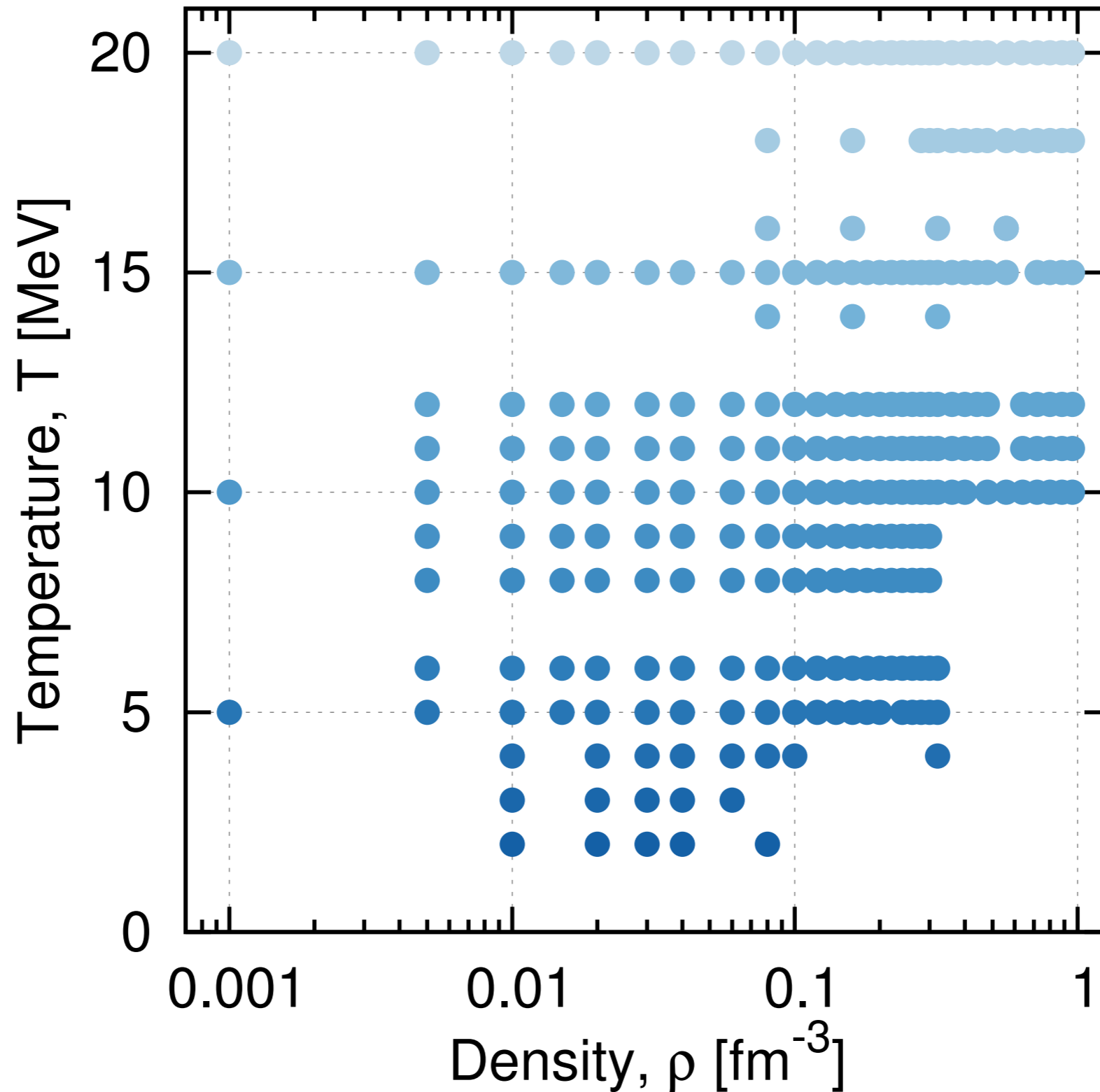
• Advantages

- **All** kinds of NN interactions ✓
- **3N** interactions ✓
- Short-range & tensor **correlations** ✓
- Density & **isospin** dependence ✓
- Access to **off-shell** spectral function ✓
- **Thermodynamically** consistent ✓

• Limitations

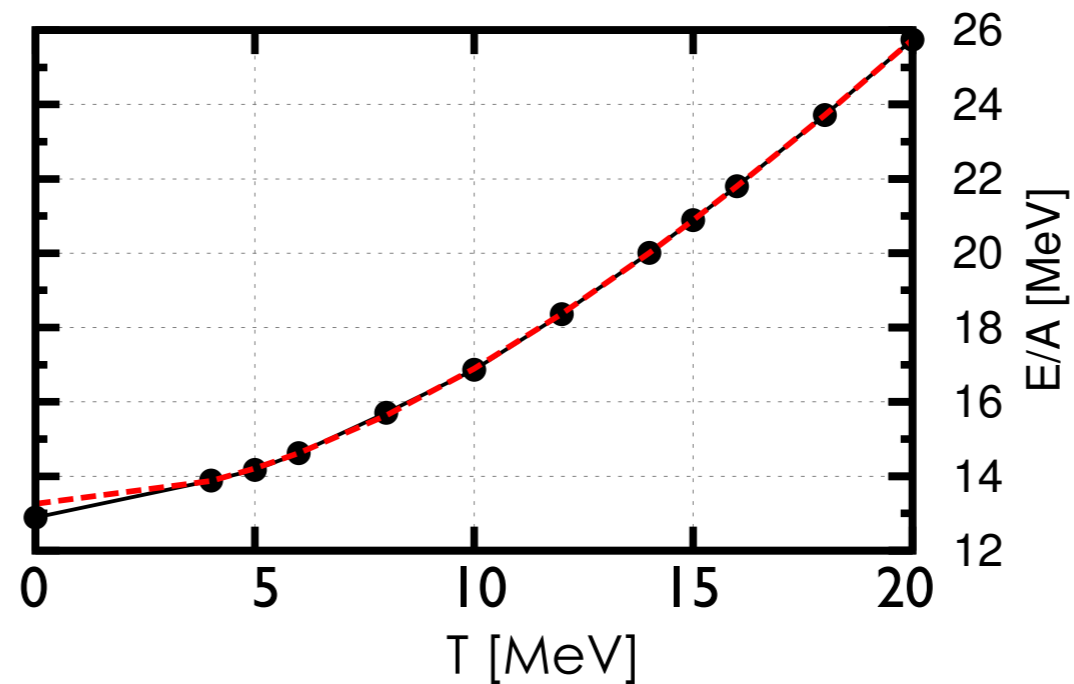
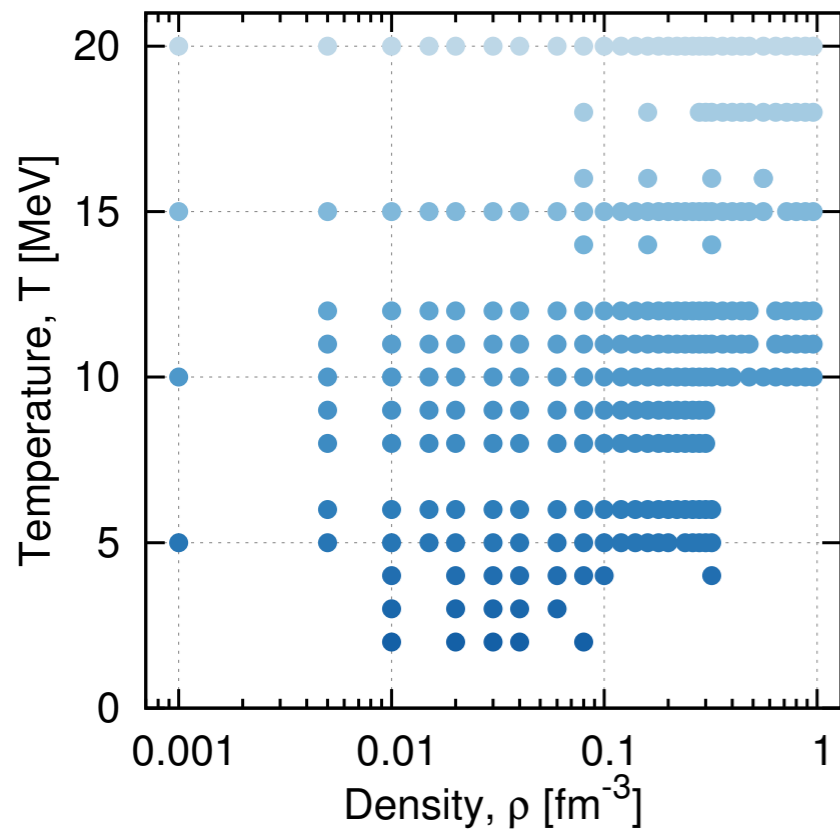
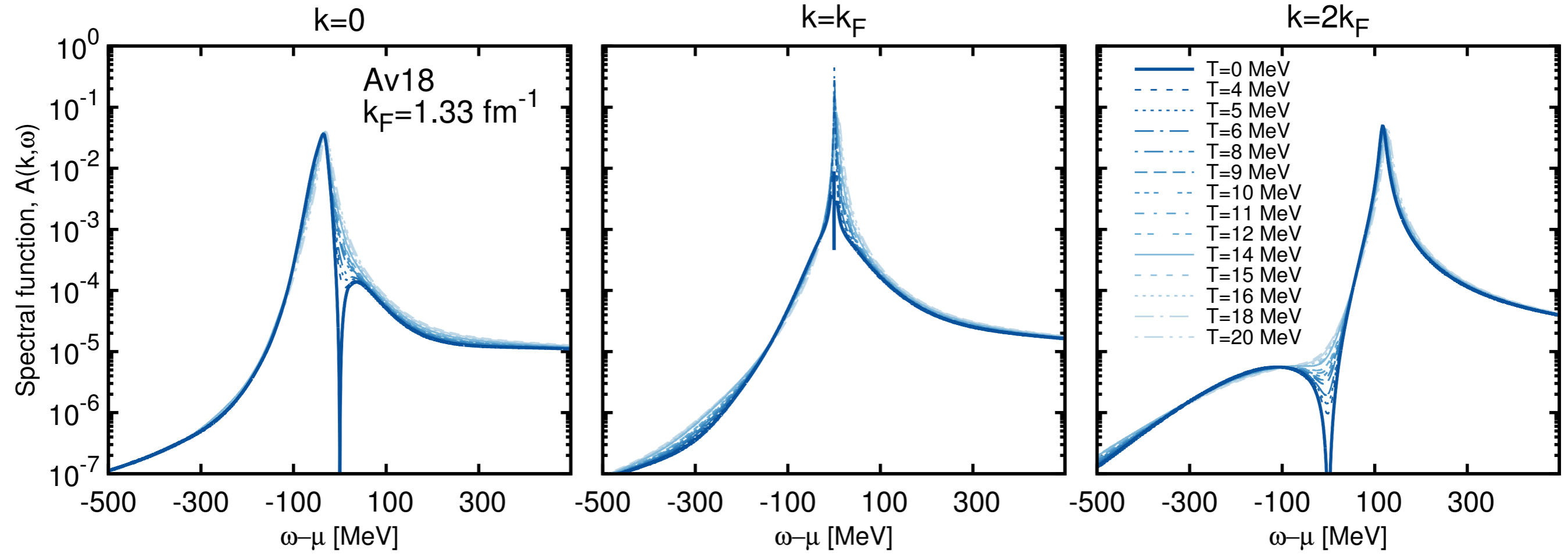
- Non-relativistic ✗
- Missing diagrams? ✗
- No nuclear surface ✗

Av18



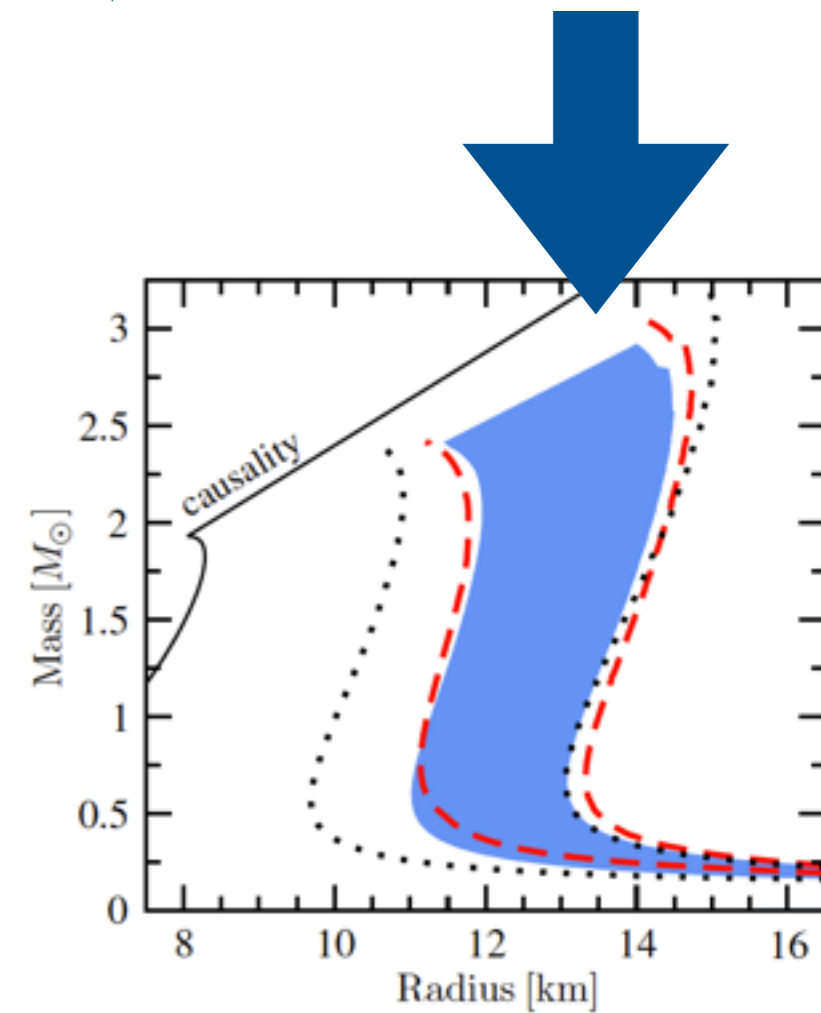
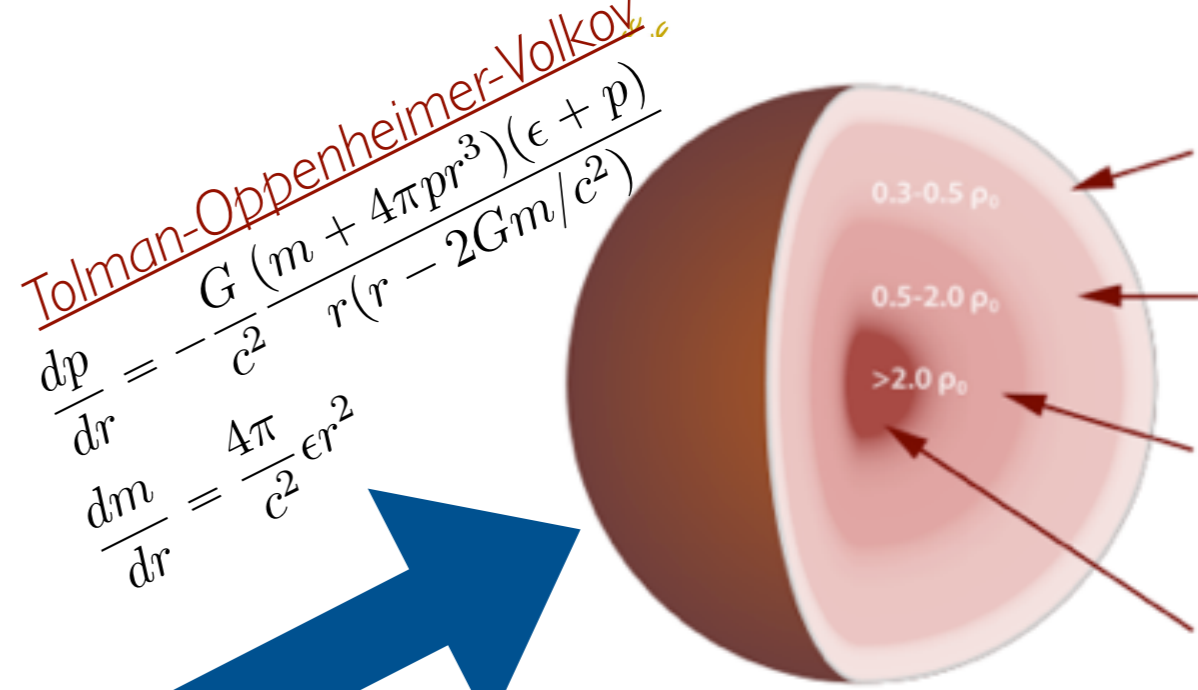
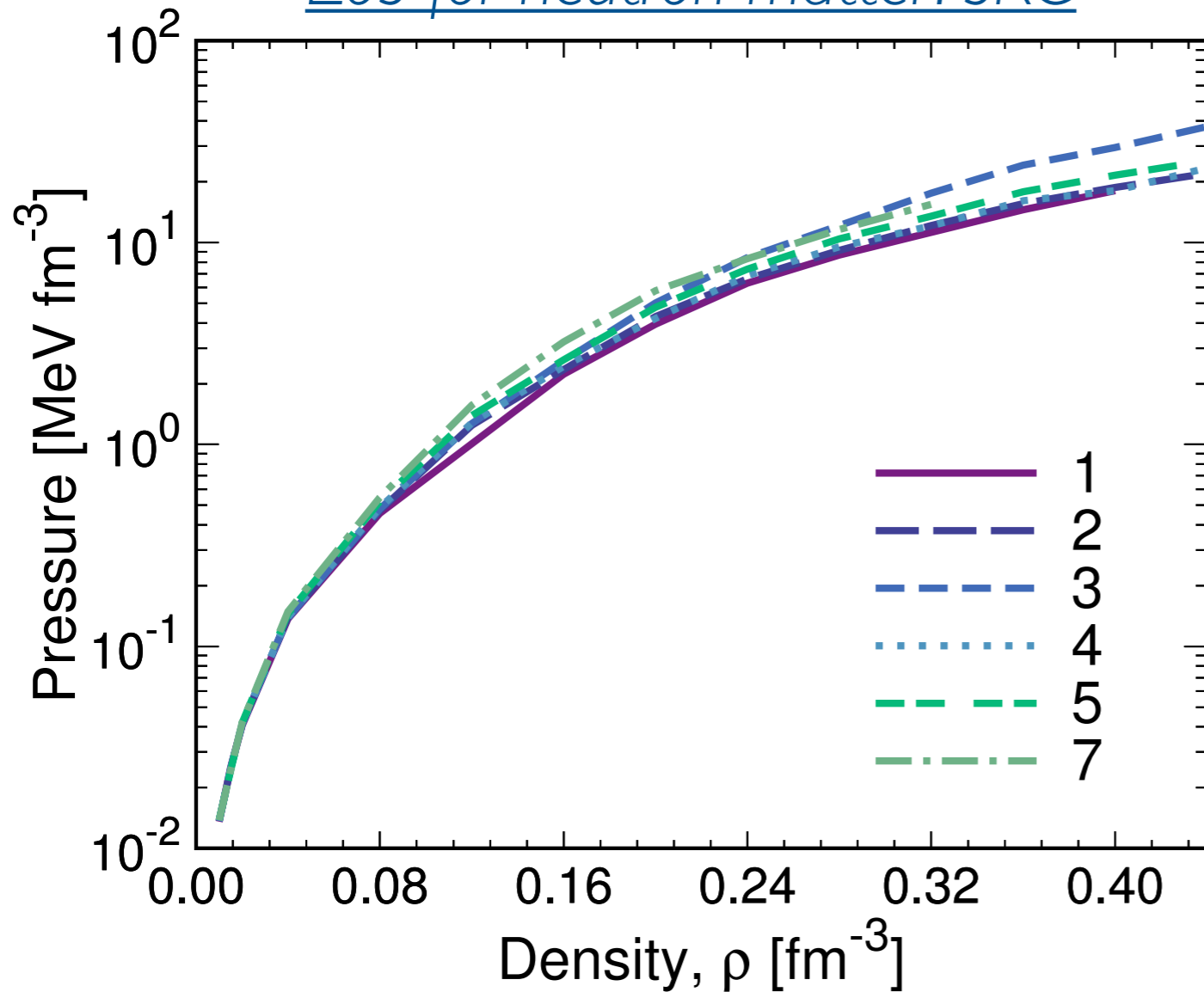
Self-energy, spectral function & thermodynamics

$T=0$ extrapolations



Neutron matter

EoS for neutron matter: SRG



Hebeler, Lattimer, Pethick, Schwenk
ApJ 773 11 (2013)

- Error band from fits in ChPT c_1, c_3 parameters
- Finite temperature & higher densities available

Carbone, Polls & Rios, PRC 90 054322 (2014)

A. Carbone, Drischler, Hebeler, Schwenk, arXiv:1608.05615 12

Momentum distribution

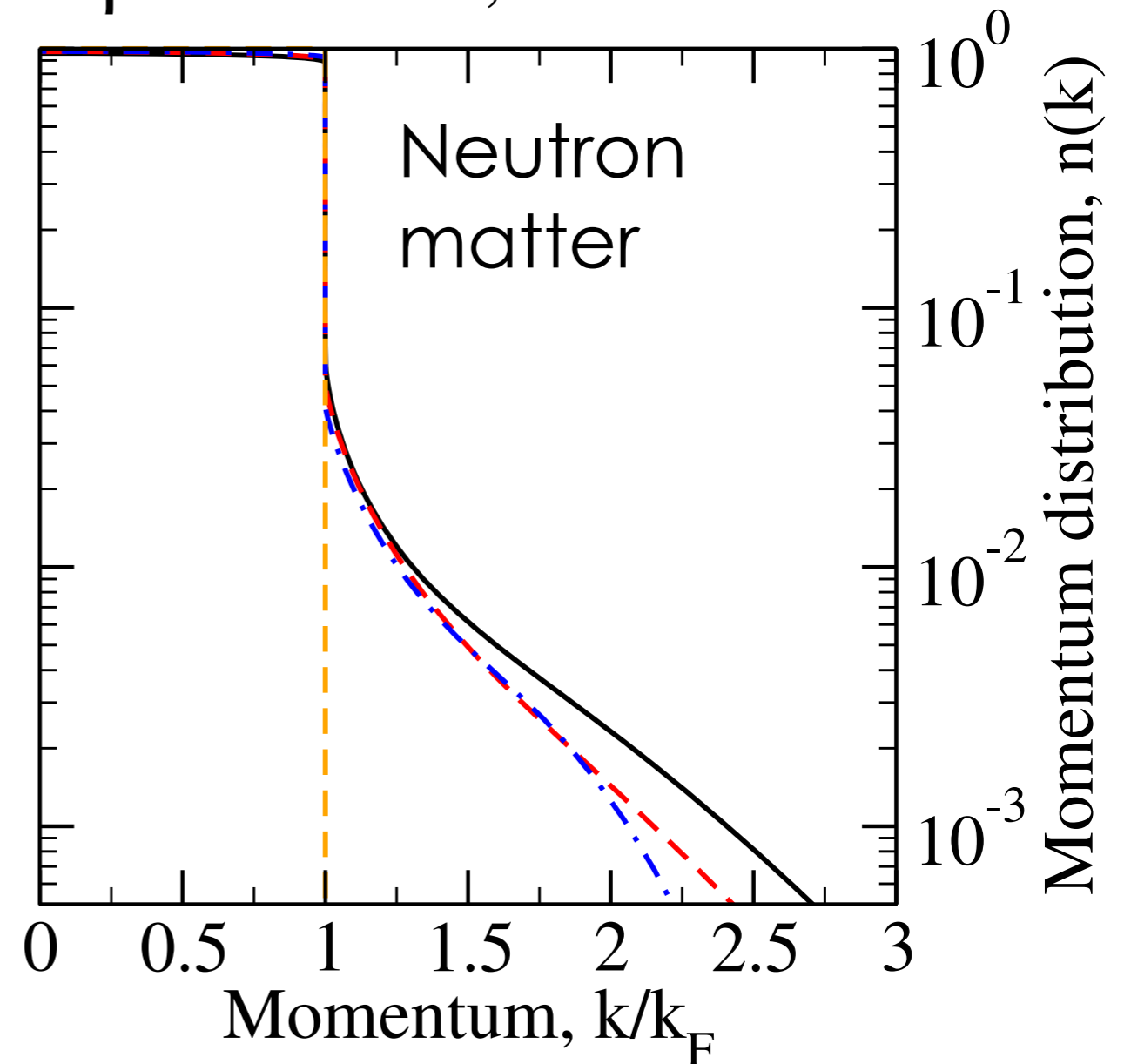
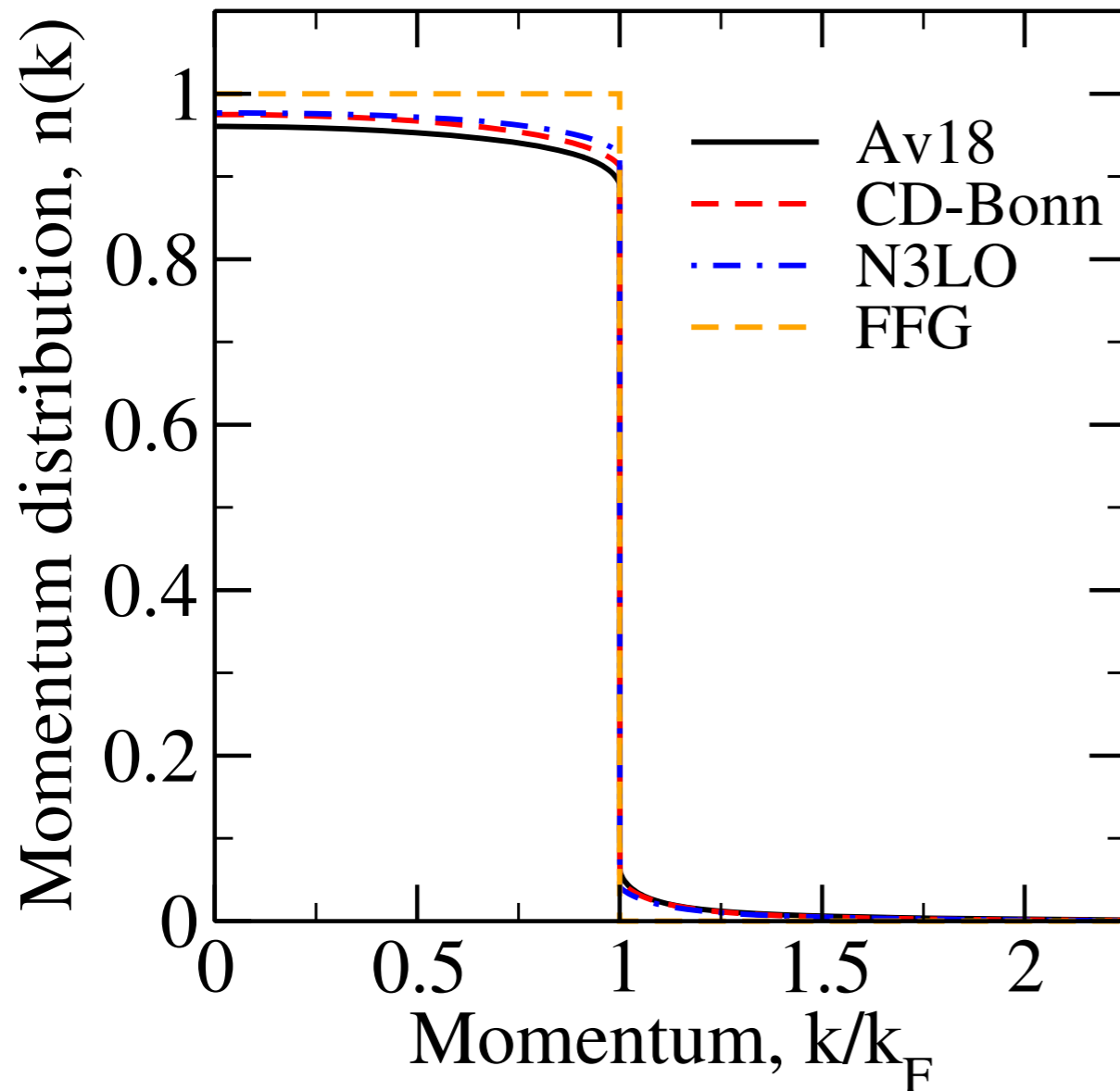
Single-particle occupation

$$n(k) = \langle a_k^\dagger a_k \rangle$$

$$\nu \int \frac{d^3 k}{(2\pi)^3} n(k) = \rho$$

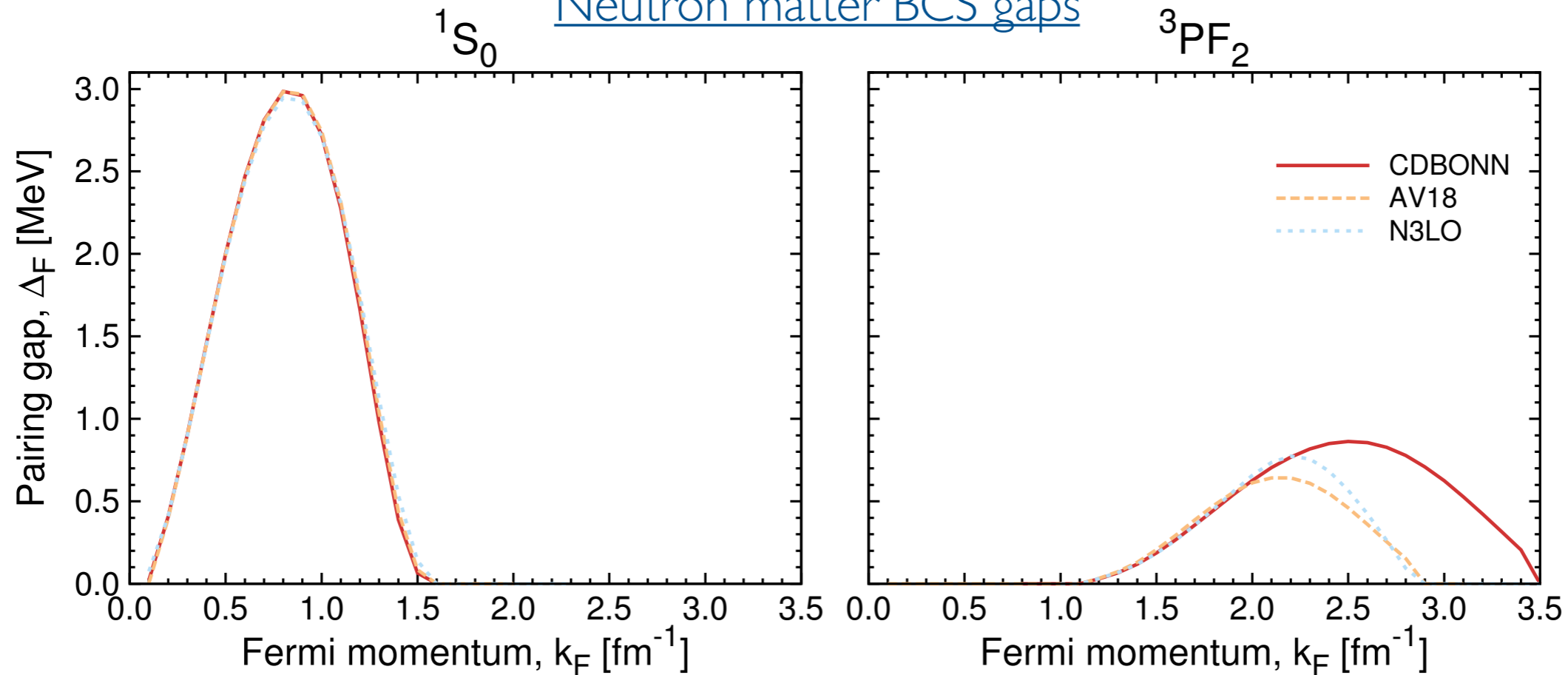
$\rho = 0.16 \text{ fm}^{-3}$, $T = 0 \text{ MeV}$

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- SNM: 11-13% depletion at low k , population at high k
- Dependence on NN interaction under control
- PNM: 4-5% depletion at low k

Neutron matter BCS gaps



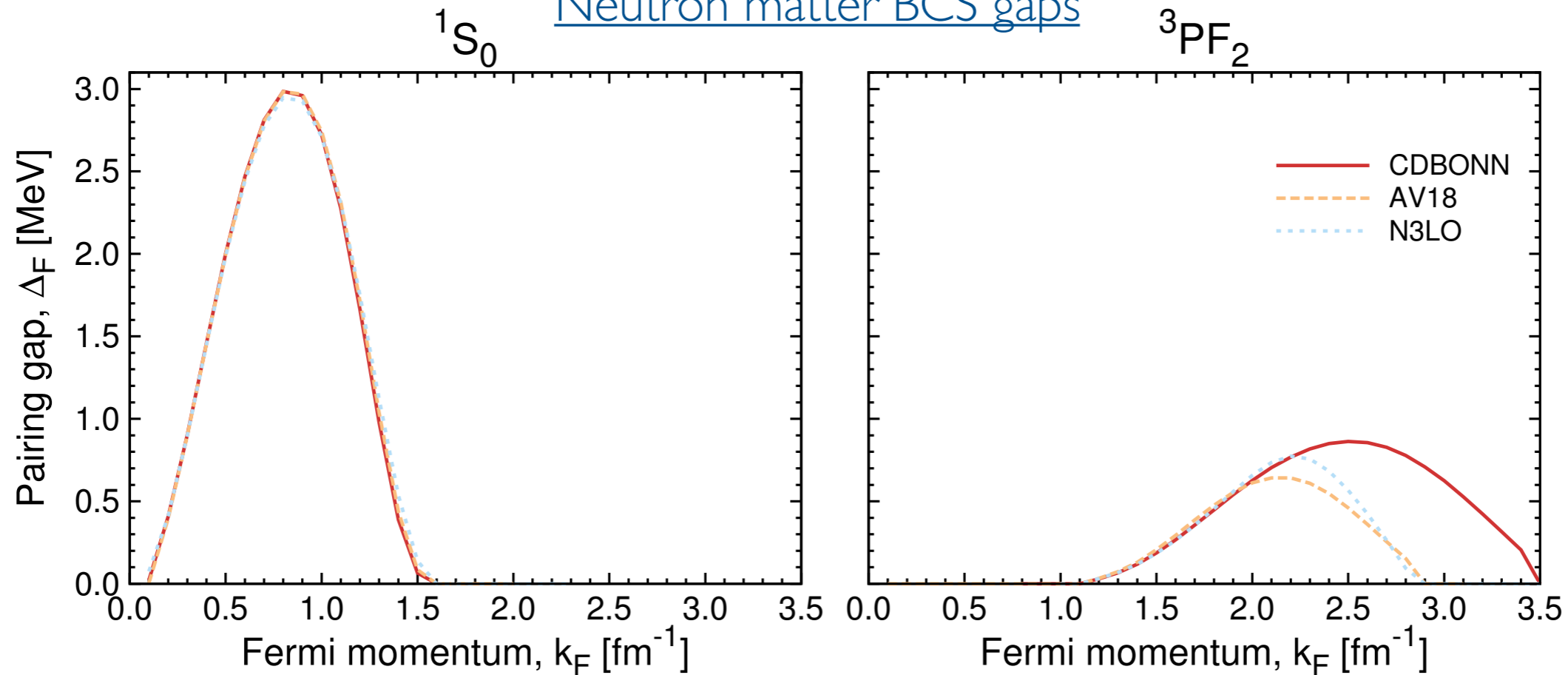
Dean & Hjorth-Jensen, *Rev. Mod. Phys.* **75** 607 (2003)

BCS equation

$$\Delta_k^L = - \sum_{L'} \int_{k'} \frac{\langle k | V_{nn}^{LL'} | k' \rangle}{2\sqrt{\chi_{k'}^2 + |\Delta_{k'}^{L'}|^2}} \Delta_{k'}^{L'} + \quad \chi_k = \varepsilon_k - \mu$$

- Single-particle spectrum choice: $\varepsilon_k = \frac{k^2}{2m} + U(k) - \mu$
- Angular gap dependence: $|\Delta_k|^2 = \sum_L |\Delta_k^L|^2$

Neutron matter BCS gaps



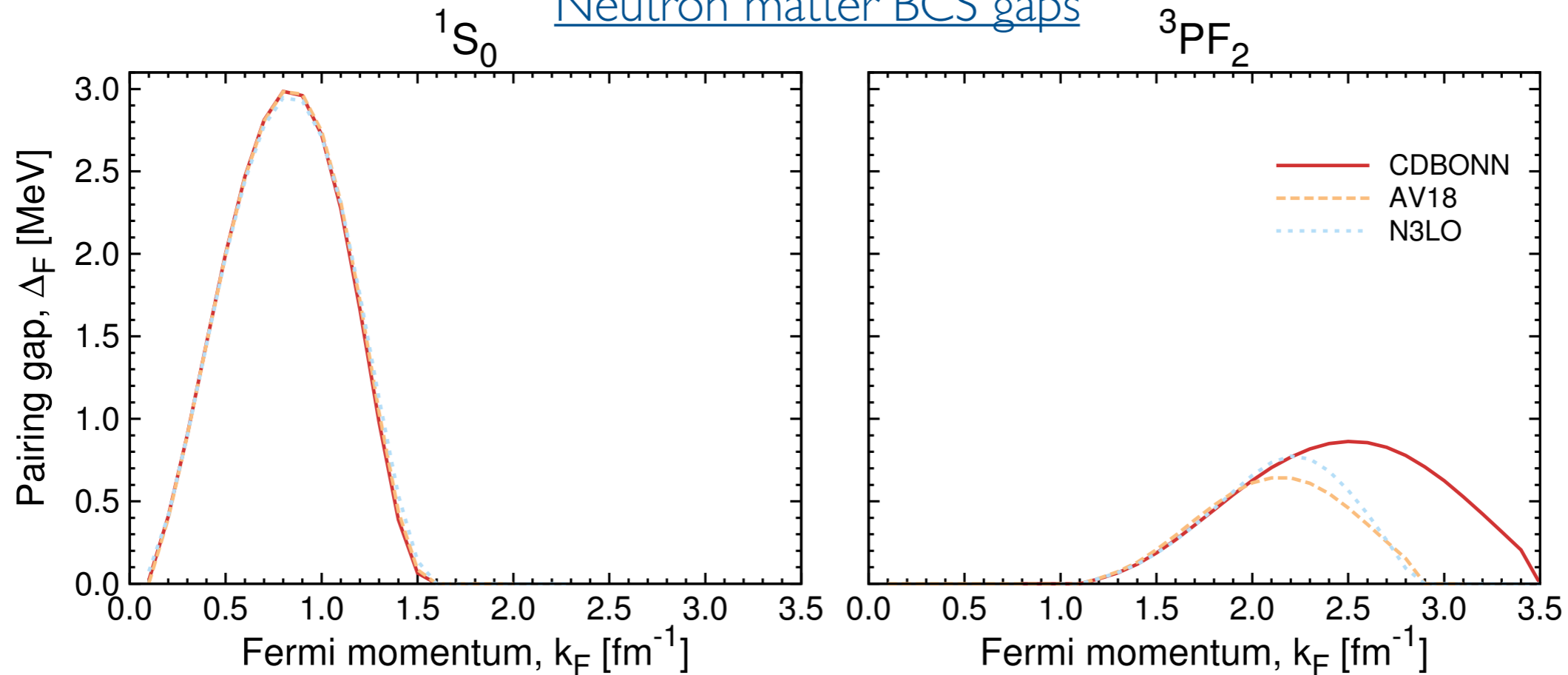
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Dean & Hjorth-Jensen, *Rev. Mod. Phys.* **75** 607 (2003)

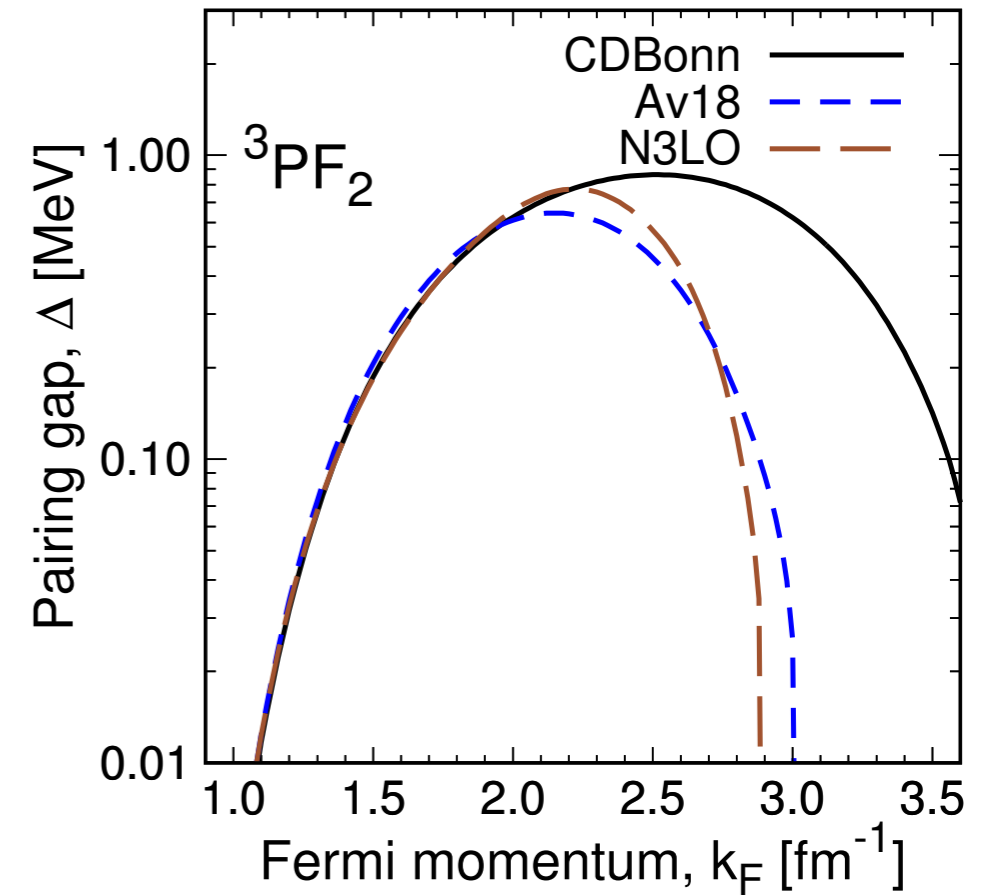
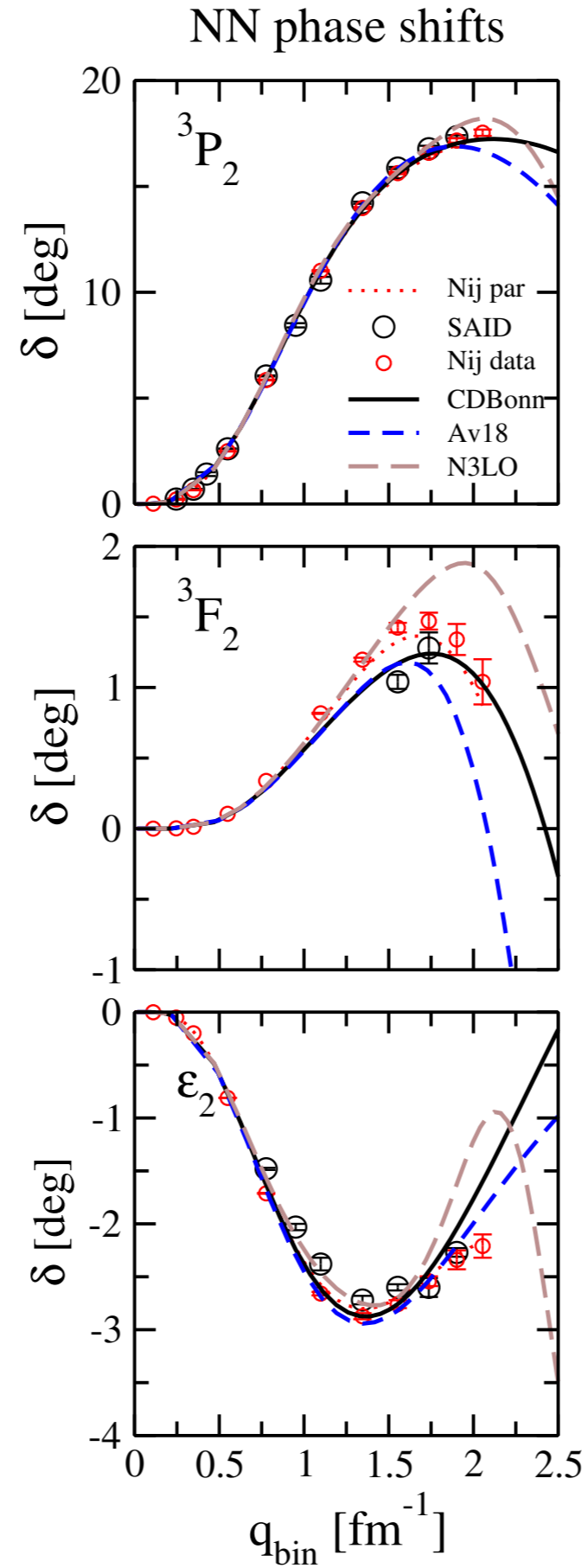
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- Angular gap dependence: $|\Delta_k|^2 = \sum_L |\Delta_k^L|^2 \approx |\Delta_k^L|^2$

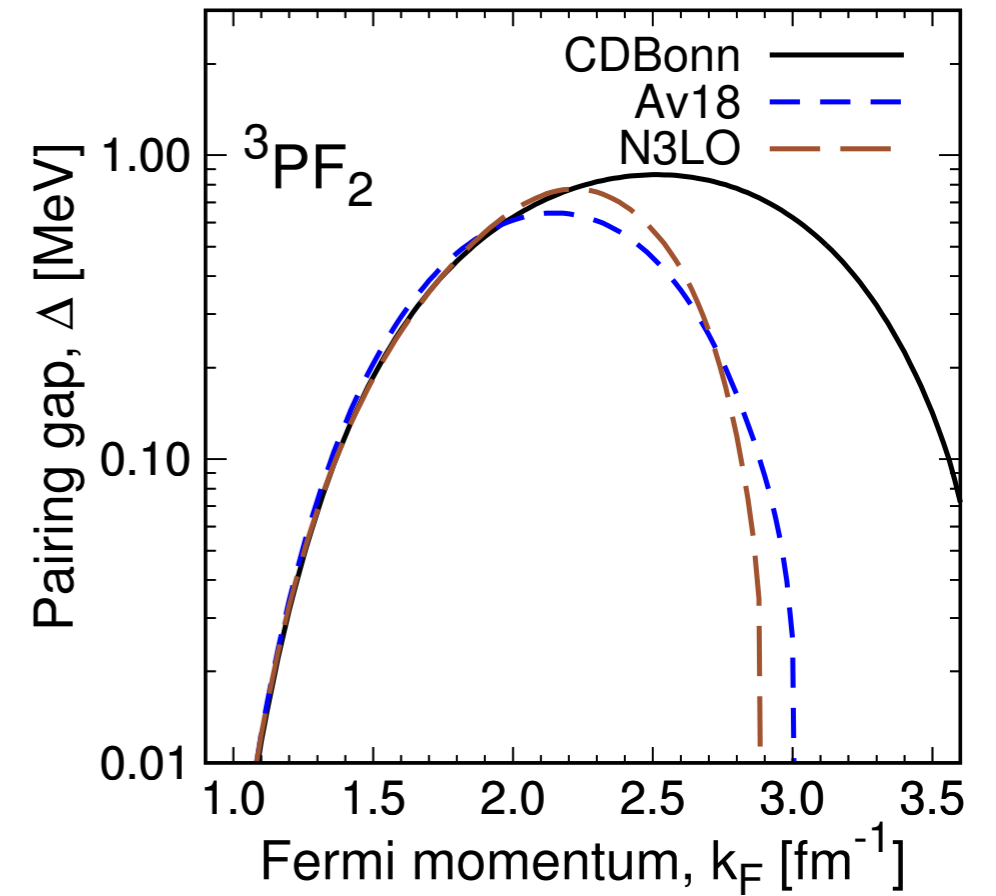
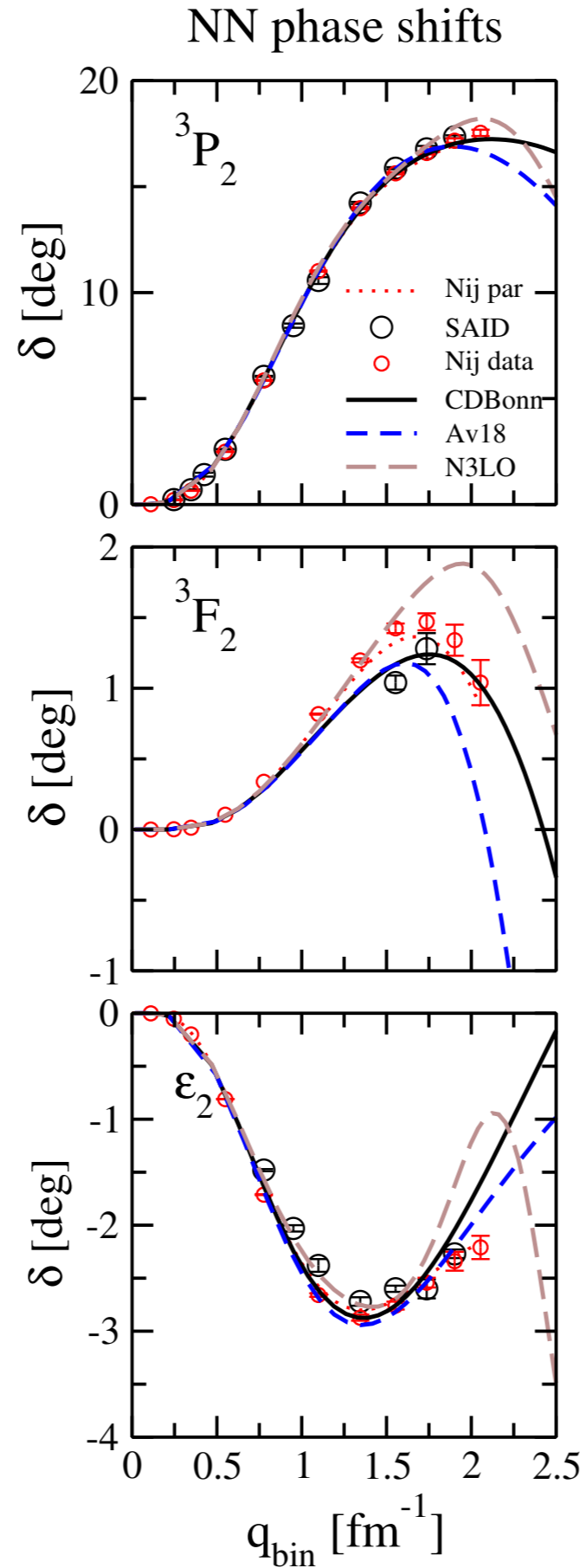
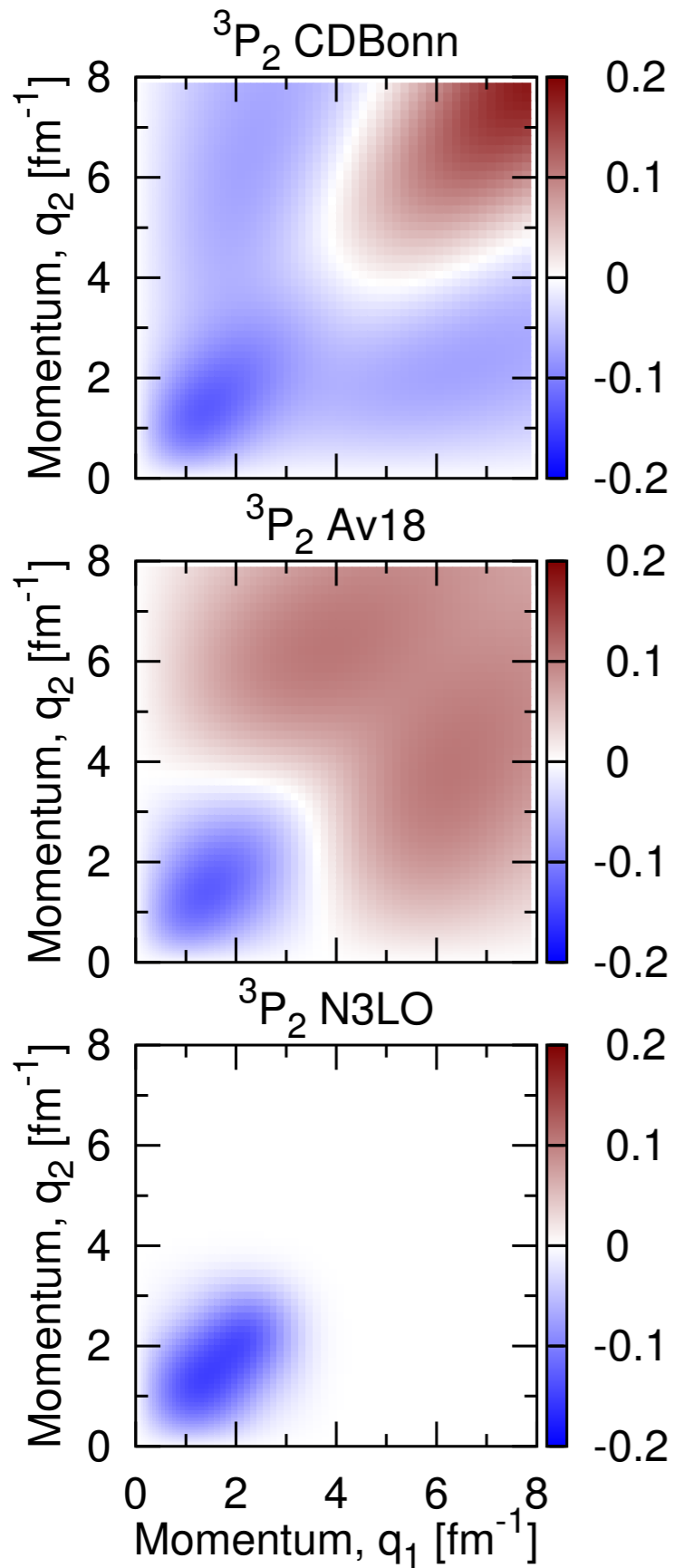
Triplet pairing

Phase-shift equivalent?



Triplet pairing

Phase-shift equivalent?



(B) $iF(1, 2) = \langle T\{\psi(1)\psi(2)\} \rangle = \longleftrightarrow = \rightarrow \textcircled{\Sigma} \longleftrightarrow + \rightarrow \textcircled{\Delta^*} \rightarrow$

(B') $iF^\dagger(1, 2) = \langle T\{\psi^\dagger(1)\psi^\dagger(2)\} \rangle = \longleftrightarrow = \rightarrow \textcircled{\Sigma} \longleftrightarrow + \rightarrow \textcircled{\Delta} \leftarrow$

(C) $iG(1, 2) = \langle T\{\psi(1)\psi^\dagger(2)\} \rangle = \longleftrightarrow = \rightarrow + \rightarrow \textcircled{\Sigma} \longleftrightarrow + \rightarrow \textcircled{\Delta} \leftarrow$

(D) $\rightarrow \textcircled{\Sigma} \rightarrow = \rightarrow \textcircled{K} \rightarrow$

(E) $\leftarrow \textcircled{\Delta} \leftarrow = \leftarrow \textcircled{K} \leftarrow$

Normal
state

Superfluid
 $\Delta(k_F)$

BCS+SRC equation

$$\Delta_k^L = - \sum_{L'} \int_{k'} \frac{\langle k | V_{nn}^{LL'} | k' \rangle}{2\sqrt{\bar{\chi}_{k'}^2 + |\Delta_{k'}|^2}} \Delta_{k'}^{L'} + \frac{1}{2\chi_k} = \int_{\omega} \int_{\omega'} \frac{1 - f(\omega) - f(\omega')}{\omega + \omega'} A(k, \omega) A_s(k, \omega')$$

- **BCS** is lowest order in Gorkov Green's function expansion
- T-matrix can be extended to paired systems
- But full self-consistency is still missing

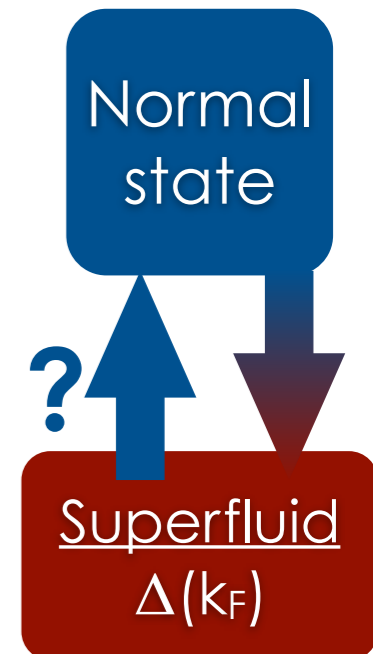
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(D) $\text{---} \Sigma \text{---} = \text{---} \text{K} \text{---}$

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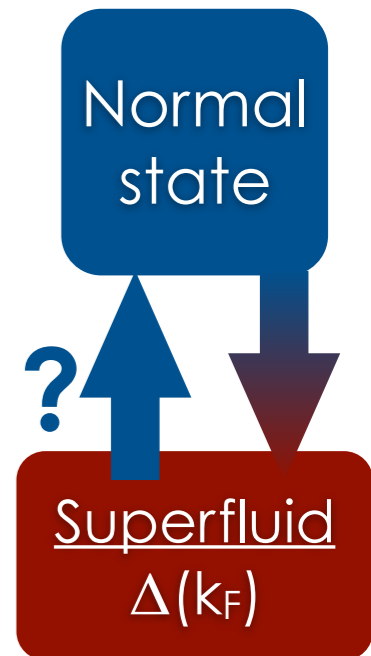
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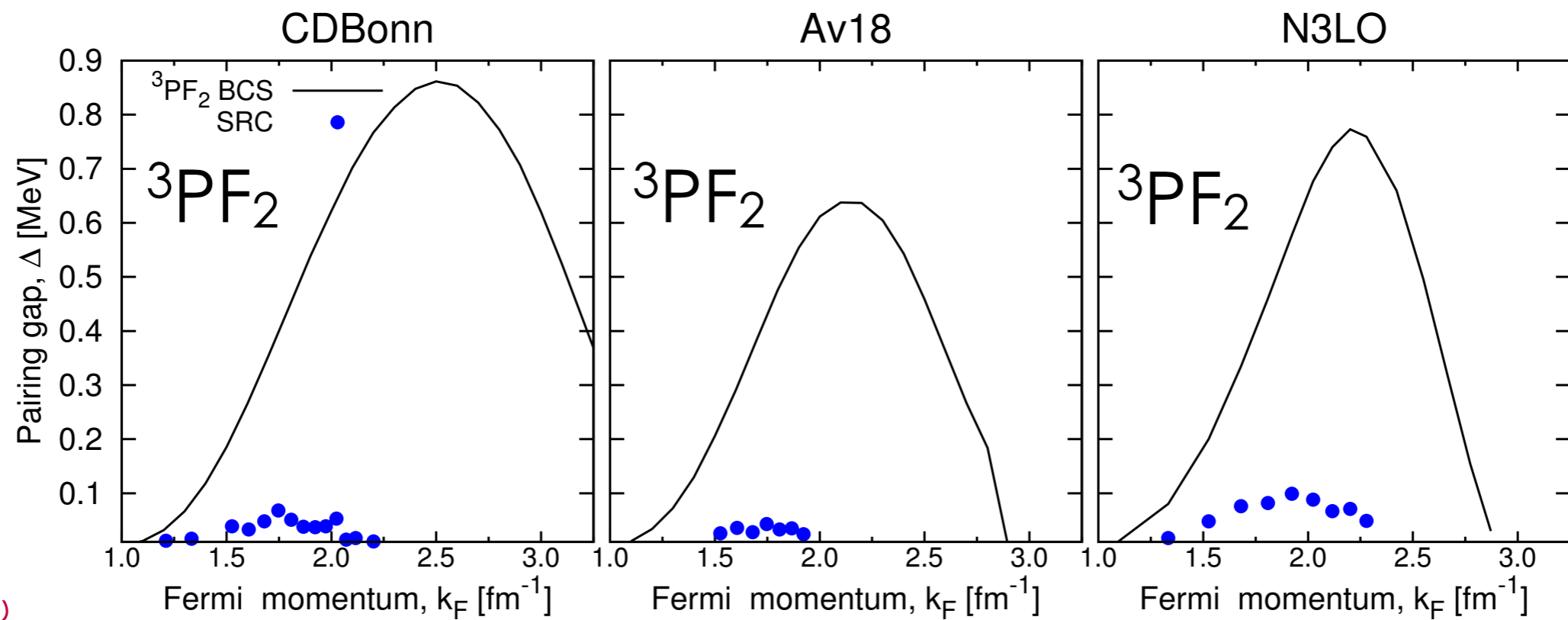
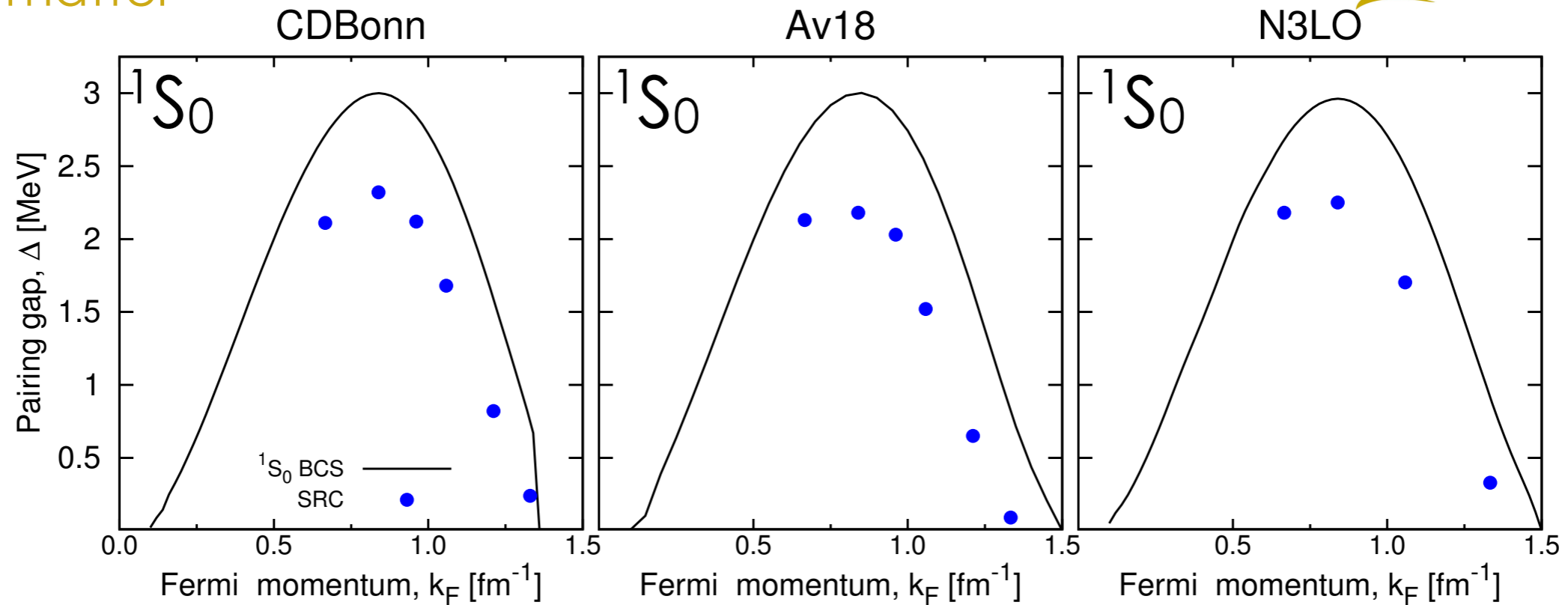
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Beyond BCS 101: SRC

Neutron matter



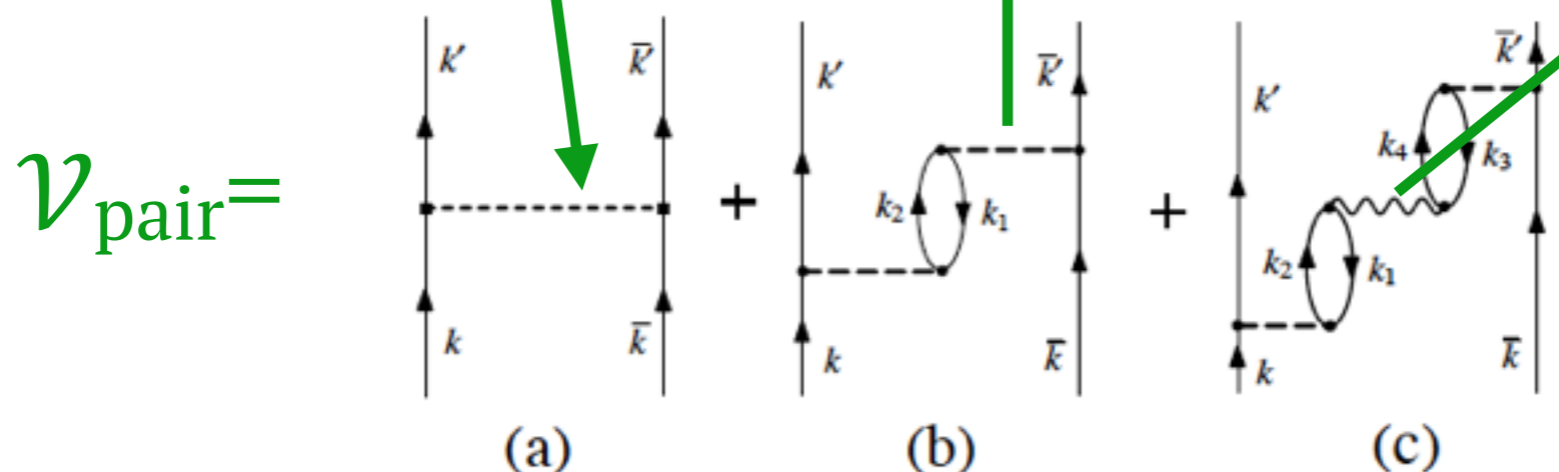
BCS+SRC

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ph recoupled G-matrix

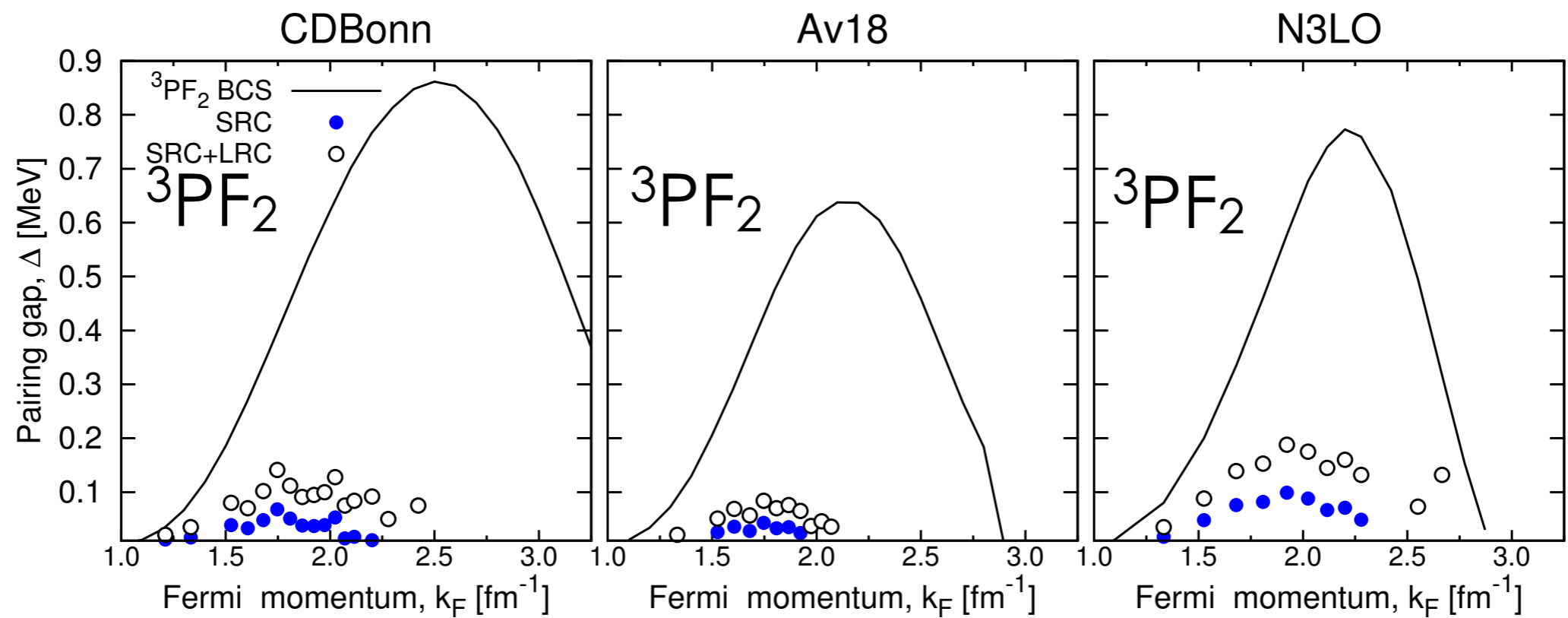
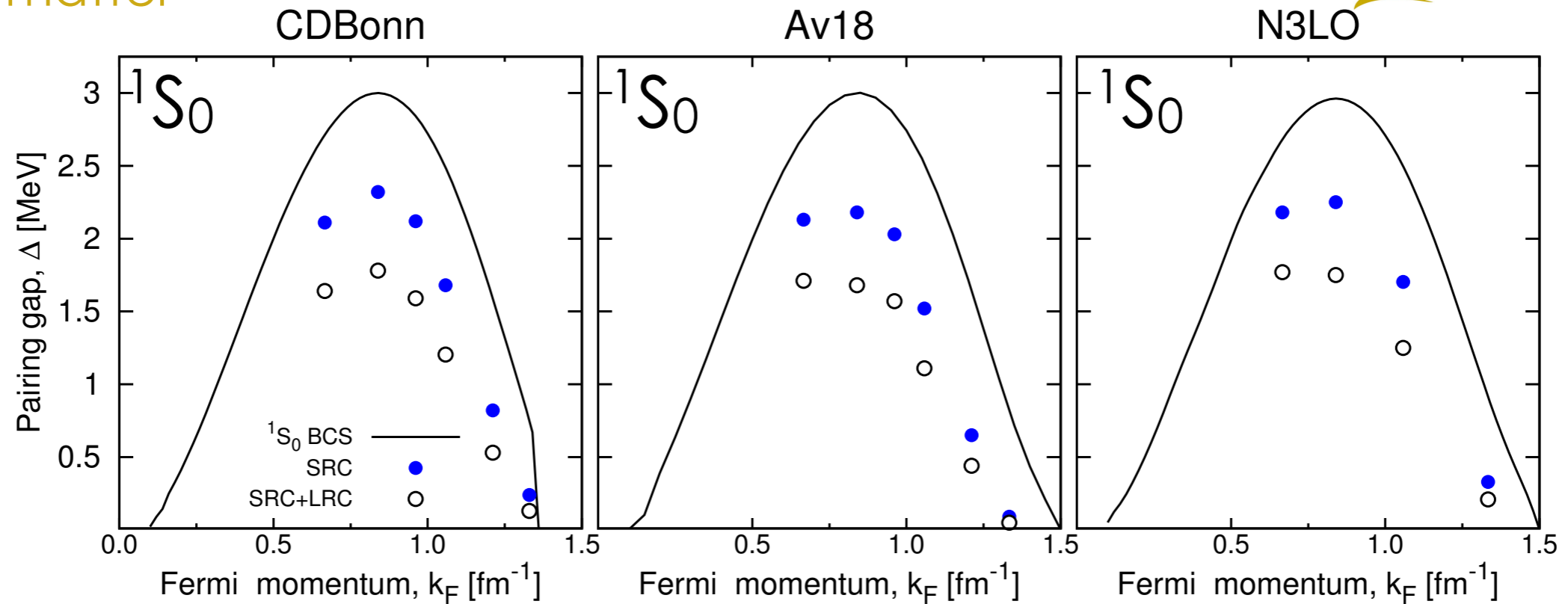
Effective Landau parameters



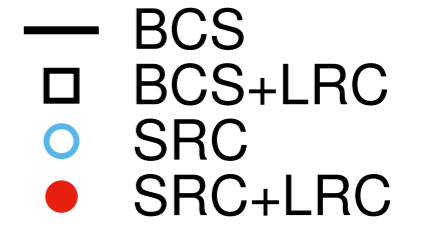
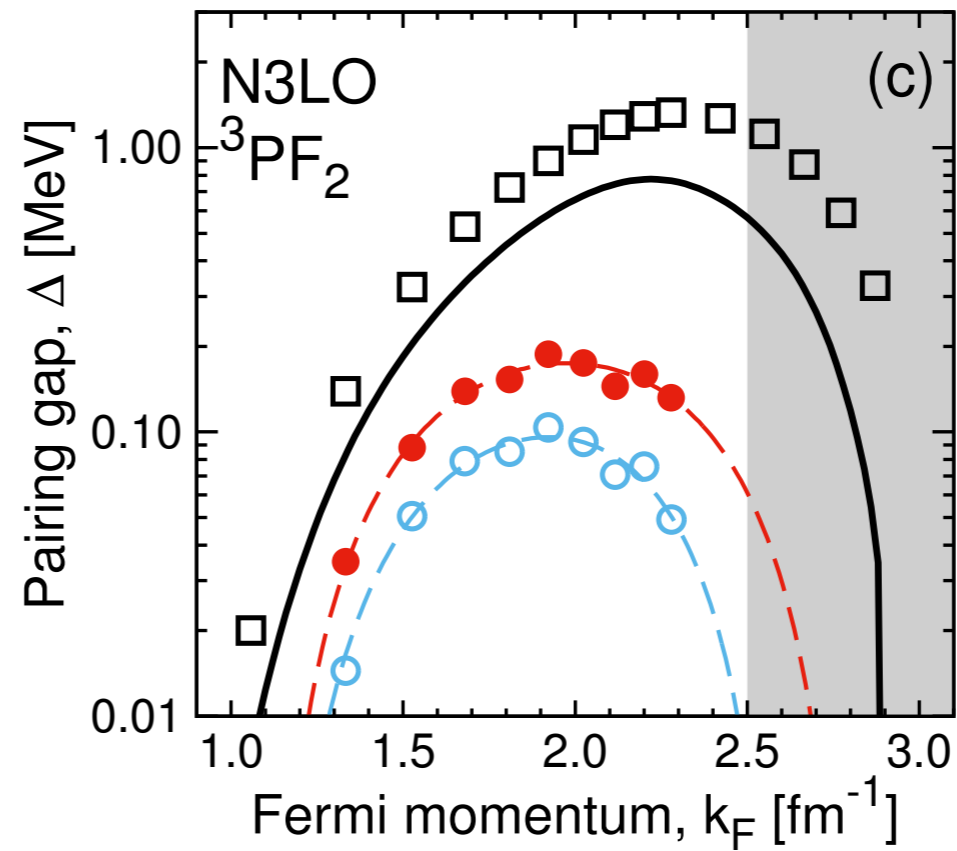
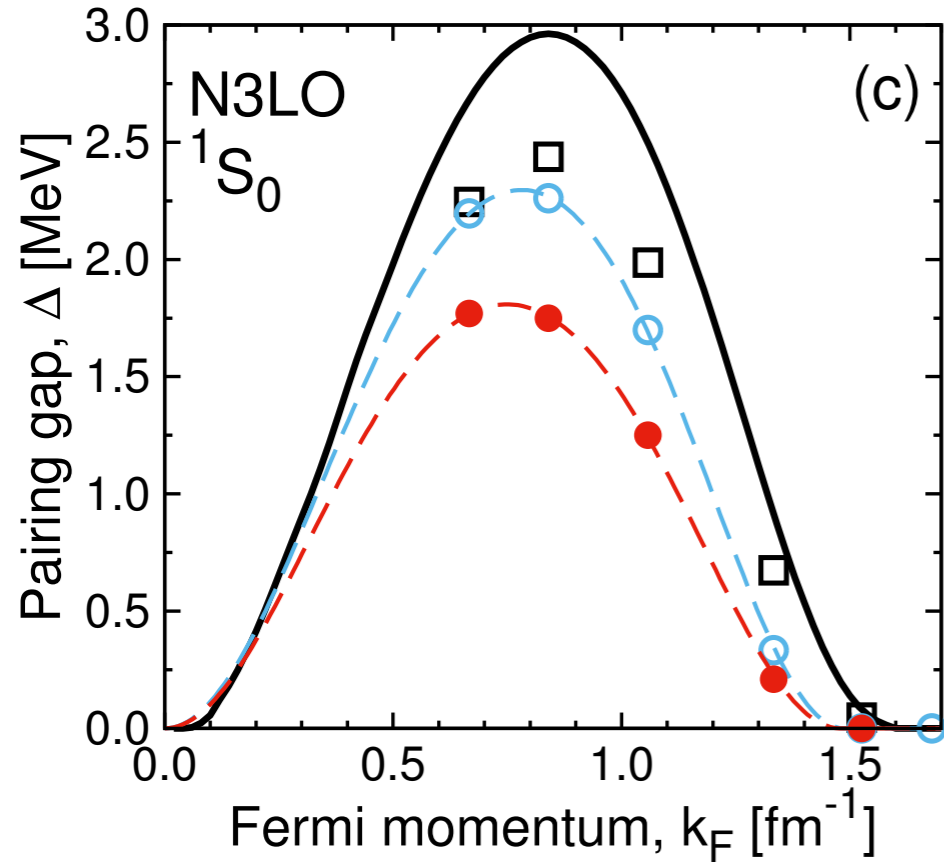
$$\langle 1\bar{1} | \mathcal{V} | 1\bar{1} \rangle = \frac{1}{4} \sum_{2,2'} \sum_{S,T} (-)^S (2S+1) \langle 12 | G_{ST}^{\text{ph}} | 1'2' \rangle_A \langle 2'\bar{1} | G_{ST}^{\text{ph}} | 2\bar{1}' \rangle_A \Lambda(22')$$

$$\Lambda_{ST}(q) = \frac{\Lambda_{ST}^0(q)}{1 - \Lambda_{ST}^0(q) \times F_{ST}}$$

- Bare NN potential only is not the only possible interaction
- Diagram (a): nuclear interaction
- Diagram (b): in-medium interaction, density and spin fluctuations
- Diagram (c): included by Landau parameters

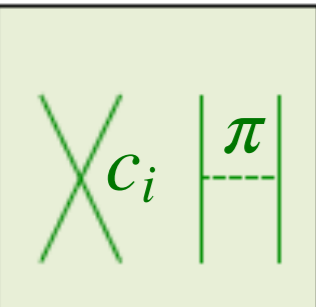
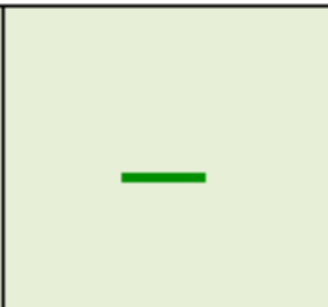
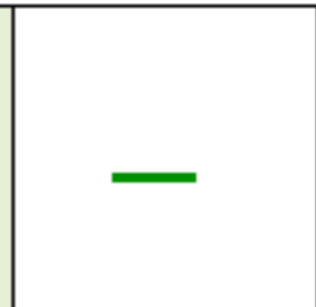
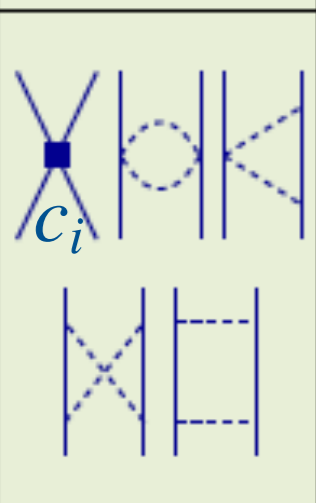
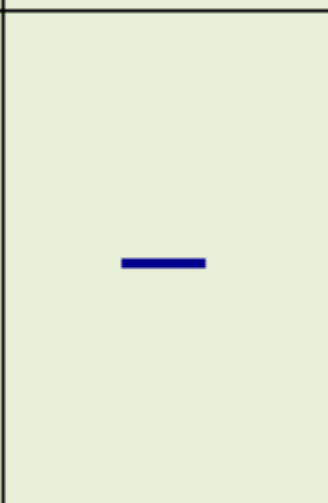

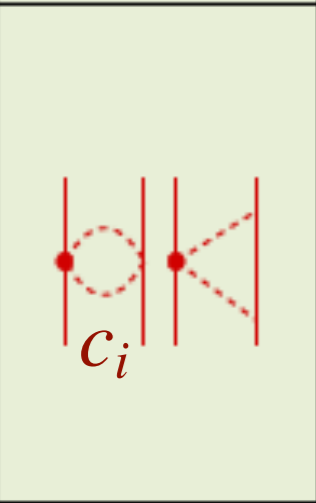
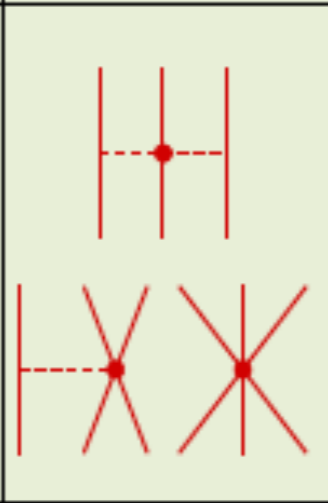

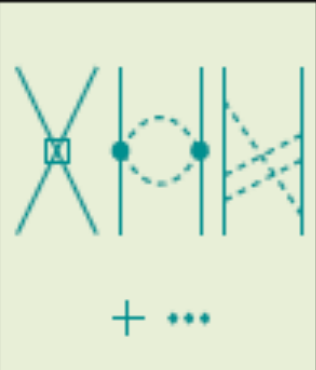




- LRC 1S_0 (3PF_2) produces (anti-)screening



$$\Delta_L^{JST}(k_F) = \Delta_0 \frac{(k_F - k_0)^2}{(k_F - k_0)^2 + k_1} \frac{(k_F - k_2)^2}{(k_F - k_2)^2 + k_3}$$

		Δ_0	k_0	k_1	k_2	k_3
Singlet	CDBonn	18.18	0.05	1.39	1.45	0.81
	Av18	14.07	0.04	1.00	1.44	0.78
	N3LO	5.85	0	0.46	1.48	0.42
Triplet	CDBonn	0.41	1.03	0.56	2.81	1
	Av18	0.17	1.1	0.35	2.18	0.05
	N3LO	0.6	1.11	0.69	2.79	0.53

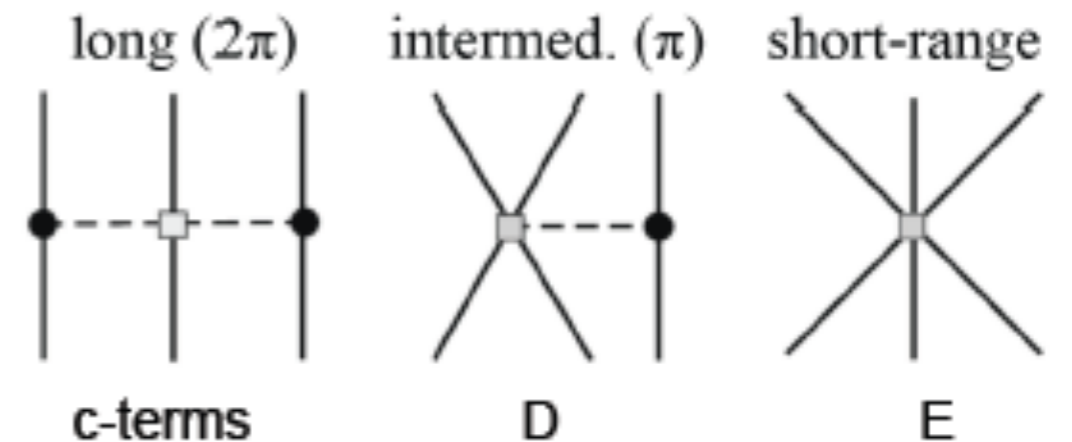
	NN	3N	4N
LO			
NLO			
N ² LO			
N ³ LO			

Chiral perturbation theory

- π and N as dof
- Systematic expansion
- 2N at N³LO - LECs from π N, NN
- 3N at N²LO - 2 more LECs
- (Often further renormalized)

$$\mathcal{O}\left(\frac{Q}{\Lambda}\right)$$

$$\Lambda \sim 1 \text{ GeV}$$

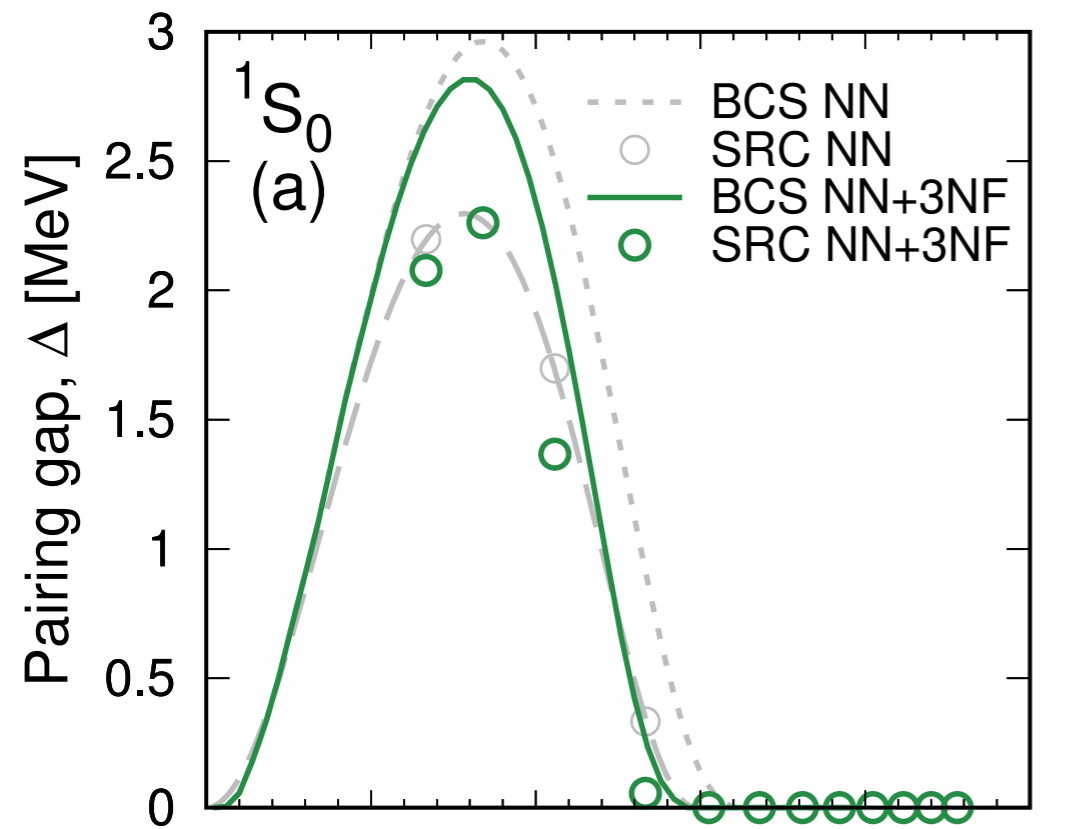


Weinberg, *Phys. Lett. B* **251** 288 (1990), *Nucl. Phys. B* **363** 3 (1991)

Entem & Machleidt, *Phys. Rev. C* **68**, 041001(R) (2003)

Tews, Schwenk et al., *Phys. Rev. Lett.* **110**, 032504 (2013) 21

3BF effect: estimate

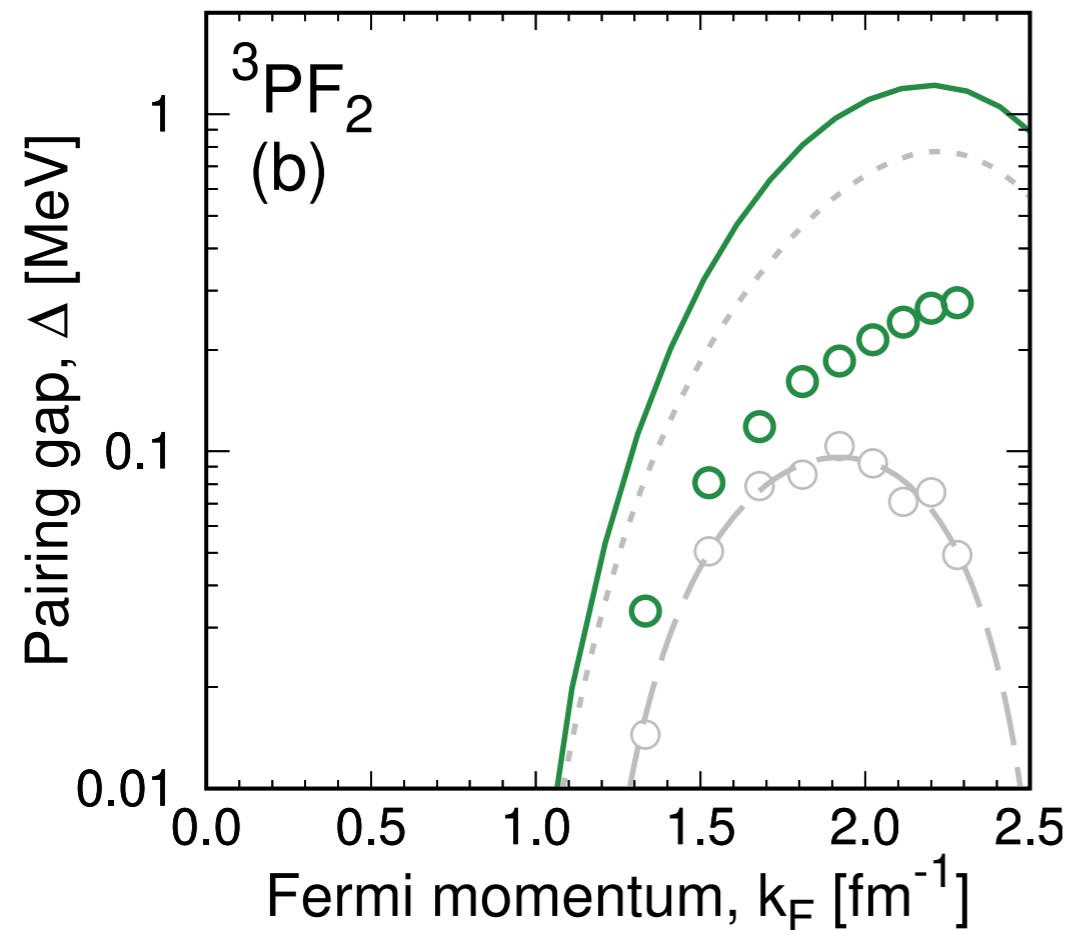


Effective one-body force \Rightarrow spectrum

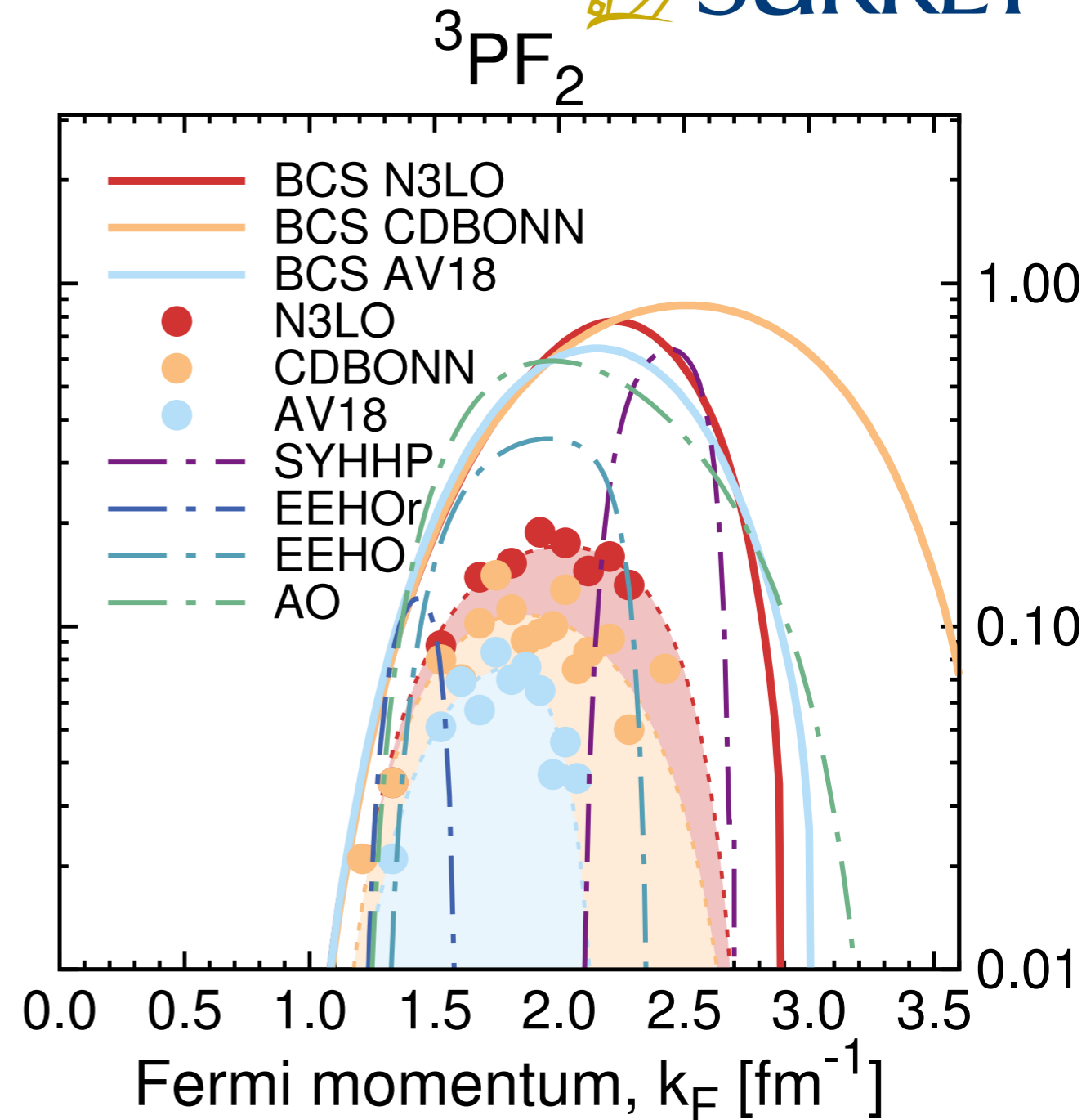
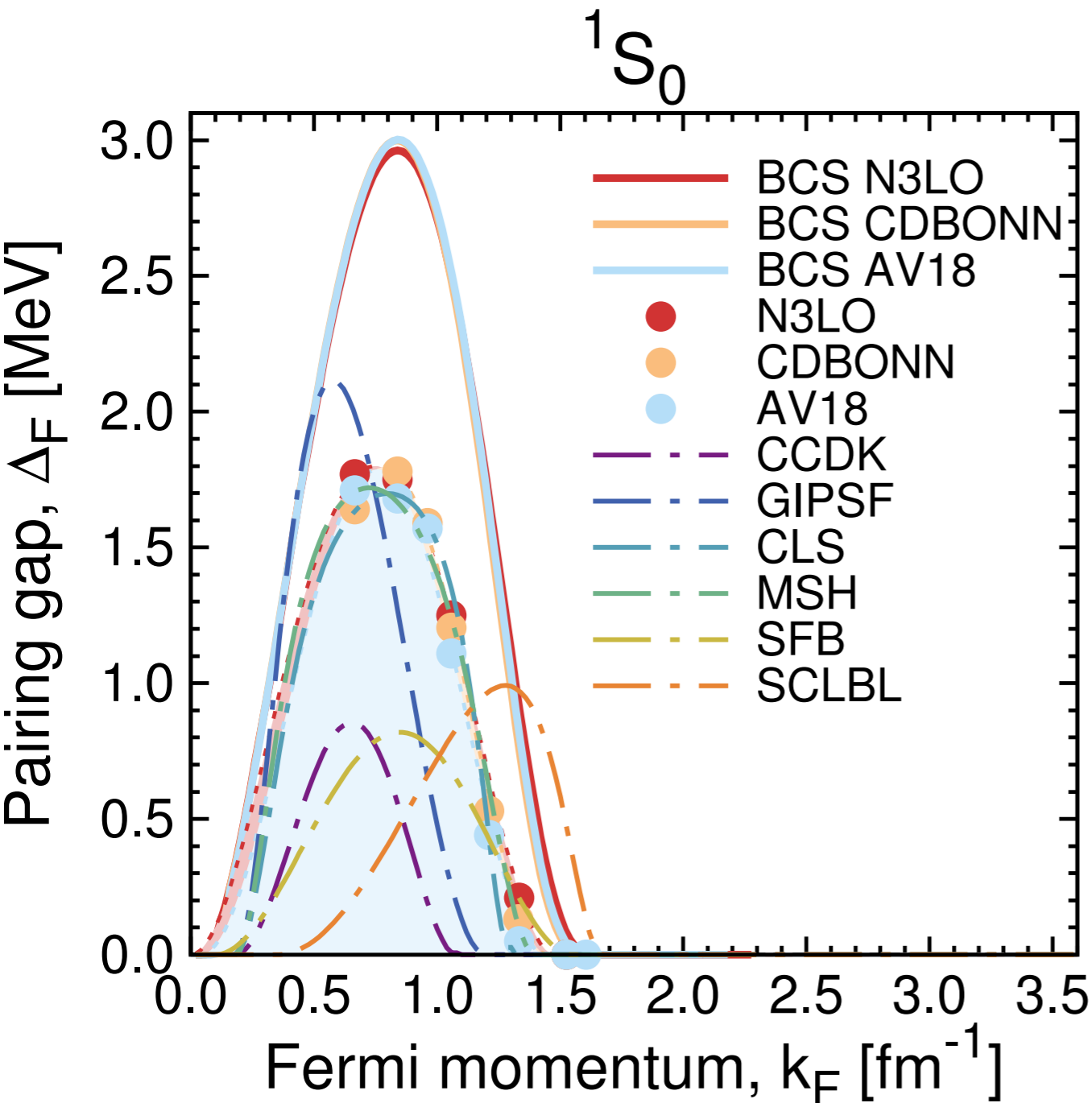
$$\bullet \text{---} \times = \bullet \text{---} \bullet \text{---} \bullet + \frac{1}{2} \bullet \text{---} \bullet \text{---} \bullet \text{---} \bullet$$

Effective two-body force \Rightarrow NN forces

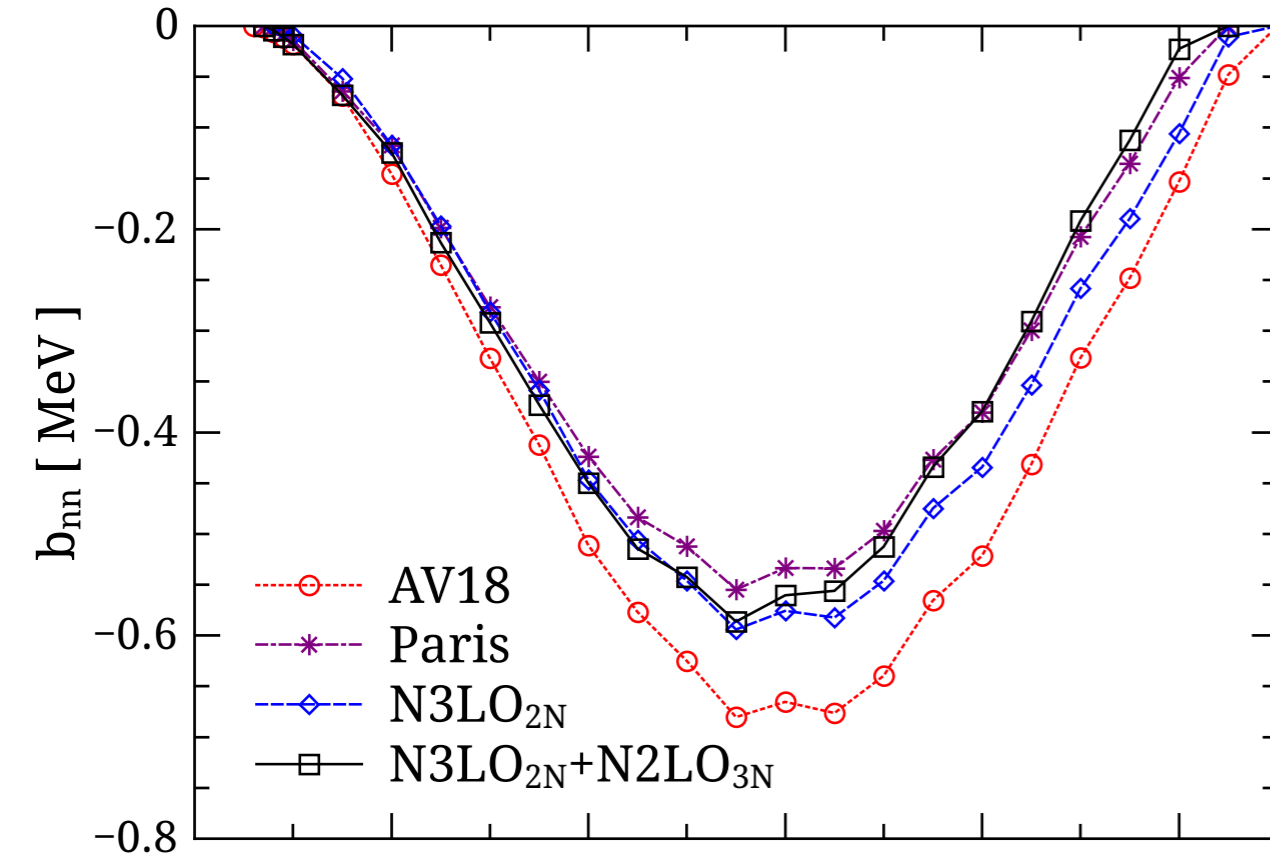
$$\bullet \text{---} \bullet \text{---} \bullet = \bullet \text{---} \bullet + \bullet \text{---} \bullet \text{---} \bullet$$



- Singlet gap: 3NF **reduce** closure
- Triplet gap: 3NF **increase** gap
- **Model dependence** to be explored
- SRG dependence for systematics



- Effect is **robust**: independent of NN potential
- 3NF effect **not** included in SRC, BCS indicates **small**
- **Singlet** channel under control in **astrophysical** situations



BHF G-matrix

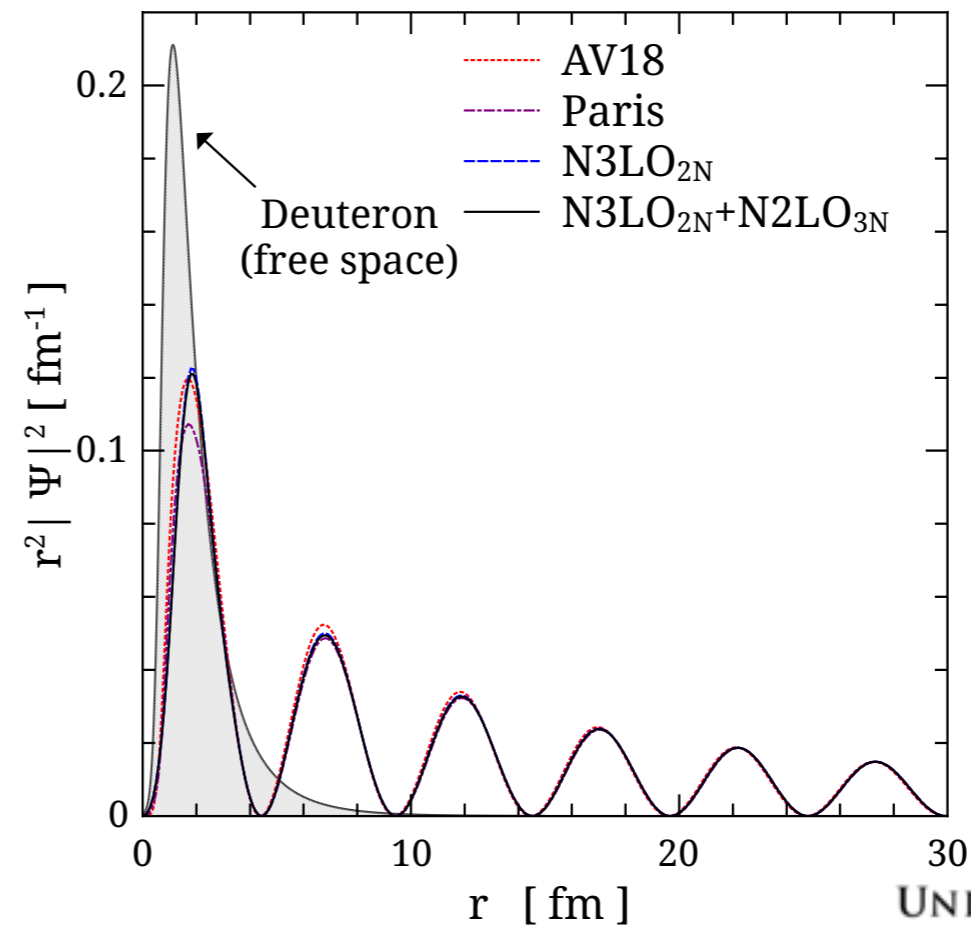
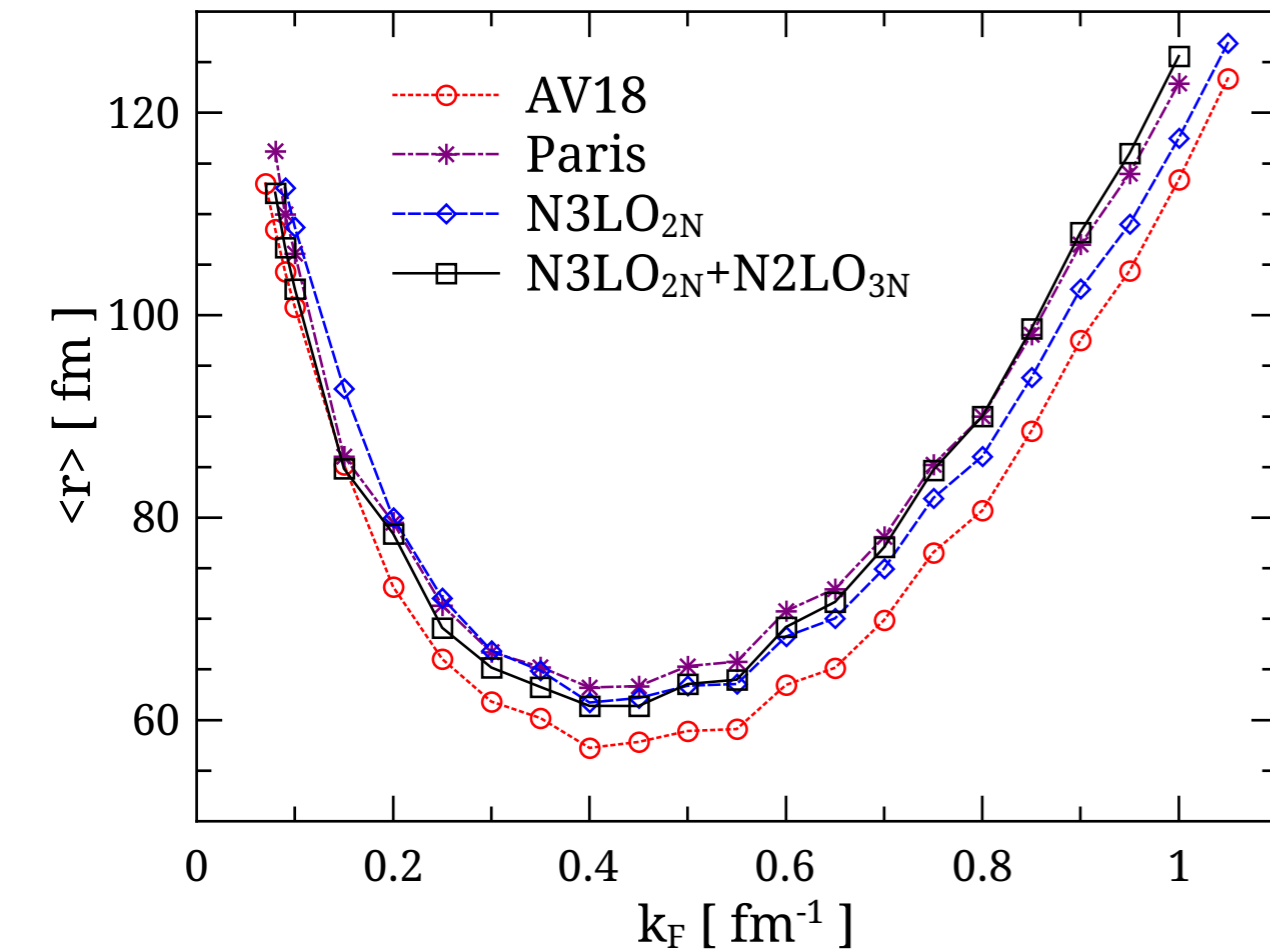
$$\mathcal{G}(\omega) = \mathcal{V} + \mathcal{V}Q(\omega)\mathcal{G}(\omega)$$

Poles

$$\det [1 - \mathcal{V}Q(\omega_i)] = 0$$

$$b_{nn} = \omega_i - \omega_{th}$$

$$k_F = 0.6 \text{ fm}^{-1}$$



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UNIVERSIDAD DE CHILE

- Ab initio nuclear **theory** to treat **correlations**
- **Talk** to us if you need **quantitative** predictions!
- **Different** NN forces give robust predictions
- Challenges ahead:
 - **Pairing** in isospin **asymmetric** matter
 - **Consistent** treatment of **cooling, glitch & EoS**
 - Improvements of **many-body theory**



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