Cosmic Superstrings in Warped Spacetime

Sarah Chadburn Durham University Work with Ruth Gregory & Tasos Avgoustidis

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Outline

Why study cosmic strings?

Warped spacetime

Modelling cosmic string motion in warped spacetime

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The formation of cosmic strings

- GUT models
- String theory models → brane inflation → cosmic superstrings

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What a cosmic string network might look like

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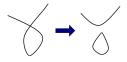
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Martins and Shellard

Detecting cosmic strings

- Lensing
- The CMB
- Gravitational radiation...



Closed loops give off a particularly distinctive gravitational wave signal. [Damour and Vilenkin 2001]

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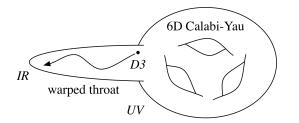
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Extra dimensions and brane inflation



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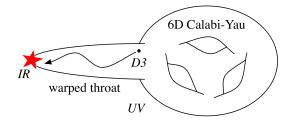
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Cosmic strings at the tip?



- Friction is required to stabilize strings at the tip.
- Model strings classically- relevant for the low energies of the current universe.

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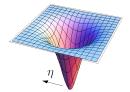
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Warped metric

We use the Klebanov Strassler metric as an explicit solution for ds_6^2 and $h(\eta)$ - the warped, deformed conifold.

$$ds^{2} = h^{-\frac{1}{2}}(\eta) \left(dt^{2} - a(t)^{2} d\mathbf{x}^{2} \right) - h^{\frac{1}{2}}(\eta) ds_{6}^{2}$$

 η is a radial coordinate in the internal space.



For simplicity, look at motion in only one of the angular directions of the conifold, $\phi.$

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Modelling cosmic string motion

We model the string classically using the Nambu action.

$${\cal S}=-\mu\int d au d\sigma \sqrt{-\gamma}$$

This gives equations of motion to solve: PDEs in τ and σ .

$$\frac{\partial}{\partial t} \left(\frac{\dot{x}^{a} x^{\prime 2}}{\sqrt{-\gamma}} \right) + \frac{\partial}{\partial \sigma} \left(\frac{x^{\prime a} \dot{x}^{2}}{\sqrt{-\gamma}} \right) + \frac{1}{\sqrt{-\gamma}} \Gamma^{a}_{bc} \left(x^{\prime 2} \dot{x}^{b} \dot{x}^{c} + \dot{x}^{2} x^{\prime b} x^{\prime c} \right) = 0$$

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Studying the equations of motion

1. Look at conserved quantities. We see how energy could be transferred between different terms.

$$E = \sqrt{\frac{\mathbf{x}'^2 + h\epsilon^{\frac{4}{3}} \left(\frac{\eta'^2}{6K^2} + B\phi'^2\right)}{h\left(1 - \dot{\mathbf{x}}^2 - h\epsilon^{\frac{4}{3}} \left(\frac{\dot{\eta}^2}{6K^2} + B\dot{\phi}^2\right)\right)}}$$

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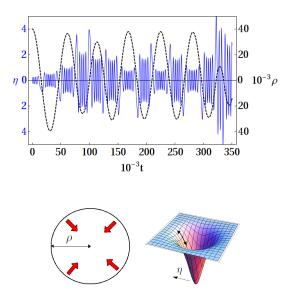
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2. Look at explicit trajectories to see that this transfer of energy really occurs.



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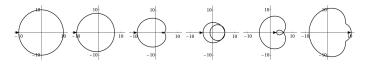
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3. Look at more complicated trajectories- numerical solutions to the full PDEs.



The same behaviour patterns show up, along with other interesting differences from the flat space case.

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Investigating the system further

We've considered other factors such as:

- Additional fields associated with the Klebanov-Strassler solution.
- Loss of energy from the strings via gravitational radiation.
- The effect of the expansion of space.

Our conclusion that extra dimensional motion is likely to be significant holds true in every case.

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