Next steps in the simulation/implementation of Einstein equations

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# **Overview:**

disconnected pieces of (some) reality...

- Book of stories, old readings, new readings...
- Preface.
- Review chapter. What we already learned
- Book #5. Tale of a disillusion –coordinate conditions—
- Book #6. Tale of search –striving for accuracy—
- Book #7. Kid's book –and they lived happily--

## Goal: obtain a robust implementation of Einstein equations

- Applicable to a large set of problems
- Independent of particular 'tweaks' *as much as possible*
- Able to yield accurate simulations (in theory & in practice)

## **Ingredients:**

- 1. Formulation with good properties (sym/strongly hyperbolic) (IVP ok at continuum) [Choquet-Bruhat,Fisher-Marsden,Friedrich,Reula,York...]
  - But which formulation is better?
- 2. Boundary treatment (*IVBP OK*) [*Friedrich-Nagy*, Calabrese-Pullin-Reula-Sarbach-Tiglio, Winicour-Szliagyi, Scheel-Lindblom-Teukolsky....]
  - Good for max. dissipative boundary conditions, radiative BC's require additional care!
- 3. Initial data
  - Far from ideal scenario... but more than enough for now to testing the 'GOAL'
- 4. Realize/translate (1+2) to the discrete arena
- 5. Adopt 'useful' coordinate conditions
- 6. Ensure simulation is well resolved and errors are on target

# Reminder...

- 'continuum stability'  $\rightarrow$  discrete stability
  - In FD, employ operators satisfying summation by parts so that

 $\int \dot{u}u \to \int u_{,x}u \to u^2 \mid_a^b \qquad \qquad \partial \to D \ s.t. \quad (u, Du) = -(Du, u) + u^2$ 

 Use these derivative operators + suitable time integrators (RK3, RK4..) + appropriate boundary conditions (Olsson method, SAT technique)

 $D_t u = A^i(t,x) D_i u + B^i(t,x) u + C(t,x)$ 

is stable, no tweaks or knobs required

 This might still not be enough, Leibnitz rule isn't satisfied at the discrete level... one can help out

$$u_{t} = 2cDu \quad "\neq "D(cu) + (Du) \quad c_{t,x} - (Dc) \quad u$$

 At the non-linear level, dissipation might be needed, dissipative operators have been constructed that do not spoil SBP

#### Combinations yield a strong starting point

[Calabrese,LL,Neilsen,Pullin,Reula,Sarbach,Tiglio] motivated by [Olsson,Tadmor,Kreiss,Carpenter,...]

## Coordinate conditions

## • A priori, requirements

#### **Fundamental**

- Artificial of coordinate singularities
  - Eg, 'geodesic' slicings t\_a = n\_a



[Palenzuela,LL,Bona]

- 'symmetry-seeking' (take advantage of possible approx symmetries)
  - Eg, adapted to the congruence defined by a Killing field
- If no symmetry, minimize the rate of change of appropriately defined geometrical quantities
  - Eg, minimal distortion of slices at t=const (York-Smarr)

#### <u>Additional</u>

- Suitable behavior near singularities
  - Eg. If no-excision singularity avoiding (Maximal slices), if excision be such that characteristics are all inflow at excision boundary  $v = (\pm \alpha + \beta) \rightarrow \alpha < \beta$
- Appropriate asymptotic behavior
- Covariantly defined if possible

#### God strongly suggested

- Combined with the main (evolution) eqns, whole system be at least strongly hyperbolic



• Minimal distortion/strain

Minimize action of the 'strain' of the hypersurface with respect to  $\beta$ 

$$\Sigma_{ij} = 2\nabla_{(i}\beta_{j)} - 2\alpha K_{ij}; \widetilde{\Sigma}_{ij} = \Sigma_{ij} - \lambda \Sigma_{ij}^{j}$$
$$S = \int \widetilde{\Sigma}_{ij} \widetilde{\Sigma}^{ij} \to \delta S = 0 \to \nabla^{j} \widetilde{\Sigma}_{ij} = 0$$

• Dynamical options:

$$\partial_t \alpha = -\alpha^2 Q$$
  
 $\partial_t \beta^i = -\alpha F^i$ 

– Need conditions for Q, F<sup>i</sup>; for instance, for Harmonic coords

$$\alpha Q = -\beta^{l} \partial_{l} \ln \alpha + \alpha K_{i}^{i}$$
$$\alpha F^{i} = -\beta^{l} \partial_{l} \beta^{i} - \alpha^{2}(g, \alpha, ...)$$

→ Symm hyperbolic, but not symmetry seeking without source functions.

## Lots of freedom...cutting to the chase..

• For lapse. Adopt...

$$\alpha Q = -\alpha \beta^l \partial_l \ln \alpha + f(\alpha) K_i^i$$

a=0,1 (along the *t* flow or *n* flow); f=0,1,2/ $\alpha$  (geodesic,time harmonic,1+log)

• For shift.... 'drive' the generalized minimal distortion condition

$$\partial_{tt}\beta^{i} = -\eta\alpha\partial_{t}\beta^{i} + g\alpha^{2}\nabla_{k}\widetilde{\Sigma}^{ki}$$

 $(\partial_t - a\beta^i \partial_i)\beta^i = -\alpha F^i \quad ; \quad (\partial_t - bL_\beta)F^i = -\eta F^i - \nabla_k (2g\alpha \widetilde{\Sigma}^{ki})$ 

- Props, FOR Z4 & Sarbach-Tiglio formulation and others...
  - a=1,b=1. Strongly hyperbolic; though doesn't freeze without sources
  - a=0,b=1. Weakly hyperbolic if  $\beta = \alpha$ . Freezes along the *n* flow?
  - a=0,b=0. Strongly hyperbolic, but needs boundary conditions at inner boundary, eigenvector mixes gauge & main variables. Freezes.

## Good example

- GR eqns:
  - Z4 [Bona].  $R_{ab} + 2\nabla_{(a}Z_{b)} = 0$
  - & Sarbach-Tiglio formalism (R<sub>ab</sub>=0 + constraints + judicial choice of pars)
- Gauge waves

$$ds^{2} = F(t,x)(-dt^{2} + dx^{2}) + dy^{2} + dz^{2}$$





- Not so good ones...
  - Schwarzschild with a perturbation: needed dissipation to deal with weakly hyperbolic pts.
    - During evolution hard to control that  $\alpha < \beta$ , when this is violated, instabilities arise.
    - Possibility... add suitable extra terms so that this condition is satisfied. But then... minimization might be broken
- Other alternatives?
  - Induce eqns from 'minimal' extended-Killing eqns, I.e. minimize

 $L = \xi_{(a;b)} \xi^{(a;b)} - 1/2 (\nabla_a \xi^a)^2$ 

- Killing, Homotetic Killing and Affine Killing vectors satisfy the eqn
- 'minimize' the Lagrangian  $\rightarrow$  dynamical conditions for  $\alpha,\beta$
- Resulting eqns are strongly hyperbolic except at  $\alpha = \beta$ !
- When there is a killing timelike vector, nice evolutions. When not...
   could be good if 'source' terms are added.
  - Eg, in Gowdy, drive to trK.

# Notes...

- How did we get in this mess?
  - Looked for a geometrical way of defining 'less spurious dynamics'. Dynamical (coupled to the main system) → elliptic equation
  - Adopted an evolution equation to obtain the solution, didn't want to hide 'under the rag' problems and stuck to the full equation.
    - → End up with options where freezing give nice conditions as long as BH's are not present.
    - → When BHs are present, either have to deal with weakly hyperbolic systems and/or need to give boundary conditions at the hole coupling main variables and gauge conditions
- How do we get out?
  - Stick to the elliptic problem and consider an elliptic-hyperbolic problem
    - Boundary conditions are known (e.g. Dain et.al.)
    - Elliptic gauge might be needed anyways (Sarbach's talk)
  - 'Abandon' symmetry seeking property
    - Strong/symmetric hyperbolicity ensured
    - Conditions seem to not be on the way of complicated problems (see Pretorius talk)

# The search: Resolution issue [LL,Liebling,Reula]

- Need to guarantee a simulation resolves the solution accurately
   Finer structure can develop without us being able to anticipate them
- Efficient use of (limited) resources

 $\rightarrow$  Need both adaptivity and higher accuracy operators.

Adaptive mesh refinement provides a way to do so.

Δx	$\Delta x/2$	

## Ground rules

- 1. Use what gives a stable *unigrid* evolution
- 2. Employ a 2-time-level scheme
- 3. Allow for higher order accuracy. Phase is especially delicate, LIGO needs it!!
  - Want error of ~ 1%? Need ~1%/(N\_time steps) per time-step.
  - Per-wavelength,  $2^{nd}$  order code,  $\varepsilon = 10^{-5} \rightarrow 816$  pts;  $4^{th}$  order  $\rightarrow 48$  pts;  $6^{th}$  order  $\rightarrow 19$  pts
  - Wavelength O(M) so...2<sup>nd</sup> order off after a few 10's of steps.. Need to go higher
  - Finer scale might develop, need to be adaptive
- What are the options?
  - Standard approach 'Berger-Oliger', provide boundary data by suitable interpolation at refinement boundaries to *all fields*
  - 'Penalty approach'. Provide boundary data *only* to incoming modes via penalty method
  - Tapered boundary approach.... 'fool-proof' way, integrate on a larger domain and drop points affected by whatever was done at the boundary.



## What is stable?

- No formal result available, intertwined evolutions so it isn't surprising.
- For lack of any better/smarter option, proceed as:
  - Employ SBP, boundary treatment + RK3/4 to guarantee stable unigrid part
  - Assess stability of parent  $\leftarrow \rightarrow$  child interaction on a model problem
  - Leave aside dissipation to begin with.
- Ingredients:
  - Child grid creation:  $J_{12}$  parent to child injection, *direct/interpolate to order (N+2)*
  - Update operator (RKs): U
  - Boundary data definition: B (involves interpolation from parent values at boundaries, if needed)
  - Parent fields redefined: J<sub>21</sub>, child to parent, *direct injection*.
  - Projector operator: P = 1 within child grid region of interest, 0 otherwise

$$u_p^{n+1} = Q \ u_p^n$$
$$Q \equiv (1-P)U + P(J_{21} \ B \ U \ B \ U \ J_{12})$$

If eigenvalue(Q)>1, unstable.

## Bottom line

- 2<sup>nd</sup> order, all obvious possibilities are stable. Tapered errors smaller, but not a big deal.
- 4<sup>th</sup> order
  - Standard: unstable for all obvious and not so obvious possibilities
    - Sideways derivatives, SBP derivatives, 2 ghost zones filled by interpolation (2<sup>nd</sup> or 3<sup>rd</sup> order accurate). ε
  - Penalty method: unstable for only 'allowed' possibilities (1 ghost zone filled by  $2^{nd}$  or  $3^{rd}$  order accurate interp)  $\epsilon/10$
  - Tapered boundary. Stable (obviously, and unfortunately?)  $\epsilon/100$ .
- Cost assessment:
  - For target error ε, 2<sup>nd</sup> standard order vs 4<sup>th</sup> order tapered, assuming N and 2N pts used respectively. C<sub>2</sub> 3<sup>3</sup>~C<sub>4</sub>. → 4<sup>th</sup> order 'wins' if ε<10<sup>-2</sup>
  - Dissipation helps... but errors are >> than those in tapered.





## Penalty





### Tapered

- Simple test;  $u_{t} = u_{x}$
- Parent grid + 2 levels of refinement (fixed in place).
- Look for long term behavior with dissipation added.





#### Linearized Einstein eqns over a gauge wave.

- First order in time/space symmetric hyperbolic eqn with time/space variable coefficients.
- Domain [-1,1], with base grid + 2 child grids [-.5/n,.5/n]; n=1,2.





- Non-linear wave eqn
- $\Box u = k u^p$
- Reduced to first order in time & space.
- Implemented RK3, 4<sup>th</sup> order derivative.
- Solution converges to 3<sup>rd</sup> order

• For p=7, singularity forms



# Notes (2)

- For higher order AMR, tapered approach is a (the?) viable way
  - Extra cost involved offset by gains in accuracy, reflections and CFL allowed
- Is this the only way? Perhaps (hopefully) not, but certainly will required more involved options than those considered
- Easy to implement... in fact... it should be there already in all codes
- For multiple grids (Reula's talk), should be no problem.





## Bubble spacetimes. A 'closed' story [Sarbach,LL]

Positive mass thm (Witten) requires existence of certain (asymptotically constant) spinors. In 5d Kaluza-Klein theory (asymptotically  $R^3xS^1$ ) these spinors are not guaranteed.

1. Are there negative mass configurations?

2.Is cosmic censorship valid?

#### Answer to 1. Yes, negative mass configurations found

- -Witten bubble (82): associated with instability of KK vacuum. More than 1 state with zero total energy.
- -Brill-Pfister (89): explicit solutions to 5D vacuum constraints with negative mass.
- -Brill-Horowitz (91): generalization to include 'gauge' fields.

#### •Qn: What's the space-time like?

-Corley-Jacobson (94). Analyze area of the bubble, conclusion: It starts out expanding [collapsing], if this trend continues, unlikely to form a singularity.

•Conjecture: It will keep expanding [collapsing] out (otherwise go through another moment of time symmetry).

•But..... This only from estimates at the initial hypersurface... what does really happen?... Need to solve the eqns...

-Numerical effort (2000). Conclusion: negative mass bubbles expand but not forever.... At some point a naked singularity appears!!! (or does it?)

# Revisiting the problem

• Consider:  $ds^2 = -dt^2 + U(r)dz^2 + U(r)^{-1}dr^2 + r^2d\Omega^2$ With U(r) greater or =0 (for  $r=r_+$ ) a smooth function ( $U \rightarrow I$  asymptotically) Bubble is at  $r_+$ . Electrovac case, consider  $\int dx^u - k(r^{-n} - r^{-n}) = k \in R$ ; n = 2, 3

$$A_{u}dx^{u} = k(r_{+}^{-n} - r^{-n}) \quad k \in R; n = 2,3....$$

Time symmetry (mom const =0); Hamiltonian constraint  $U(r) = 1 - m/r + b/r^2 + \tilde{k}^2/r^{2n}$ 

With *m*,*b* constants. In particular  $M_{ADM} = m/4$ ...but this <u>can be negative</u>

Initial acceleration of the bubble's area [extending Corley-Jacobson]

$$\ddot{A} \propto \left(1 - m - k[n-1][n-2]\right)$$

- n=2. If m<0, bubble expands; m>0 both cases possible
- *n*>2. For *k* large, arbitrary negative acceleration with negative mass...sounds promising!

## Numerical evolution

- Variables functions of *(t,r)* only (1D evolution)
- At outer boundary: constraint preserving boundary conditions [a-la Calabrese,LL,Tiglio 02]
- At bubble, regularity conditions used.
- Proved well posedness at continuum level, translated to the numerical arena thanks to SBP in a first order formulation.
  - Improved resolution at bubble with a non-uniform radial coordinate.



$$k=0$$

Case studied numerically previously, no naked singularity found, m < 0 expands even faster than m > 0

## More than we asked for...

What happens with a non-zero gauge field?
Choose *n*=2, and stick to cases where bubble starts out collapsing (positive mass)





Depending on field strength, the bubble either collapses ( $k < k^*$ ) to a black string or bounces back to expand ( $k > k^*$ ).



Changes behavior almost always without going through another moment of time symmetry

Last... it appears to approach a stationary solution... does it exist?

## Curvature invariant, sub/supra critical behavior



Observation... there must be a static solution at the threshold

• Put static anzats, solve resulting constraint and...

$$ds^{2} = -V(r)dt^{2} + V(r)/U(r)dr^{2} + U(r)/V(r)^{2}dz^{2} + r^{2}V(r)dz$$
$$A_{a}dx^{a} = \pm \frac{1}{2}\sqrt{3(r_{+}/r_{-}-1)}dz/V(r)$$

With  $V=(1-r_r/r)$ ;  $U=(1-r_r/r)$ . And the parameters are obtained from

 $P=4\pi r_{+}(1-r_{-}/r_{+})^{(3/2)}$  and  $M=r_{+}/4$ .

- New solution?... nah... obtained by 'just' analytically continuing that of a charged black string....[found in Horowitz-Maeda 03]
  - Analyzed perturbation off this solution:
    - Obtained single growing mode



 Work analytical continuation 'backwards'. Used to show a family of charged black strings becomes <u>more unstable</u> as charged is added (opposite to what was conjectured)

- What happened with the negative mass data that started contracting with arbitrary negative acceleration?
  - Bubble shrinks to arbitrarily small sizes, but 'bounces' back... cosmic censorship stood its ground



# And they lived happily ever after the end...