

CHEM3023: Spins, Atoms and Molecules

Lecture 6

Gaussian basis sets

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Learning outcomes

- Be able to convert expressions involving molecular orbitals to expressions involving basis functions and expansion coefficients and vice versa
- Be able to distinguish the characteristics and quality of different Gaussian basis sets
- Be able to select the basis set that provides the best compromise between required accuracy and computational efficiency for a given problem

The SCF procedure for Hartree-Fock calculations

- A Hartree-Fock calculation provides the optimum molecular orbitals for a Slater determinant wavefunction
- These orbitals are the eigenfunctions of the Hartree-Fock eigenvalue equation, which involves the Fock operator

$$\hat{f}[\{\psi_i\}_{i=1}^{N_{elec}/2}] \psi_j(\mathbf{r}) = \varepsilon_j \psi_j(\mathbf{r})$$

- The Fock operator is very unusual: it is constructed from its own eigenfunctions
- An iterative scheme called the “Self Consistent Field” (SCF) procedure has been developed to find its eigenfunctions and hence the optimum molecular orbitals of Hartree-Fock theory

Basis functions

- Even for small molecules, the SCF procedure involves a huge amount of computation
- However, a computer performs operations on numbers rather than mathematical functions
- We need to transform the Hartree-Fock equations in a format that involves operations with numbers rather than functions
- To achieve this we write the molecular orbitals as linear combinations of some **fixed** set of functions which we call “basis functions”
- Often the basis functions are atomic orbitals: in this case a molecular orbital is a Linear Combination of Atomic Orbitals (LCAO)

$$\text{molecular (spatial) orbital} \longrightarrow \psi_i(\mathbf{r}) = \sum_{\alpha=1}^{N_{BF}} G_{\alpha}(\mathbf{r}) C_{\alpha i} \longleftarrow \text{coefficient (a number, to be determined by the SCF calculation)}$$

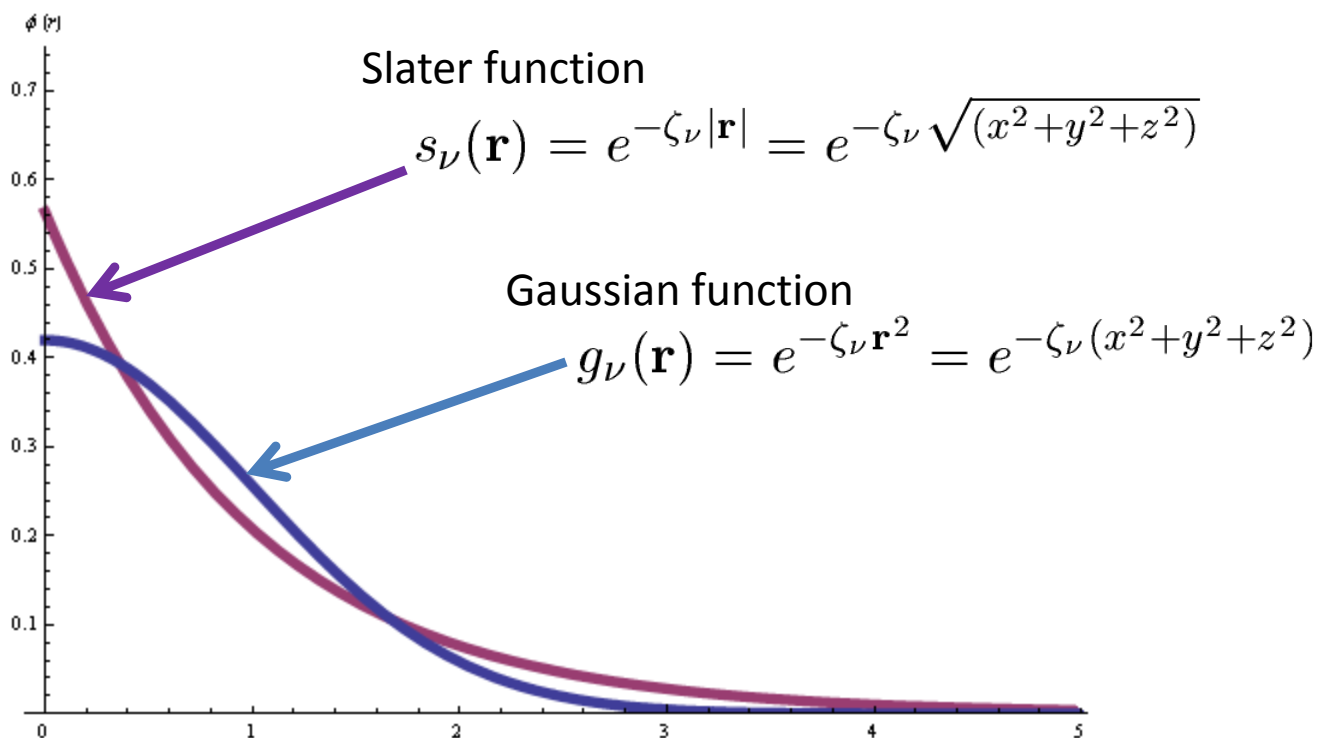
↑
basis function (fixed)

Gaussian basis functions

Functions that resemble hydrogen atomic orbitals (Slater functions) are a very good basis set because they have the correct shape:

- Near the nucleus (cusp)
- Far from the nucleus (decay like e^{-ar})

Even though Gaussian functions do not behave that well, they are preferred in practice because they allow for efficient computation of molecular integrals



Gaussian functions $g_\nu(\mathbf{r}) = x^k y^m z^n e^{-\zeta_\nu \mathbf{r}^2} = x^k y^m z^n e^{-\zeta_\nu (x^2 + y^2 + z^2)}$

Similar to atomic orbitals, we define **Cartesian primitive Gaussian** atomic functions by an angular momentum “quantum number” l . We have $l = k + m + n$ which leads to the following types of function:

s functions

$$l = 0 \quad \left\{ \begin{array}{l} e^{-\zeta \mathbf{r}^2} \end{array} \right.$$

p functions

$$l = 1 \quad \left\{ \begin{array}{l} x e^{-\zeta \mathbf{r}^2} \\ y e^{-\zeta \mathbf{r}^2} \\ z e^{-\zeta \mathbf{r}^2} \end{array} \right.$$

d functions

$$l = 2 \quad \left\{ \begin{array}{l} x^2 e^{-\zeta \mathbf{r}^2} \\ y^2 e^{-\zeta \mathbf{r}^2} \\ z^2 e^{-\zeta \mathbf{r}^2} \\ x y e^{-\zeta \mathbf{r}^2} \\ x z e^{-\zeta \mathbf{r}^2} \\ z y e^{-\zeta \mathbf{r}^2} \end{array} \right. \quad \text{etc.}$$

Contracted Gaussian functions

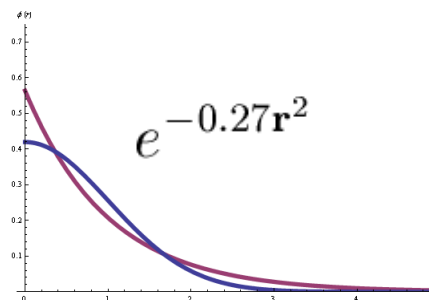
- In practice, fixed linear combinations of “primitive” Gaussian functions are used
- These are called “Contracted Gaussians” (CGs):

$$G_{\alpha}(\mathbf{r}) = \sum_{\nu=1}^{N_{\alpha}} c_{\nu} g_{\nu}(\mathbf{r})$$

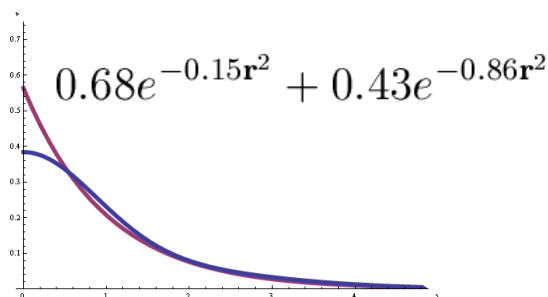
- The simplest kind of CGs are the **STO-nG** basis sets
- These basis sets approximate Slater-type orbitals (STOs) by n primitive Gaussians

E.g. STO-nG functions for the 1s orbital of a hydrogen atom

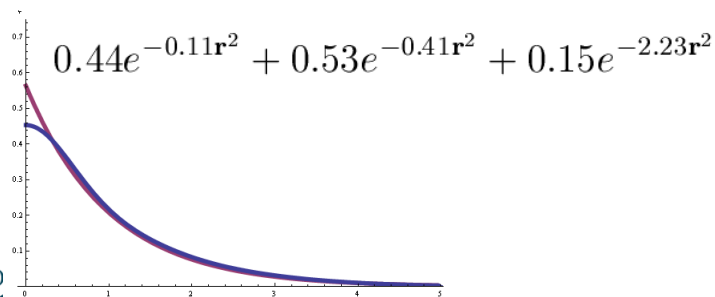
STO-1G



STO-2G



STO-3G



Gaussian basis sets

The STO-nG basis sets are rather unsatisfactory as they include only one contracted Gaussian (CG) per atomic orbital. Improved basis sets are obtained by including:

- More than one CG per atomic orbital, e.g.: **DZ** (“double zeta”), **TZ** (“triple zeta”), **QZ** (“quadruple zeta”)
- One CG per “core” atomic orbital and more than one for the valence atomic orbitals, e.g.: **SV**, **3-21G**, **4-31G**, **6-31G**, **6-311G**

Write down how many CGs some of the above basis sets will include for the following atoms: H, C, S

And for the following molecules: H₂O, CH₄

Polarisation and diffuse functions

Increasing the number of CGs per atomic orbital will never result in a good quality basis set. Other types of CGs need to be included, such as:

- CGs of angular momentum higher than in the valence orbitals of each atom. These “**polarisation functions**” enhance the “flexibility” of atoms to form chemical bonds in any direction and hence improve calculated molecular structures. Examples: **3-21G***, **6-31G***, **6-31G****, **DVP**, **TZP**, **cc-pVDZ**, **cc-pVTZ**
- CGs which extend further from the nucleus than the atomic orbitals. Such “**diffuse functions**” improve the predicted properties of species with extended electronic densities such as anions or molecules forming hydrogen bonds. Examples: **4-31+G**, **6-31+G**
- Basis sets are considered “balanced” when they include both polarisation and diffuse functions. Examples: **6-31+G***, **6-311++G****, **aug-cc-pVDZ**

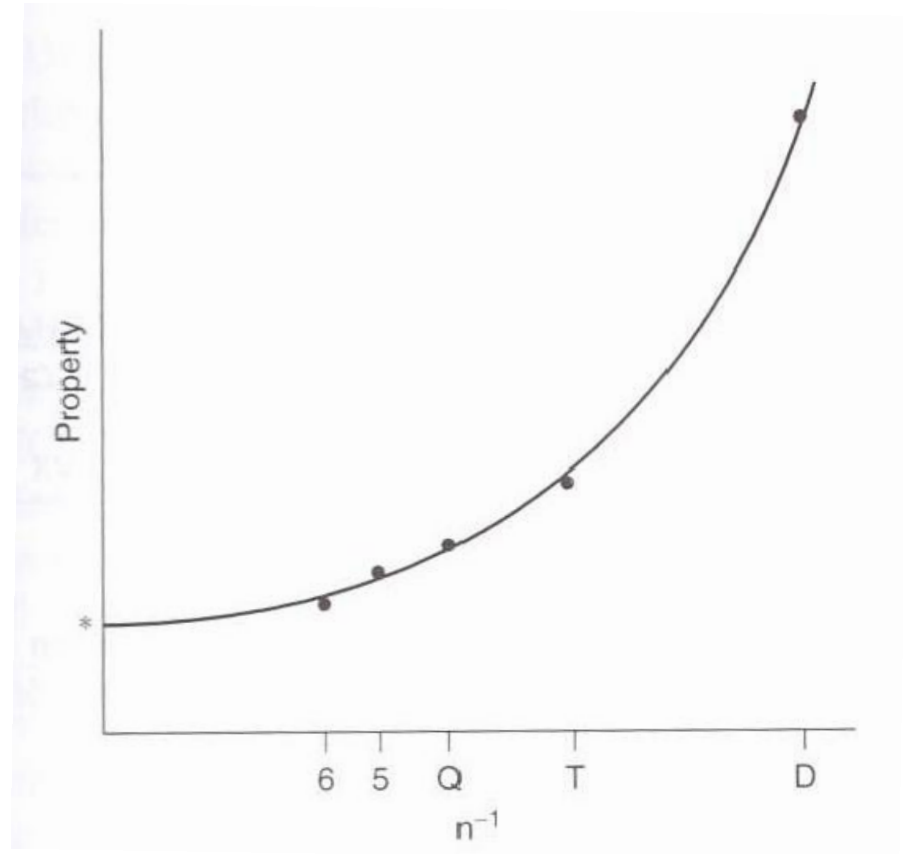
Write down how many CGs some of the above basis sets will include for the following atoms: H, F, S and molecules: H₂O, CH₄

The complete basis set limit

- Basis sets are an approximation introduced in order to solve equations for molecular orbitals on a computer
- The molecular orbitals obtained are solutions of the Hartree-Fock equations only within the “function space” of the basis set used (so solutions within the STO-3G set of functions, or the 6-31G set, etc.)
- Improving the quality of the basis set results in increasing the number of CGs
- A **complete** basis set can represent exactly **any molecular orbital**
- Unfortunately, complete basis sets tend to have an infinite number of functions and are therefore not practical for calculations

Extrapolation to the complete basis set limit

- We can estimate the complete basis set result by systematically increasing the number of basis functions and extrapolating to an infinite-size basis set
- In Hartree-Fock calculations, the complete basis set solutions are called the “**Hartree-Fock limit**”
- The cc-pVDZ, cc-pVTZ, cc-pVQZ, etc, basis sets are an example of a systematic series of basis sets that can be extrapolated to the complete basis set limit



Examples of H-F calculations with Gaussian basis sets

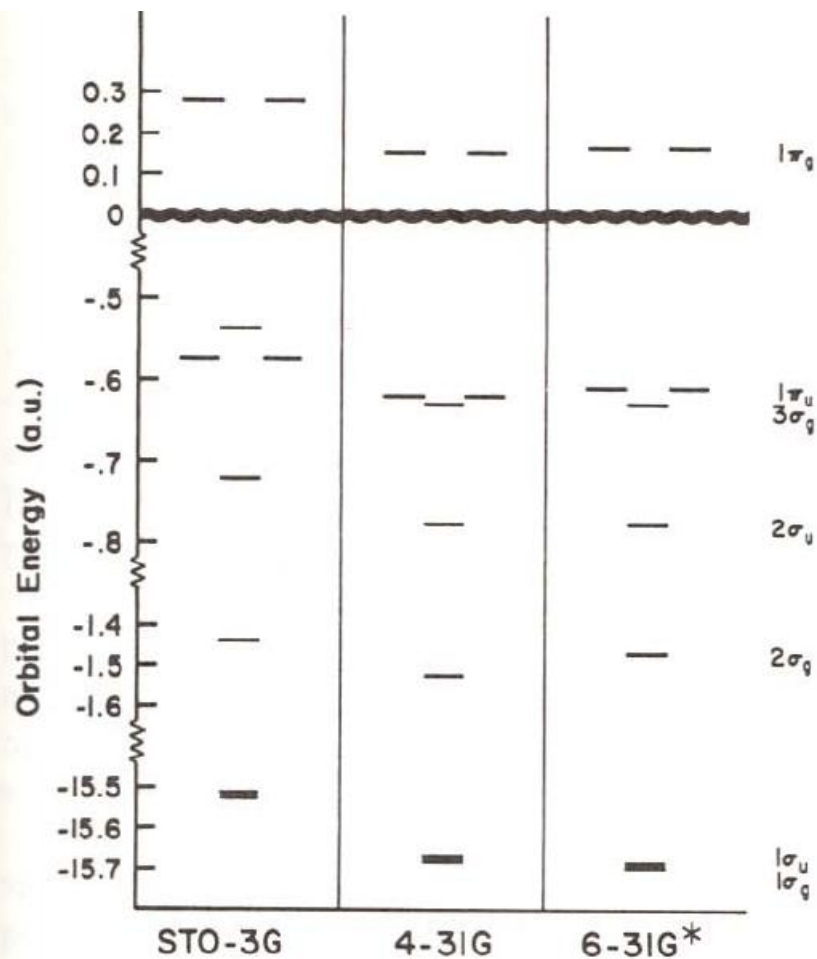
SCF total energies (a.u.) of N₂ and CO with the standard basis sets

Basis set	N ₂	CO
STO-3G	-107.496	-111.225
4-31G	-108.754	-112.552
6-31G*	-108.942	-112.737
HF-limit ^a	-108.997	-112.791

SCF equilibrium bond angles for NH₃ and H₂O

Basis set	NH ₃	H ₂ O
STO-3G	104.2	100.0
4-31G	115.8	111.2
6-31G*	107.5	105.5
6-31G**	107.6	106.0
Near-HF-limit	107.2 ^c	106.1 ^b
Experiment	106.7	104.5

Orbital energies of N₂ with Gaussian basis sets



Example: available basis sets in a quantum chemistry program

3-21g	aug-cc-pvqz	cc-pvtzfi_sf_fw	lanl2dzdp_ecp_polarization
3-21++g	aug-cc-pvqz_diffuse	cc-pvtzfi_sf_lc	lanl2dz_ecp
3-21gs	aug-cc-pvt+dz	cc-pvtzfi_sf_sc	mclean_chandler_vtz
3-21++gs	aug-cc-pvt+dz_diffuse	cc-pvtz-fit2-1	midi!
3-21gsp	aug-cc-pvtz	cc-pvtzpt_sf_fw	midi_huzinaga
3-21gs_polarization	aug-cc-pvtz_diffuse	cc-pvtzpt_sf_lc	mini_huzinaga
4-22gsp	bauschlicher_ano	cc-pvtzpt_sf_sc	mini_scaled
4-31g	binning-curtiss_ld_polarization	cc-pvtzsseg-opt	nasa_ames_ano
6-311g	binning-curtiss_df_polarization	chipman_dzp_+_diffuse	nasa_ames_cc-pcv5z
6-311++g2d_2p	binning_curtiss_sv	core_val._functions_cc-pcv5z	nasa_ames_cc-pcvqz
6-311g2df_2pd	binning_curtiss_svp	core_val._functions_cc-pcv6z	nasa_ames_cc-pcvtz
6-311++g3df_3pd	binning_curtiss_vtz	core_val._functions_cc-pcvdz	nasa_ames_cc-pv5z
6-311gs	binning_curtiss_vtzp	core_val._functions_cc-pcvqz	nasa_ames_cc-pvqz
6-311+gs	blaudeau_polarization	core_val._functions_cc-pcvtz	nasa_ames_cc-pvtz
6-311gs_polarization	cc-pcv5z	crenbl_ecp	partridge_uncontr._1
6-311gss	cc-pcv6z	crenbs_ecp	partridge_uncontr._2
6-311++gss	cc-pcvdz	d-aug-cc-pv5z	partridge_uncontr._3
6-311gss_polarization	cc-pcvqz	d-aug-cc-pv5z_diffuse	people_2d_2p_polarization
6-31g	cc-pcvtz	d-aug-cc-pv6z	people_2df_2pd_polarization
6-31++g	cc-pv5+dz	d-aug-cc-pv6z_diffuse	people_3df_3pd_polarization
6-31g3df_3pd	cc-pv5z	d-aug-cc-pvdz	people-style_diffuse
6-31g-blaudeau	cc-pv5z_dk	d-aug-cc-pvdz_diffuse	pv6z
6-31gs	cc-pv5zfi_sf_fw	d-aug-cc-pvqz	qmmm_zhang_3-21g_ecp
6-31+gs	cc-pv5zfi_sf_lc	d-aug-cc-pvqz_diffuse	qmmm_zhang_6-31gs_ecp
6-31++gs	cc-pv5zfi_sf_sc	d-aug-cc-pvtz	sadlej_pvtz
6-31gs-blaudeau	cc-pv5zpt_sf_fw	d-aug-cc-pvtz_diffuse	sbkjc_vdz_ecp
6-31gs_polarization	cc-pv5zpt_sf_lc	demon_coulomb_fitting	sdb-aug-cc-pvqz
6-31gss	cc-pv5zpt_sf_sc	dgauss_a1_dft_coulomb_fitting	sdb-aug-cc-pvqz_diffuse
6-31++gss	cc-pv6+dz	dgauss_a1_dft_exchange_fitting	sdb-aug-cc-pvtz
6-31gss_polarization	cc-pv6z	dgauss_a2_dft_coulomb_fitting	sdb-aug-cc-pvtz_diffuse
ahlrichs_coulomb_fitting	cc-pvd+dz	dgauss_a2_dft_exchange_fitting	sdb-cc-pvqz
ahlrichs_polarization	cc-pvdz	dhms_polarization	sdb-cc-pvtz
ahlrichs_pvdz	cc-pvdz_dk	dunning-hay_diffuse	sto-2g
ahlrichs_tzv	cc-pvdzfi_sf_fw	dunning-hay_double_rydberg	sto-3g
ahlrichs_vdz	cc-pvdzfi_sf_lc	dunning-hay_rydberg	sto-3gs
ahlrichs_vtz	cc-pvdzfi_sf_sc	dz_+_double_rydberg_dunning-hay	sto-3gs_polarization
aug-cc-pcv5z	cc-pvdz-fit2-1	dz_dunning	sto-6g
aug-cc-pcvdz	cc-pvdzpt_sf_fw	dzp_+_diffuse_dunning	stuttgart_rlc_ecp
aug-cc-pcvqz	cc-pvdzpt_sf_lc	dzp_dunning	stuttgart_rsc_1997_ecp
aug-cc-pcvtz	cc-pvdzpt_sf_sc	dzp_+_rydberg_dunning	stuttgart_rsc_ano_ecp
aug-cc-pv5+dz	cc-pvdsseg-opt	dz_+_rydberg_dunning	stuttgart_rsc_segmented_ecp
aug-cc-pv5+dz_diffuse	cc-pvq+dz	dzvp2_dft_orbital	sv_+_double_rydberg_dunning-hay
aug-cc-pv5z	cc-pvqz	dzvp_dft_orbital	sv_dunning-hay
aug-cc-pv5z_diffuse	cc-pvqz_dk	feller_misc._cvdz	svp_+_diffuse_dunning-hay
aug-cc-pv6+dz	cc-pvqzfi_sf_fw	feller_misc._cvqz	svp_+_diffuse_+_rydberg
aug-cc-pv6+dz_diffuse	cc-pvqzfi_sf_lc	feller_misc._cvtz	svp_dunning-hay
aug-cc-pv6z	cc-pvqzfi_sf_sc	gamess_pvtz	svp_+_rydberg_dunning-hay
aug-cc-pv6z_diffuse	cc-pvqzpt_sf_fw	gamess_vtz	sv_+_rydberg_dunning-hay
aug-cc-pvd+dz	cc-pvqzpt_sf_lc	glendering_polarization	tz_dunning
aug-cc-pvd+dz_diffuse	cc-pvqzpt_sf_sc	hay-wadt_mb_n+1_ecp	tzvp_dft_orbital
aug-cc-pvdz	cc-pvqsseg-opt	hay-wadt_vdz_n+1_ecp	wachters+f
aug-cc-pvdz_diffuse	cc-pvt+dz	hondo7_polarization	wtbs
aug-cc-pvq+dz	cc-pvtz	huzinaga_polarization	
aug-cc-pvq+dz_diffuse	cc-pvtz_dk	lanl2dzdp_ecp	

Basis sets on the Web

- All kinds of basis sets have been developed over the years
- Most are available for download from websites, such as
 - <https://bse.pnl.gov/bse/portal>
 - <http://www.emsl.pnl.gov/forms/basisform.html>

The screenshot shows the EMSL Basis Set Exchange website in a Mozilla Firefox browser. The page title is "EMSL Basis Set Exchange - Mozilla Firefox". The address bar shows "https://bse.pnl.gov/bse/portal". The website header includes the EMSL logo (Environmental Molecular Sciences Laboratory) and the text "BASIS SET EXCHANGE". There are login and contributor buttons. A navigation bar shows "Basis Set Exchange: v 1.1" and links for "Feedback", "About", "ReleaseNotes", and "Help".

The main content area displays a search for "6-31G*" basis sets. A list on the left shows search results, with "6-31G*" selected. The main area shows a periodic table with elements highlighted in blue, indicating the search results. The text "Total: 139 published basis sets that contain C" is displayed. Below the periodic table, there are search options: "Format: NWChem" and "Optimized General Contractions" (checked). A "Get Basis Set" button is present.

Below the search results, there is a section titled "6-31G*" Basis Set Information. It includes the following details:

Abstract:	VDZP Valence Double Zeta + Polarization (Li-Ar)	Contributor:	Dr. David Feller
Primary Developer:	N/A	Curation Status:	published
Last Modified:	Mon, 15 Jan 2007 23:47:00 GMT		

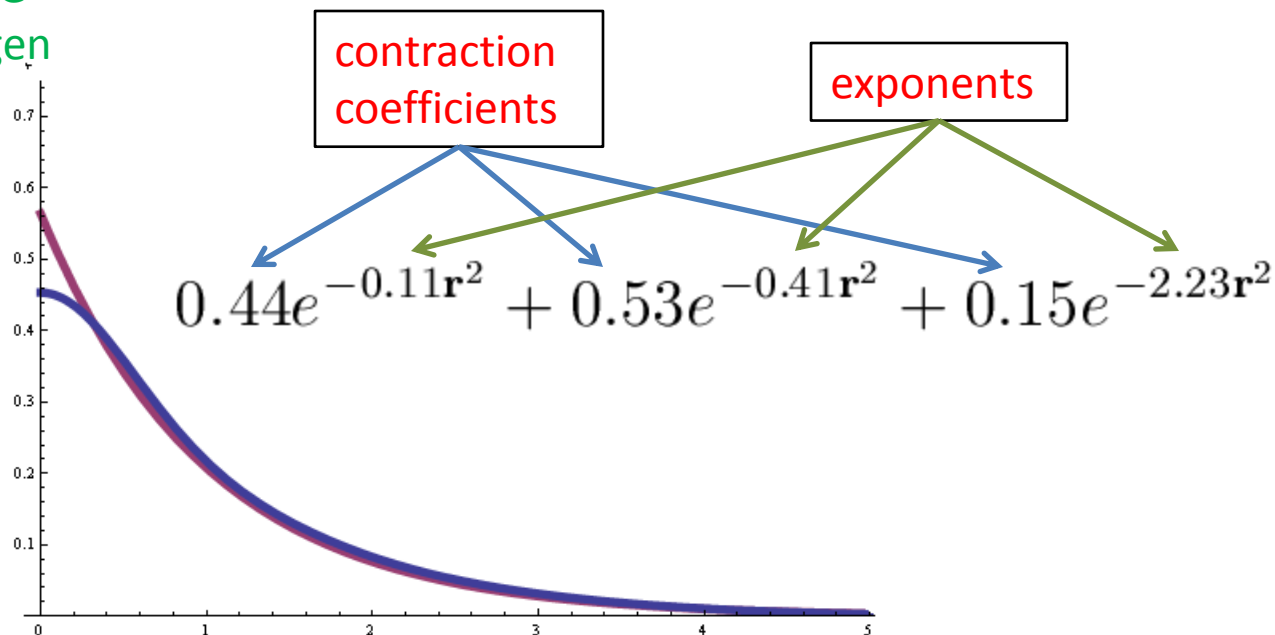
Exponents and contraction coefficients

Basis sets are essential input data for calculations. Gaussian bases are represented by two kinds of numbers:

- 1) Exponents
- 2) Contraction coefficients

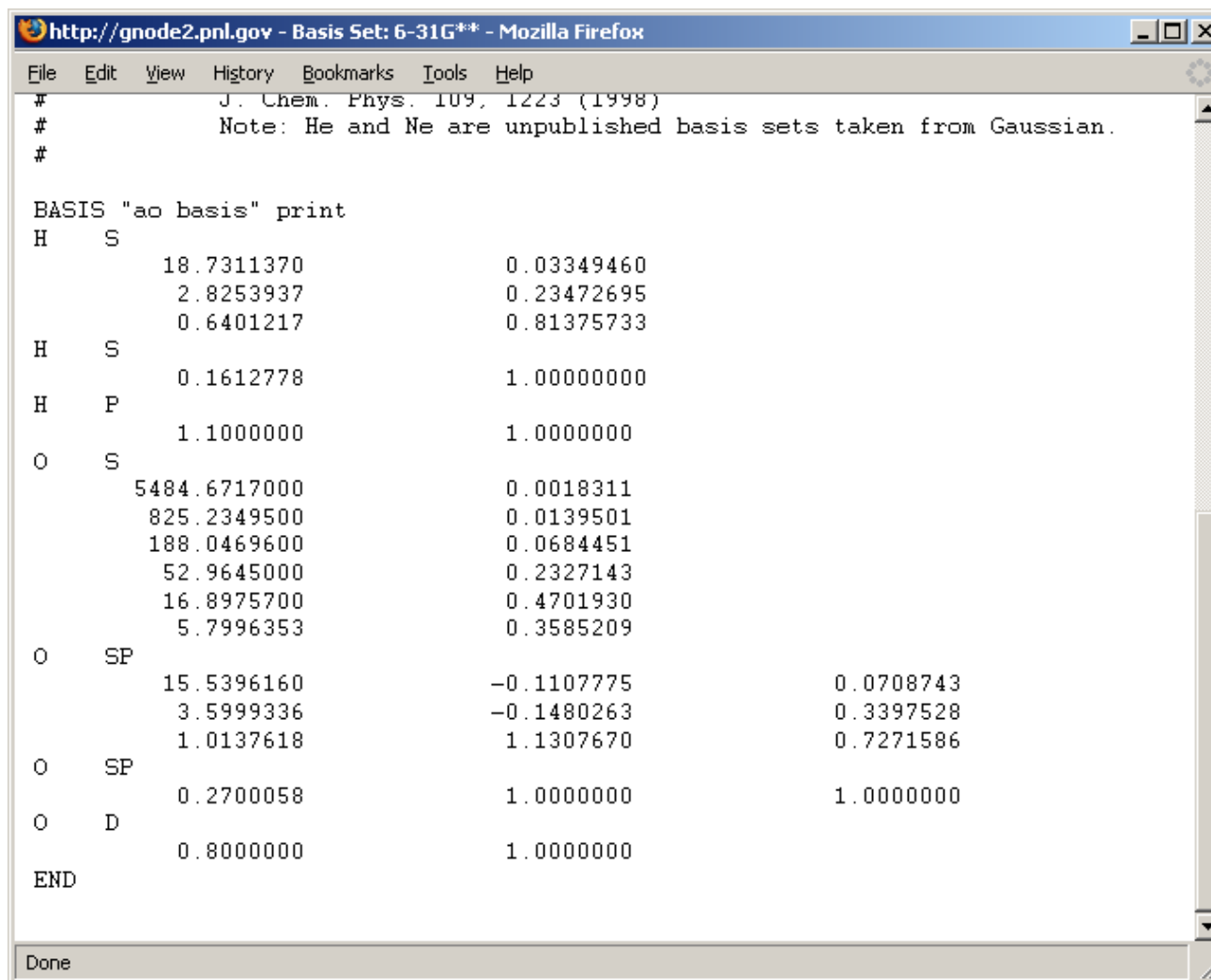
$$G_{\alpha}(\mathbf{r}) = \sum_{\nu=1}^{N_{\alpha}} c_{\nu} e^{-\zeta_{\nu} \mathbf{r}}$$

Example: STO-3G
basis for hydrogen



Downloading basis sets

Example: Data needed for calculation on water with 6-31G** basis



```
http://gnode2.pnl.gov - Basis Set: 6-31G** - Mozilla Firefox
File Edit View History Bookmarks Tools Help
# J. Chem. Phys. 109, 1223 (1998)
# Note: He and Ne are unpublished basis sets taken from Gaussian.
#
BASIS "ao basis" print
H S
    18.7311370      0.03349460
    2.8253937      0.23472695
    0.6401217      0.81375733
H S
    0.1612778      1.00000000
H P
    1.1000000      1.0000000
O S
    5484.6717000    0.0018311
    825.2349500     0.0139501
    188.0469600     0.0684451
    52.9645000      0.2327143
    16.8975700      0.4701930
    5.7996353       0.3585209
O SP
    15.5396160      -0.1107775      0.0708743
    3.5999336       -0.1480263      0.3397528
    1.0137618        1.1307670       0.7271586
O SP
    0.2700058        1.0000000        1.0000000
O D
    0.8000000        1.0000000
END
Done
```

Summary / Reading assignment

- LCAO approach (Cramer, pages 111-113)
- Gaussian basis sets (Cramer, pages 166-178)