

Non-abelian dyons in anti-de Sitter space

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University
Of
Sheffield.

Outline

- 1 A brief history of Einstein-Yang-Mills
- 2 $\mathfrak{su}(N)$ EYM with $\Lambda < 0$
- 3 Dyonic solutions
- 4 Conclusions and outlook

A brief history of Einstein-Yang-Mills

Asymptotically flat EYM studied for over 20 years

- Purely magnetic $\mathfrak{su}(2)$ solitons and black holes found numerically 1989-90
- Have no magnetic charge

Non-abelian baldness of asymptotically flat $\mathfrak{su}(2)$ EYM

If $\Lambda = 0$, the only solution of the $\mathfrak{su}(2)$ EYM equations with a non-zero (electric or magnetic) charge is Abelian Reissner-Nordström

Rules out dyonic $\mathfrak{su}(2)$ asymptotically flat solutions

[Ershov and Gal'tsov, PLA **138** 160 (1989), **150** 159 (1990)]

- Ershov/Gal'tsov result no longer holds for larger gauge group in asymptotically flat space
- What about asymptotically anti-de Sitter space?

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The model for $\mathfrak{su}(N)$ EYM

Einstein-Yang-Mills theory with $\mathfrak{su}(N)$ gauge group

$$S = \frac{1}{2} \int d^4x \sqrt{-g} [R - 2\Lambda - \text{Tr} F_{\mu\nu} F^{\mu\nu}]$$

Field equations

$$\begin{aligned} R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} + \Lambda g_{\mu\nu} &= T_{\mu\nu} \\ D_\mu F_\nu^\mu &= \nabla_\mu F_\nu^\mu + [A_\mu, F_\nu^\mu] = 0 \end{aligned}$$

Stress-energy tensor

$$T_{\mu\nu} = \text{Tr} F_{\mu\lambda} F_\nu^\lambda - \frac{1}{4} g_{\mu\nu} \text{Tr} F_{\lambda\sigma} F^{\lambda\sigma}$$

Static, spherically symmetric, configurations

Metric

$$ds^2 = -\mu(r)\sigma(r)^2 dt^2 + [\mu(r)]^{-1} dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

$$\mu(r) = 1 - \frac{2m(r)}{r} - \frac{\Lambda r^2}{3}$$

$\mathfrak{su}(N)$ gauge potential [Kunzle *Class. Quant. Grav.* **8** 2283 (1991)]

Static, dyonic, gauge potential

$$A_\mu dx^\mu = \mathcal{A} dt + \frac{1}{2} (C - C^H) d\theta - \frac{i}{2} \left[(C + C^H) \sin \theta + D \cos \theta \right] d\phi$$

$N - 1$ electric gauge field functions $h_j(r)$ in matrix \mathcal{A}

$N - 1$ magnetic gauge field functions $\omega_j(r)$ in matrix C

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Field equations

Yang-Mills equations

$$\begin{aligned}
 h_k'' &= h_k' \left(\frac{\sigma'}{\sigma} - \frac{2}{r} \right) + \sqrt{\frac{2(k+1)}{k}} \frac{\omega_k^2}{\mu r^2} \left(\sqrt{\frac{k+1}{2k}} h_k - \sqrt{\frac{k-1}{2k}} h_{k-1} \right) \\
 &\quad + \sqrt{\frac{2k}{k+1}} \frac{\omega_{k+1}^2}{\mu r^2} \left(\sqrt{\frac{k}{2(k+1)}} h_k - \sqrt{\frac{k+2}{2(k+1)}} h_{k+1} \right) \\
 0 &= \omega_k'' + \omega_k' \left(\frac{\sigma'}{\sigma} + \frac{\mu'}{\mu} \right) + \frac{\omega_k}{\sigma^2 \mu^2} \left(\sqrt{\frac{k+1}{2k}} h_k - \sqrt{\frac{k-1}{2k}} h_{k-1} \right)^2 \\
 &\quad + \frac{\omega_k}{\mu r^2} \left(1 - \omega_k^2 + \frac{1}{2} [\omega_{k-1}^2 + \omega_{k+1}^2] \right)
 \end{aligned}$$

Einstein equations

Give $\mu'(r)$ and $\sigma'(r)$ in terms of ω_k , h_k and their derivatives

Solving the field equations

Numerically solve the field equations for $\mathfrak{su}(2)$ and $\mathfrak{su}(3)$, with different values of $\Lambda < 0$

Soliton solutions

- Regular at the origin $r = 0$
- Solutions parameterized by $\omega_k''(0)$ and $h_k'(0)$

Black hole solutions

- Regular event horizon at $r = r_h = 1$
- Solutions parameterized by $\omega_k(r_h)$ and $h_k'(r_h)$

Electric functions h_k are monotonically increasing

Colour-code solution space by number of zeros of magnetic gauge functions ω_k

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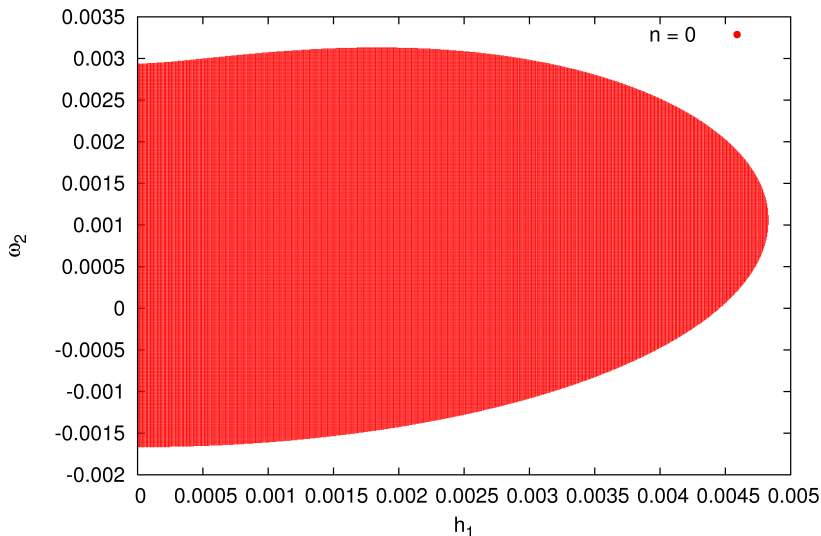
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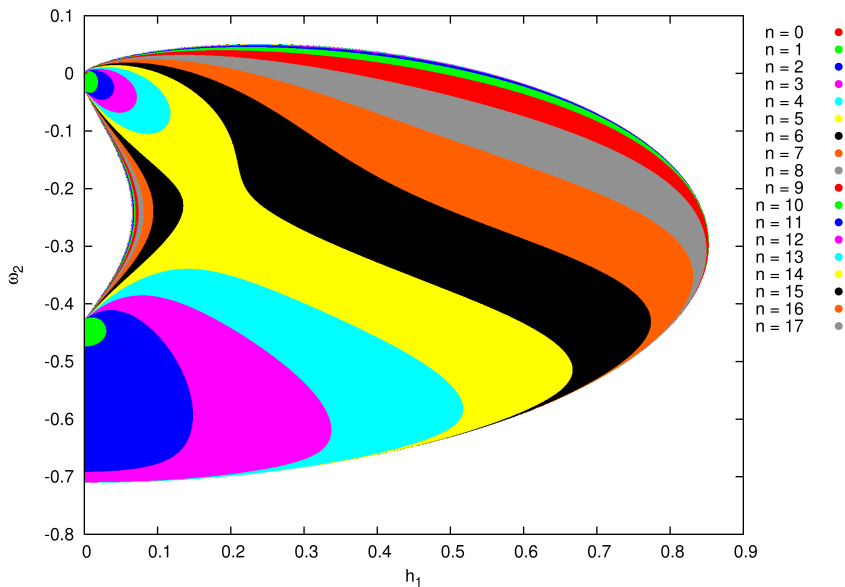
Colour-code solution space by number of zeros of magnetic gauge functions ω_k

$\mathfrak{su}(2)$ solitons, $\Lambda = -0.01$

[Bjoraker and Hosotani, PRL **84** 1853 (2000), PRD **62** 043513 (2000)]

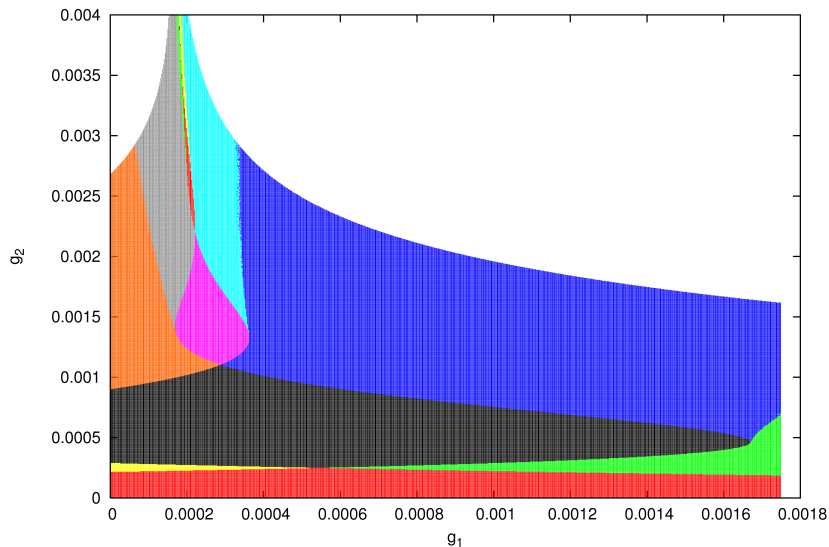


$\mathfrak{su}(2)$ solitons, $\Lambda = -0.01$ [Shepherd and EW]

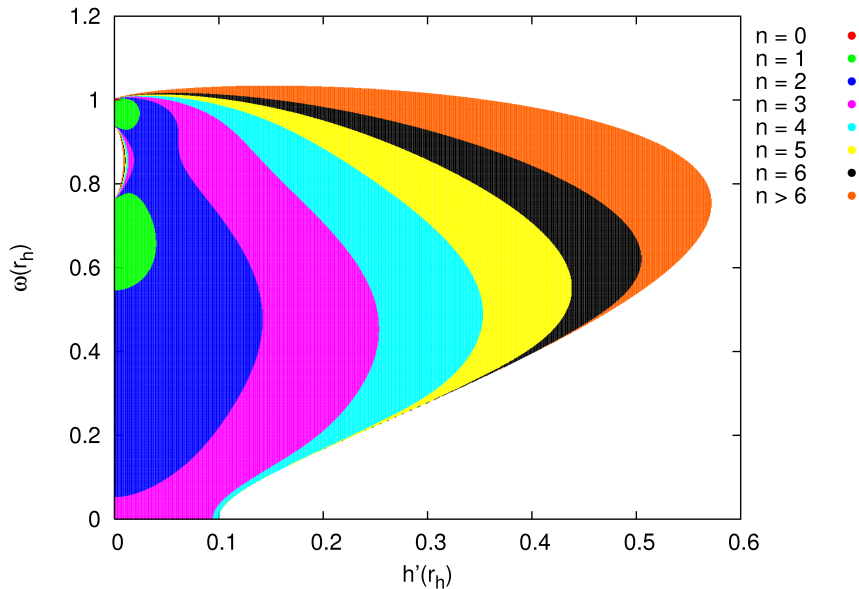


$\mathfrak{su}(3)$ solitons - part of the solution space for $\Lambda = -0.01$

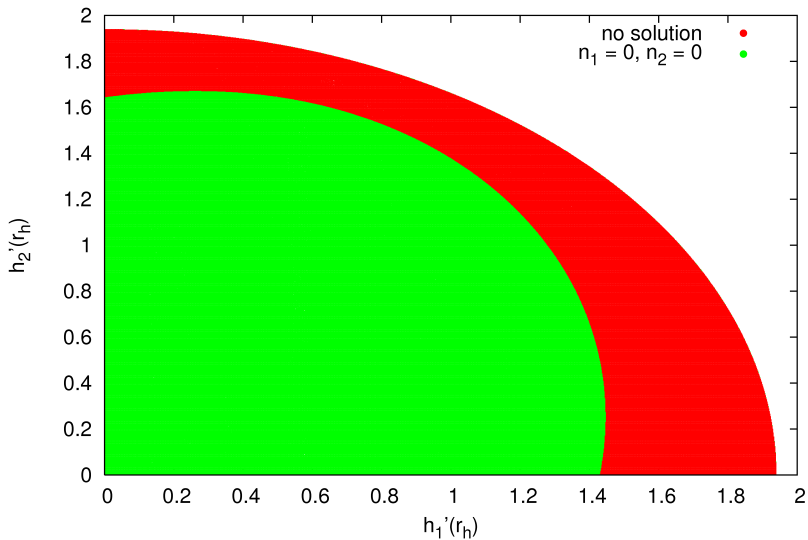
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$su(2)$ black holes, $\Lambda = -0.01$, $r_h = 1$ [Shepherd and EW]



$\mathfrak{su}(3)$ black holes $\Lambda = -3$, $r_h = 1$, $\omega_1(r_h) = 1.3$,
 $\omega_2(r_h) = 1.2$ [Shepherd and EW]



Conclusions and outlook

Dyons in $\mathfrak{su}(N)$ EYM in adS

Black hole and soliton solutions with both electric and magnetic gauge fields:

- $\mathfrak{su}(2)$ dyonic solutions found previously by Bjoraker and Hosotani
- Examination of larger part of solution space shows a very rich menagerie of solutions
- New $\mathfrak{su}(3)$ dyonic solutions
- Solutions in which magnetic functions ω_k have no zeros for larger $|\Lambda|$

Open questions

For sufficiently large $|\Lambda|$

- Prove analytically the existence of dyonic solutions for all N
- Stability when ω_k have no zeros?

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