
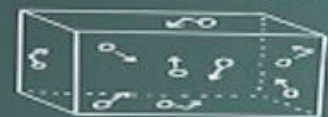


# PHYS1013


## Energy and Matter

$U_i (n_i, P_i, V_i, \dots)$   $U_f (n_f, P_f, V_f, \dots)$   $W = -nRT \int_{V_i}^{V_f} \frac{dV}{V} = -nRT \ln\left(\frac{V_f}{V_i}\right)$   $H = U + pV$   $T(K) = T(^{\circ}C) + 273.15$   
 $dH = dU + d(pV)$   $dH = dU + pdV + Vdp$   $C_p = (\Delta H / \Delta T)_p$   $\Delta U = Q - W$   $\Delta S = nRT \ln\left(\frac{V_f}{V_i}\right)$   
 $dU = dq + dw$   $dH = dq - pdV + Vdp$   $C_p = \left(\frac{\partial H}{\partial T}\right)_p$   $W = P\Delta U$   $W = \int_{V_1}^{V_2} P dV$   
 $H = U + PV$   $dH = C_p dT$   $\Delta H = q_p = C_p \Delta T$   $C_v = (\Delta U / \Delta T)_v$   $ds \geq \frac{dq}{T}$   
 $dw = -pdv$   $\Delta S = \frac{\Delta_{\text{trans}} H}{T}$   $dS = \frac{dq_{\text{rev}}}{T}$   $\Delta S = \int_1^f \frac{dq_{\text{rev}}}{T}$   
 $C_v = \left(\frac{\partial U}{\partial T}\right)_v$   $ds \geq \frac{dq}{T}$






$\Delta U = m(u_2 - u_1) \Delta KE$   
 $= \frac{1}{2} m (v_2^2 - v_1^2) \Delta PE$   
 $= mg(z_2 - z_1)$

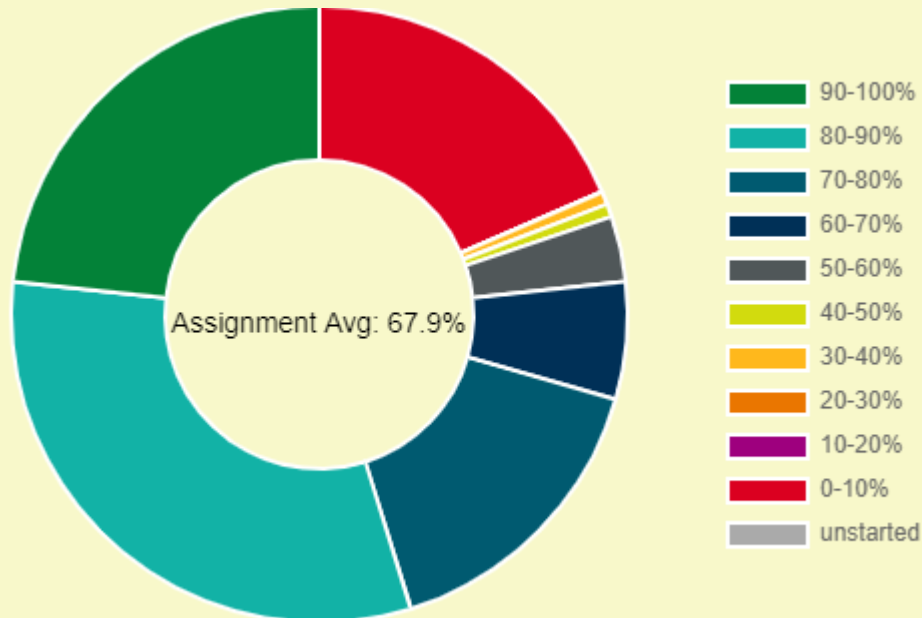
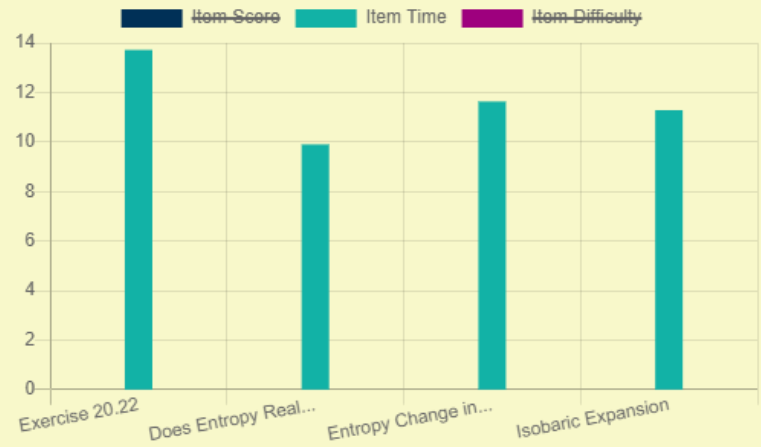
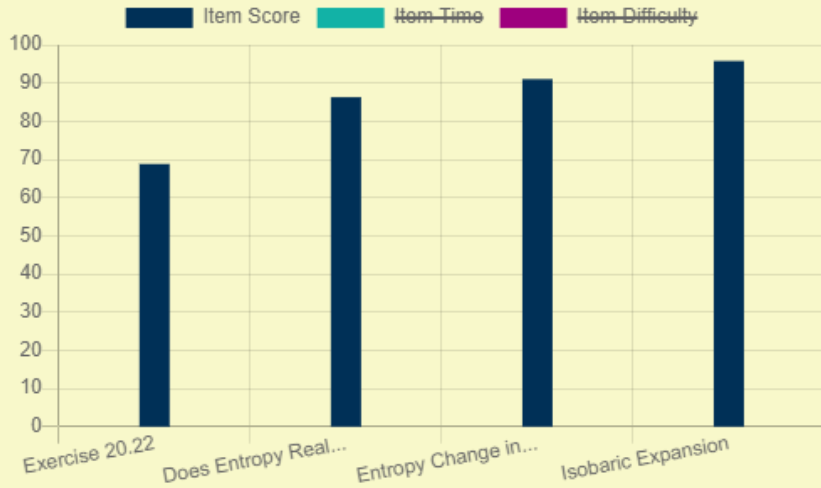


$W_b = \frac{P_1 V_1 - P_2 V_2}{1 - \gamma}$   $\eta_{th} = \frac{W_{net}}{Q_{in}} = 1 - \frac{Q_{out}}{Q_{in}}$   
 $dH = dq + vdp$   $\Delta H = \Delta U + V\Delta p$   $Q = \Delta U + P\Delta V$   
 $dH = (dq)_p$   $\Delta H = q_p$   $T_R = \frac{T}{T_c}$   $dU = C_v dT$   
 $dU = (dq)_v$   $\Delta U = q_v$   $\Delta U = q_v = C_v \Delta T$   $\Delta U = U_f - U_i = q(\text{heat}) + w(\text{work})$   
 $P_R = \frac{P}{P_{cr}}$   $W_b = P_1 V_1 \ln \frac{V_2}{V_1}$   $\frac{P_1}{P_2} = RT_1 \ln \frac{P_1}{P_2}$   $\gamma_k = \frac{\gamma P_{cr}}{RT_{cr}}$



MP Week 7 - average score 67.9%  
(last year 68.3%)

Average time  
47 min  
(last year 42 min)



# Work on our 2023 Undergraduate Open Days!

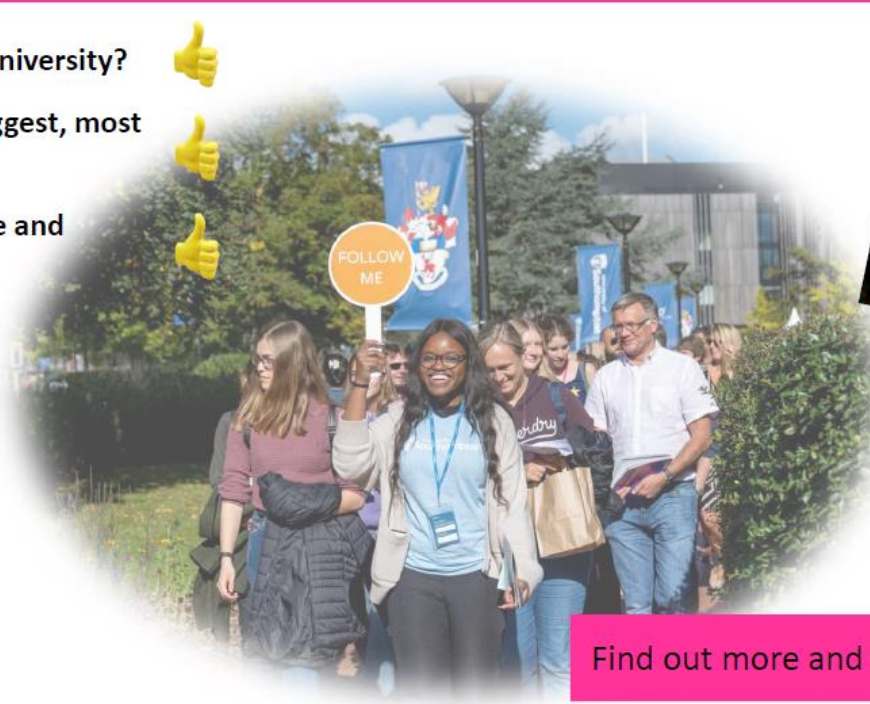
## Friendly and Enthusiastic Students Wanted

- ✓ Are you passionate about your University?
- ✓ Do you want to be part of the biggest, most exciting University event?
- ✓ Do you want to take on a fun role and enhance your employability



Come and see us at our Student Ambassador Pop Up in front of Hartley Library to find out more!

3<sup>rd</sup>, 10<sup>th</sup>, 18<sup>th</sup>, 23<sup>rd</sup> 31<sup>st</sup> May  
10:00 – 14:00



Pay rate:

£11.41 per hour (under 23)  
£11.68 per hour (23+)  
*(rates inclusive of holiday pay)*

Find out more and apply!



The exam is 14:30 on the 25<sup>th</sup> May.

It will consist of

Section A - all questions to be answered

Section B - answer 2 of 3 questions

**Answers to Section A and Section B must be in separate answer books**

Answer **all** questions in **Section A** and **only two** questions in **Section B**.

**Section A** carries  $\frac{1}{3}$  of the total marks for the exam paper and you should aim to spend about 40 mins on it.

**Section B** carries  $\frac{2}{3}$  of the total marks for the exam paper and you should aim to spend about 80 mins on it.

## SCHOOL OF PHYSICS & ASTRONOMY

### PHYSICAL CONSTANTS

Candidates are advised that they should only use the number of significant figures appropriate for the problem they are attempting to solve.

#### GENERAL CONSTANTS:

Charge on electron	$-e = -1.60217733 \times 10^{-19} \text{ C}$
Mass of electron	$m_e = 9.1093897 \times 10^{-31} \text{ kg} (\equiv 0.510998902 \text{ MeV}/c^2)$
Mass of proton	$m_p = 1.6726231 \times 10^{-27} \text{ kg} (\equiv 938.27200 \text{ MeV}/c^2)$
Mass of neutron	$m_n = 1.6749286 \times 10^{-27} \text{ kg} (\equiv 939.56533 \text{ MeV}/c^2)$
Permeability of vacuum	$\mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1}$
Permittivity of vacuum	$\epsilon_0 = 8.854187817 \times 10^{-12} \text{ F m}^{-1}$
Fine structure constant	$\alpha = 1/137.035989$
Gravitation constant	$G = 6.67259 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$
Boltzmann's constant	$k_B = 1.3806503 \times 10^{-23} \text{ J K}^{-1}$
Atmospheric pressure	$1 \text{ atm.} = 1.01325 \times 10^5 \text{ N m}^{-2} (\text{Pa})$
Stefan-Boltzmann constant	$\sigma = 5.6704 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$
Avogadro's number	$N = 6.0221367 \times 10^{23}$
Velocity of light	$c = 2.99792458 \times 10^8 \text{ m s}^{-1}$
Bohr radius	$a_0 = 5.2917721 \times 10^{-11} \text{ m}$
Bohr magneton	$\mu_B = 9.274006 \times 10^{-24} \text{ J T}^{-1}$
Planck's constant	$h = 6.62607544 \times 10^{-34} \text{ J s}$
Planck's constant/ $2\pi$	$\hbar = 1.05457266 \times 10^{-34} \text{ J s}$

#### ASTRONOMICAL CONSTANTS

Astronomical unit:	$1 \text{ AU} = 1.49597871 \times 10^{11} \text{ m}$
Parsec:	$1 \text{ pc} = 3.08567758 \times 10^{16} \text{ m}$
Mass of the Earth	$M_\oplus = 5.97 \times 10^{24} \text{ kg}$
Radius of the Earth	$R_\oplus = 6.37814 \times 10^6 \text{ m}$
Mass of the Sun	$M_\odot = 1.99 \times 10^{30} \text{ kg}$
Radius of the Sun	$R_\odot = 6.96 \times 10^8 \text{ m}$
Luminosity of the Sun	$L_\odot = 3.85 \times 10^{26} \text{ W}$
Thomson cross-section	$\sigma_T = 6.652459 \times 10^{-29} \text{ m}^2$

#### ATOMIC AND NUCLEAR PHYSICS UNITS

	$1 \text{ fm} = 10^{-15} \text{ m}$
	$1 \text{ barn} = 10^{-28} \text{ m}^2$
Atomic mass unit	$1 \text{ u.} = 1.6605402 \times 10^{-27} \text{ kg}$
Atomic energy unit	$1 \text{ a.u.} = 27.2113834 \text{ eV}$
Ångstrom	$1 \text{ Å} = 10^{-10} \text{ m}$
Electron volt	$1 \text{ eV} = 1.6021765 \times 10^{-19} \text{ J}$
	$\hbar c = 197.32696 \text{ MeV fm}$

(Updated 4 June 2010)

There is not a formula sheet although you may be given some formulae in the questions (making the questions simpler).

Writing things down (blind) is supposed to be the best revision technique.

Do think about setting questions as you go...

Do use the problem sheets as example questions - solutions are all online (Blackboard)

Past exam papers are linked from the blackboard site...

The summaries below are intended to get you started with the key elements of the course... but all material taught in lectures can be examined.



# KINETIC THEORY OF GASES

## PRINCIPLES

- ① A gas is composed of a large number of molecules
- ② Molecules are small relative to separation
- ③ Molecules uniformly distributed & move randomly
- ④ Obey Newton's laws of motion
- ⑤ No forces except in collisions with each other or the walls (hard spheres)
- ⑥ Collisions with walls are elastic

$$n = N/V$$

$$p = \frac{Nm}{V}$$

$$P = \frac{1}{3} n m \bar{v}^2$$

$$U = \frac{1}{2} m \bar{v}^2 \times \text{number of molecules}$$
$$= \frac{1}{2} m \bar{v}^2 (nV)$$

$$PV = \frac{2}{3} U$$

$$PV = NkT$$

$$PV = nRT$$

$$R = k N_A$$

Equipartition Theorem: for a classical system in thermal equilibrium, the total energy is shared equally among all the degrees of freedom

& we can see each gets  $\frac{1}{2} kT$

$$U = \frac{3}{2} NkT$$

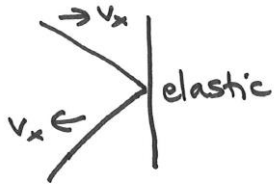


## Pressure

$$P = \frac{\text{Force}}{\text{area}}$$

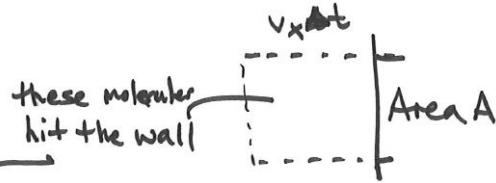
$$Pa = Nm^{-2}$$

Caused by collisions



$$F = \frac{\Delta m v}{\Delta t} \quad \text{2Mv}_x \text{ per collis}$$

In time  $\Delta t$



except  $\frac{1}{2}$  that were already travelling away

$$\# \text{ collisions in } \Delta t = \frac{1}{2} n \times (v_x \Delta t A) \quad \leftarrow \text{volume}$$

$$\text{Force} = \frac{2Mv_x}{\Delta t} \times \frac{1}{2} n (v_x \Delta t A)$$

$$P = \frac{F}{A} = n M v_x^2$$

I've assumed all molecules have the same  $v_x$ ... so  $v_x^2 \rightarrow \overline{v_x^2}$  the mean squared velocity

Since motion is random

$$\overline{v_x^2} = \overline{v_y^2} = \overline{v_z^2} = \frac{1}{3} \overline{v^2}$$

We arrive at

$$P = \frac{1}{3} n m \overline{v^2}$$

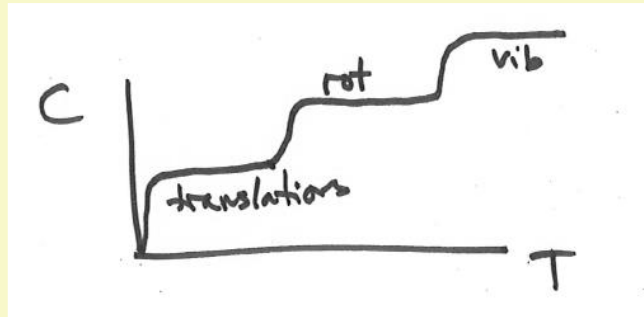
# HEAT CAPACITIES

$$\Delta Q = C \Delta T$$

N unlinked molecules have  $3N$  motions

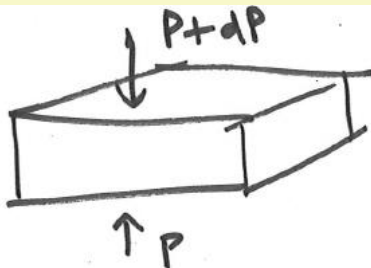
- > 3 translations (KE)
- > 2 or 3 rotations (KE)
- > rest vibrations (KE + PE)

$$U = N_A \frac{1}{2} kT \times \text{dof}$$



Quantum thresholds

# Hydrostatic Equilibrium



$$\frac{dP}{dh} = -\rho g$$

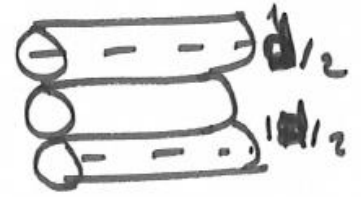
$$P(E) \propto e^{-E/kT}$$

BOLTZMAN  
FACTOR

# TRANSPORT

## Mean Free Path

$$\text{collisions/sec} = \underbrace{\pi d^2 v}_{\text{volume}} n \sim N/V$$



$$\vec{v}_{\text{rel}} \cdot \vec{v}_{\text{rel}} = v_A^2 + v_B^2 - 2\vec{v}_A \cdot \vec{v}_B$$

$$v \rightarrow \langle v_{\text{rel}} \rangle \approx \sqrt{2} v$$

$$\lambda = \frac{1}{\sqrt{2} \pi d^2 n}$$

$$\frac{\text{speed}}{\text{collision rate}}$$

$$P(\text{collision in } \delta r \text{ travel}) = \frac{\delta r}{\lambda}$$

$$\begin{aligned} P(\text{travel } r + \delta r) &= P(\text{travel } r) P(\text{travel } \delta r) \\ &= P(\text{travel } r) (1 - P_{\text{collision}}(\delta r)) \\ &= P(\text{travel } r) (1 - \delta r / \lambda) \end{aligned}$$

$$\frac{dP}{dr} = \frac{P(r + \delta r) - P(r)}{\delta r} = -\frac{P(r)}{\lambda}$$

$$P = e^{-r/\lambda}$$

# Distance Travelled On A Random Walk

$$r^2 = \vec{r} \cdot \vec{r} = s_1^2 + s_2^2 + \dots + \vec{s}_1 \cdot \vec{s}_2 + \vec{s}_2 \cdot \vec{s}_1 + \dots$$

$$r = \sqrt{N} \lambda$$

## Diffusion

### Fick's Law

number  
flowing  
across  
dotted line  
/sec

$$\frac{dN}{dt}$$

$$= -DA \frac{dn}{dx}$$

↓ this defines Diffusion Coefficient  $m^2 s^{-1}$   
 ↑ Area flowing over  
 ↑ spatial density gradient

Net flow  
L → R

$$= \frac{1}{6} \left[ n(x-\lambda) - n(x+\lambda) \right] A \bar{v}$$

$n(x) - \frac{dn}{dx} \lambda$        $n(x) + \frac{dn}{dx} \lambda$

$$= \frac{1}{3} \lambda \bar{v} A \frac{dn}{dx}$$

$$\Rightarrow D = \frac{1}{3} \lambda \bar{v}$$

Fournier's Law:

$$\frac{dQ}{dt} = -k A \frac{dT}{dx}$$

$\uparrow$  Area  
 $\uparrow$  thermal conductivity  
 $\leftarrow$  temperature gradient

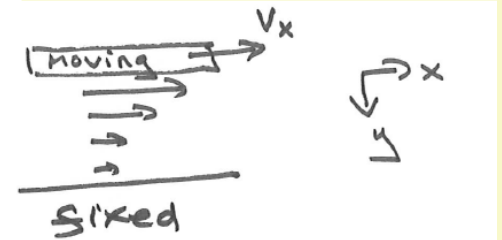
$\uparrow$   
 energy flow/s

Diffusion  $D = \frac{1}{3} \lambda \bar{v}$   
 Thermal cond.  $K = \frac{1}{3} \lambda \bar{v} C$   
 Viscosity  $\eta = \frac{1}{3} \lambda \bar{v} \rho$

Viscosity

$$F_x = -\eta A \frac{dv_x}{dy}$$

$\uparrow$  in opposite direction to  $v_x$   
 $\uparrow$  viscosity coefficient  $Nm^{-2}s^{-1}$   
 $\uparrow$  now area of moving surface



Temperature Dependence

D:  $\frac{1}{2} m \bar{v}^2 = \frac{3}{2} kT \rightarrow \bar{v} \sim T^{1/2}$

$\lambda = \frac{1}{\sqrt{2} \pi d^2 n} \sim \frac{1}{n} \sim \frac{\text{ideal } RT/P}{P} \sim T$

$D \sim T^{3/2}$

K:  $C$  is energy/volume & grows as  $n \sim \frac{P}{RT} \sim 1/T$

$K \sim T^{1/2}$

$\eta$ :  $\rho \sim n$  too  $\Rightarrow \eta \sim T^{1/2}$  They all grow with temperature

Pressure Dependence

We had  $\rho, C \sim n$   $\lambda \sim 1/n$   $n = \frac{P}{RT}$   
 $\Rightarrow \eta, K$  are predicted to be independent of  $P$ !

$\lambda$  can't be bigger than the container



**A4.** (a) Explain the two main effects that are assumed to be negligible in the definition of an ideal gas, and describe the regimes in which these effects become non-negligible.

(b) Even for warm and dilute gases, some approximations are valid only in specific cases. What are the necessary additional assumptions for the following statements to be good approximations?

(i) The internal energy of a gas with  $N$  molecules at temperature  $T$  is

$$U = \frac{3}{2}Nk_{\text{B}}T.$$

(ii) The specific heat capacity of a gas of  $\text{O}_2$  molecules is  $c_V = \frac{7}{2}R$ . [ 5 ]

**A1.** Solids which contain only one kind of atom typically follow the Dulong and Petit rule which predicts that an ideal monatomic crystal should have a molar heat capacity of  $3R$ . State the equipartition theorem and explain how it can be applied to an ideal monoatomic crystal. What causes the molar heat capacity to drop below  $3R$  at lower temperature? [ 5 ]

**A4.** Three identical flasks contain three different gases at the same pressure and temperature. Flask A contains  $\text{CO}_2$ , flask B contains  $\text{C}_2\text{H}_4$  and flask C contains He. Answer the following questions with a brief explanation.

[ 4 ]

- (a) Which flask contains the highest number of molecules?
- (b) Which flask contains the gas with highest density?
- (c) Which flask contains the fastest molecules?
- (d) Which flask has the gas with lowest heat capacity at high temperature?

(Note: the correct answers may include all or none of the flasks)



**B2.** Fourier's law for heat conduction says that the energy flow rate across area  $A$  driven by a temperature gradient  $dT/dx$  for a material of thermal conductivity  $\kappa$  is given by

$$\frac{dQ}{dt} = -\kappa A \frac{dT}{dx}.$$

A long rod has a uniform cross-sectional area of  $4.0 \text{ cm}^2$ . It is made of two sections joined end-to-end: a  $75.0 \text{ cm}$  long section of copper, and a  $25.0 \text{ cm}$  section of steel. The rod is insulated to prevent heat loss along its sides. The copper end is placed in perfect thermal contact with boiling water at  $100^\circ\text{C}$  and the steel end is placed in perfect thermal contact with an ice-water mixture at  $0^\circ\text{C}$ . The thermal conductivities of copper and steel are  $385.0 \text{ W m}^{-1} \text{ K}^{-1}$  and  $50.2 \text{ W m}^{-1} \text{ K}^{-1}$  respectively.

- (a) Draw a sketch graph to show how the temperature varies along the rod in the steady state. [ 4 ]
- (b) What is the temperature of the copper-steel junction in the steady state? [ 6 ]
- (c) What is the rate of heat transfer from the steam bath to the ice-water mixture in the steady state? [ 4 ]
- (d) What is the rate of entropy production? [ 4 ]
- (e) What would be the junction temperature, heat flow and entropy creation rate if the cross-sectional area of the rod were doubled? [ 2 ]