

PHYS1013

Energy and Matter

$U_i (n_i, P_i, V_i, \dots)$ $W = -nRT \int_{V_i}^{V_f} \frac{dV}{V} = -nRT \ln\left(\frac{V_f}{V_i}\right)$ $H = U + pV$ $T(K) = T(^{\circ}C) + 273.15$

$U_f (n_f, P_f, V_f, \dots)$ $C_p = (\Delta H / \Delta T)_p$ $\Delta U = Q - W$ $dH = dU + p dV + V dp$

$dU = dq + dw$ $dH = dq - p dV + V dp$ $C_p = \left(\frac{\partial H}{\partial T}\right)_p$ $W = P \Delta U$ $\Delta S = nRT \ln\left(\frac{V_f}{V_i}\right)$

$H = U + PV$ $dH = C_p dT$ $\Delta H = q_p = C_p \Delta T$ $C_v = (\Delta U / \Delta T)_v$ $W = \int_{V_i}^{V_f} P dV$

$dw = -p dV$ $\Delta S = \frac{\Delta_{\text{rev}} H}{T}$ $dS \geq \frac{dq}{T}$ $dS = \frac{dq_{\text{rev}}}{T}$

$C_v = \left(\frac{\partial U}{\partial T}\right)_v$ $\Delta S = \int_1^f \frac{dq_{\text{rev}}}{T}$

$\Delta U = m(u_2 - u_1) \Delta KE$
 $= \frac{1}{2} m (v_2^2 - v_1^2) \Delta PE$
 $= mg(z_2 - z_1)$

$W_b = \frac{P \Delta V}{1 - \eta}$ $\eta_{\text{th}} = \frac{W_{\text{net}}}{Q_{\text{in}}} = 1 - \frac{Q_{\text{out}}}{Q_{\text{in}}}$ $Q = \Delta U + P \Delta V$

$dH = (dq)_p$ $\Delta H = q_p$ $T_R = \frac{T}{T_{\text{cr}}}$ $dU = C_v dT$

$dU = (dq)_v$ $\Delta U = q_v$ $\Delta U = q_v = C_v \Delta T$

$P_{\text{cr}} = \frac{P}{P_{\text{cr}}}$ $W_b = P_i V_i \ln \frac{V_f}{V_i}$
 $= P_i V_i \ln \frac{P_i}{P_f} = RT_i \ln \frac{P_i}{P_f}$ $\Delta U = U_f - U_i = q(\text{heat}) + w(\text{work})$

$x = \frac{mg}{m_f + mg}$ $\eta_{\text{cr}} = \frac{P_{\text{cr}}}{RT_{\text{cr}}}$

Week 3 Module Survey: on
Blackboard:

We hugely appreciate
comments on the course at
this stage when we can
make changes

[https://forms.office.com/e/
ALn7gQ0Pya](https://forms.office.com/e/ALn7gQ0Pya)

by 17:00 on Friday
16/02/2024

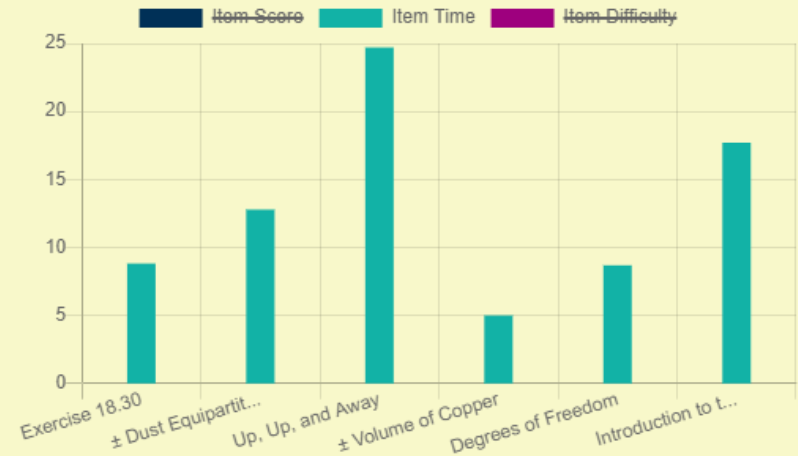
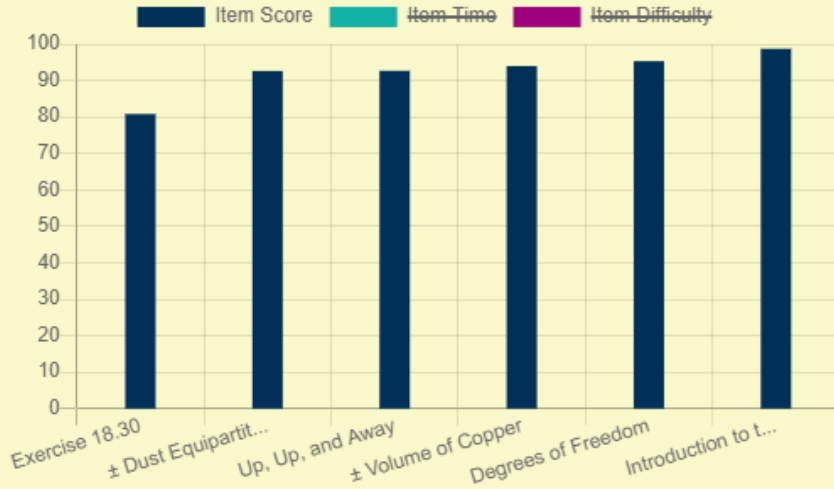
PHYS1013 Week 3 Survey



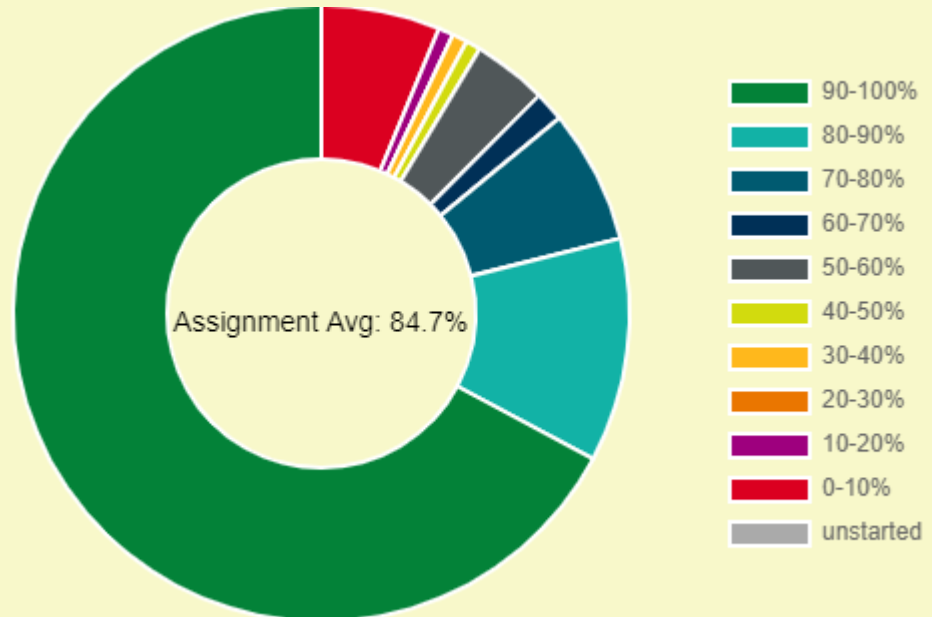
**Recommend some good music
- thank you to those avoiding
Rick Astley**

MP Week 1 - average score 90%

Average time 78 min



MP2 is due in this Sunday...

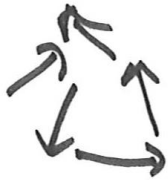


Distance Travelled On A Random Walk

18

The particle is constantly scattered randomly

$$\vec{r} = \vec{s}_1 + \vec{s}_2 + \dots$$



$$r^2 = \vec{r} \cdot \vec{r} = s_1^2 + s_2^2 + \dots + \vec{s}_1 \cdot \vec{s}_2 + \vec{s}_2 \cdot \vec{s}_1 + \dots$$

As we saw before in a random environment

$$\langle \cos \theta_{12} \rangle = 0 \quad \vec{s}_1 \cdot \vec{s}_2 = 0$$

$$\Rightarrow r^2 = \sum_{i=1}^N s_i^2 \approx N \lambda^2$$

↑ assume go λ each step.

$$\boxed{r = \sqrt{N} \lambda}$$

Note that since $N \propto t$ (more collisions in more time)

$$r \propto t^{1/2}$$

If one were to simulate a finite step walk particle by particle then $\langle \cos \theta_{12} \rangle \neq 0$ precisely... one gets a distribution of distances travelled



As $N \rightarrow \infty$ the Gaussian like peak narrows....

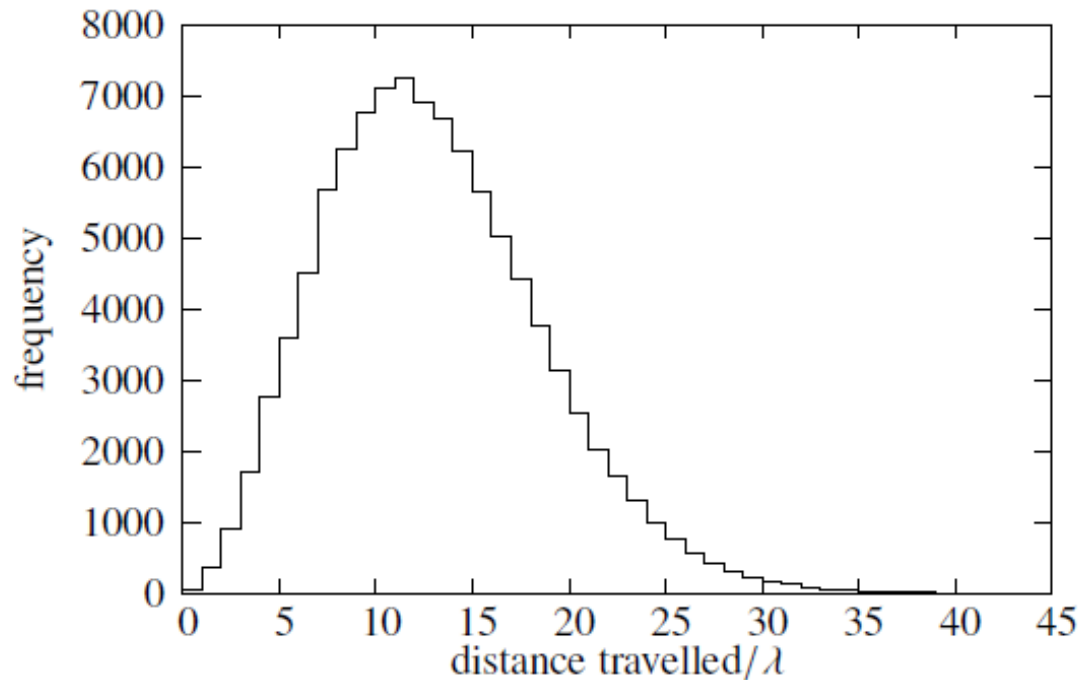
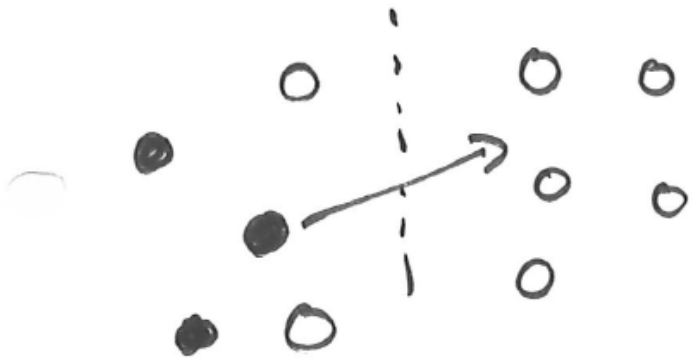


Figure 3.4 Frequency distribution of the distance travelled in 100 000 random walks, each of 100 steps, with the step-lengths themselves distributed according to an exponential distribution with mean λ . The mean distance travelled is 12.99λ , while the rms distance is 14.15λ and the maximum is 44.30λ .

Modelling Transport



black & white mix
over time

movement as:

identity \rightarrow diffusion

energy \rightarrow thermal
conductivity

momentum \rightarrow viscosity

We can do simple models as all these... assume
particles move in $\pm x, \pm y, \pm z$ only.

$\frac{1}{6}$ th do each motion.

Ignore collisions - work in small Δt where pts cross line unimpeded

Diffusion

Fick's Law

number
flowing
across
dotted line
/sec

$$\frac{dN}{dt}$$

$$= -DA$$

$$\frac{dn}{dx}$$

this defines Diffusion Coefficient $m^2 s^{-1}$

↑
Area
flowing
over

↑
spatial density
gradient

Let's use kinetic theory to study the same process.

$$\begin{array}{l} \text{Number glowing} \\ L \rightarrow R \end{array} = \frac{1}{6} \underbrace{n(x-\lambda)}_{\substack{\text{density} \\ \text{to left} \\ \text{by } \lambda}} \underbrace{A \bar{v} \Delta t}_{\substack{\text{volume that} \\ \text{can cross in} \\ \Delta t \text{ sec.}}}$$

\uparrow
 1 in 6 go
 right way

$$\begin{array}{l} \text{Number glowing} \\ R \rightarrow L \end{array} = \frac{1}{6} n(x+\lambda) A \bar{v} \Delta t$$

so there's a net glow if the density is bigger on one side.

$$\begin{array}{l} \text{Net glow} \\ L \rightarrow R \end{array} \frac{dN}{\Delta t} = \frac{1}{6} \left[\underbrace{n(x-\lambda)}_{n(x) - \frac{dn}{dx} \lambda} - \underbrace{n(x+\lambda)}_{n(x) + \frac{dn}{dx} \lambda} \right] A \bar{v}$$

$$= \frac{1}{3} \lambda \bar{v} A \frac{dn}{dx}$$

$$\Rightarrow \boxed{D = \frac{1}{3} \lambda \bar{v}}$$

EG Evaporation from a test tube



$$\text{Flow rate} = -\frac{1}{3} \lambda \bar{v} A \frac{dn}{dx}$$

? dN/dt

Assume saturation density n_s just above surface; and zero at top

$$\frac{dn}{dx} = \frac{n_s}{h}$$

Ideal gas: $P_s = n_s k T_I = \frac{P_s}{k T_I}$

$$|\text{Flow rate}| = \frac{1}{3} \frac{\lambda \bar{v} A P_s}{h k T_I}$$

$$\text{Mass loss} = \frac{M_{\text{mole}}}{N_A} \frac{\lambda \bar{v} A P_s}{3 h k T_I}$$

$$\uparrow M_{\text{mole}} = M_{\text{atom}} N_A$$

$$= \frac{M_{\text{mole}}}{R} \frac{\lambda \bar{v} A P_s}{3 h T_I}$$

$$\div \text{density} \quad \text{Volume loss} = \frac{M_{\text{mole}}}{R} \frac{\lambda \bar{v} A P_s}{3 h T_I \rho}$$

$$\div \text{area} \quad \text{Height loss} = \frac{M_{\text{mole}}}{R} \frac{\lambda \bar{v} P_s}{3 h T_I \rho}$$

$$T = 293 \text{ K} ; P_s = 1710 \text{ N m}^{-2} ; \rho = 1000 \text{ kg m}^{-3}$$

$$I_S \quad h = 0.1 \text{ m} , \lambda = 10^{-7} \text{ m} , \bar{v} = 650 \text{ m s}^{-1} \quad (\text{here weights less than } N_2)$$

$$\Rightarrow 0.24 \text{ mm/day}$$

20TH
FEBRUARY
19:00

UNIVERSITY OF
Southampton

Quantum Gravity

*Exploring String Theory
and the concept of
Brane Worlds*

PHYSOC Lecture series
Professor Nick Evans

B46 Physics and Astronomy Building
Room 3001



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